

AD609931

PHILCO.

A SUBSIDIARY OF *Ford Motor Company.*
AERONUTRONIC DIVISION

144-P

COPY	2	OF	3
HARD COPY	\$. 4.00		
MICROFICHE	\$. 1.00		



ARCHIVE COPY

DISCLAIMER NOTICE

THIS DOCUMENT IS THE BEST
QUALITY AVAILABLE.

COPY FURNISHED CONTAINED
A SIGNIFICANT NUMBER OF
PAGES WHICH DO NOT
REPRODUCE LEGIBLY.

BLANK PAGE

PHILCO

A SUBSIDIARY OF *Ford Motor Company*

AERONUTRONIC DIVISION

Publication No U-2955
W.O. 2056

SCIENTIFIC REPORT


ABSORPTION BY CO₂ BETWEEN 4500 AND 5400 cm⁻¹

Prepared for: Advanced Research Projects Agency
Washington 25, D. C.

Under Contract: NONr 3560(00)
ARPA Order No. 237/11-7-63

Prepared by: Darrell E. Burch
David A. Gryvnak
Richard R. Patty

Approved:


M. H. Johnson
Director, Physics Laboratory

15 December 1964

"This material is the result of tax-supported research and, as such may be freely reprinted with the customary crediting of the source."

ABSTRACT

The absorption by 38 different samples of CO_2 in the $4500 - 5400 \text{ cm}^{-1}$ (near 2 microns) region has been studied with the use of a spectrometer having a spectral resolution between approximately 0.3 and 1 cm^{-1} . Samples of CO_2 and $\text{CO}_2 + \text{N}_2$ were contained in two different absorption cells; one having a maximum path of 933 meters and used at a maximum pressure of 2.5 atm. The other was used at paths as long as 32.9 meters and at pressures as high as 15 atm. Photographs of the spectra of every sample as well as a table of transmittance versus wavenumber are included. Also presented is a table of the integrated absorptance

$$\int_{\nu}^{\nu} [1 - T(\nu)] d\nu$$

versus ν , where approximately 600 values are tabulated for the largest samples.

ACKNOWLEDGMENT

The authors would like to acknowledge the assistance of several people during the course of this investigation. Mr. Francis Gates performed much of the work in getting the instruments assembled and in reducing the data. Mr. Ronald Steen also helped considerably in digitizing the data. Dr. Byron Edwards provided much assistance in developing the method and the instruments used in digitizing the spectral records. Mr. Ricardo DeCasas wrote the computer program and assisted in all the work connected with the computer.

Dr. William S. Benedict of Johns Hopkins University made many worthwhile suggestions and is responsible for the identification of several new absorption bands. Several helpful discussions with Drs. Louise Gray and Lewis Kaplan of the Jet Propulsion Laboratory are also greatly appreciated.

FOREWORD

The work reported here is a portion of an extensive investigation of the absorption of infrared radiation by atmospheric gases. Reports similar to this one but dealing with other spectral regions will be published within the next few months. This report does not contain much in the way of analysis of the data, but very detailed results are included in a manner which should be convenient for other workers. We expect to use the data presented here for at least one more report which will deal with the shapes and strengths of many of the weaker bands and with the application of the results to problems dealing with paths through the atmospheres of the earth and other planets.

TABLE OF CONTENTS

<u>Section</u>		<u>Page</u>
1	INTRODUCTION AND SUMMARY.	1-1
2	EXPERIMENTAL METHODS.	2-1
	2.1 GAS SAMPLING	2-1
	2.2 RECORDING AND CALIBRATION OF SPECTRA . . .	2-4
3	DISCUSSION OF ABSORPTION BANDS.	3-1
	3.1 IDENTIFICATION OF ABSORPTION BANDS	3-1
	3.2 BAND STRENGTHS	3-3
4	SAMPLE PARAMETERS	4-1
5	TRANSMISSION SPECTRA.	5-1
6	TABLE OF TRANSMITTANCES	6-1
7	TABLE OF INTEGRATED ABSORPTANCE	7-1
8	REFERENCES.	8-1
Appendix A	30-METER MULTIPLE-PASS CELL AND SOURCE TANK . .	A-1
Appendix B	ONE-METER MULTIPLE-PASS ABSORPTION CELL	B-1
Appendix C	DATA DIGITIZING SYSTEM.	C-1

LIST OF TABLES

<u>Table No.</u>	<u>Title</u>	<u>Page</u>
2-1	RESOLUTION SCHEDULES.	2-4
3-1	CO ₂ BANDS BETWEEN 4500 AND 3400 cm ⁻¹	3-2
3-2	BAND STRENGTHS.	3-5
4-1	SAMPLE PARAMETERS	4-2, 4-3
5-1	CALIBRATION TABLE	5-22

LIST OF FIGURES

<u>Fig. No.</u>		<u>Page</u>
3-1	SPECTRUM OF A VERY LARGE SAMPLE OF CO ₂ BETWEEN 4540 AND 5380 cm ⁻¹	3-4
5-1 to 5-20	TRANSMISSION SPECTRA.	5-2 to 5-22
A-1	OPTICAL DIAGRAM OF THE 30-METER MULTIPLE-PASS ABSORPTION CELL AND SOURCE TANK	A-2
C-1	DIRECT DIGITIZING SYSTEM	C-2
C-2	REPLOTTING AND DIGITIZING SYSTEM	C-5

SECTION 1

INTRODUCTION AND SUMMARY

The major portion of the absorption by CO_2 in the 4500-5400 cm^{-1} region (near 2 microns) is due to three combination bands: $2\nu_1 + \nu_3$ whose center is at 5099.62 cm^{-1} , $\nu_1 + 2\nu_2 + \nu_3$ at 4977.79 cm^{-1} and $4\nu_2 + \nu_3$ at 4853.58 cm^{-1} . Besides these three bands of medium strength, there are approximately 40 weaker bands which have been observed in this region, and at certain wavenumbers one or more of them may produce most of the absorption.

Howard, Burch and Williams¹ have made some quantitative measurements on the absorption in this region, but the spectral resolution of their instrument was low. Courtoy² has measured the positions of several hundred absorption lines with very good accuracy and has identified many bands which had not been observed previously; but only limited information about the amount of absorption one might expect from a given sample can be obtained from his results. Weber, Holm and Penner³ have measured the strengths of the three main bands, but with an accuracy of only ± 20 percent. Breeze and Ferriso⁴ have also investigated the effect of changing the temperature on the band strengths. Kuiper⁵ has investigated the absorption by large samples; but his measurements were made with only moderate resolution and at pressures of several atmospheres.

Because of the limitations of the previous work, there was still need for data on samples covering wide ranges of pressure and absorber thickness and with resolution sufficiently good that at least some lines from most of the bands could be resolved. The present investigation was undertaken for the purpose of obtaining these data. The long absorption cell with possible paths as great as 933 meters has enabled us to study

samples having very large absorber thickness at pressures which are not so great as to smooth out the structure in the band. The short absorption cell has also made it possible to investigate samples at higher pressures in order that band strengths can be measured easily.

From the results presented in this report one should be able to determine the strengths of nearly every band of any significance in this region. The widths of most of the lines can also be determined from "curves of growth" of the average absorptance over relatively narrow intervals. Once the strengths and widths of the lines have been found, the absorptance at any wavenumber can be calculated for a great variety of samples, even though they are quite different from any of the samples studied. This is true, even for samples having non-uniform temperature and pressure.

Because of the long paths which are possible with the long absorption cell, we have been able to observe and identify five new CO₂ bands in this region which have not been reported previously. Approximate strengths of these as well as the three major bands have been determined and are tabulated in Section 3.

Spectra of 38 samples covering a wide range of pressures and path lengths have been reproduced and digitized. The reproduced spectra are shown in Section 5 and a table of transmittance $T(\nu)$ versus wavenumber ν is presented in Section 6. Section 7 contains a table of

$$\int_{\nu^1}^{\nu} [1 - T(\nu)] d\nu,$$

the integrated absorptance versus wavenumber; ν^1 is on the low wavenumber side of the region considered for a given sample, and tabulations are made for nearly 600 different values of ν for the largest samples.

Detailed descriptions of the two absorption cells used and of the digitizing apparatus are given in Appendices A, B, and C.

SECTION 2

EXPERIMENTAL METHODS

2.1 GAS SAMPLING

Samples of CO_2 and $\text{CO}_2 + \text{N}_2$ were contained in two different absorption cells which are described in detail in Appendices A and B. The longer cell has a base length of approximately 30 meters and was used at as many as 32 passes, giving a total path length of 933 meters. This cell is approximately 0.9 meters in diameter and can be evacuated to less than 1 micron of Hg or pressurized to as much as 2.5 atmospheres. The shorter cell has a base length of approximately 1 meter and was used at as many as 32 passes. It also can be evacuated and pressurized to as much as 15 atmospheres.

The CO_2 was taken from the vapor in a dewar which contained both liquid and vapor. It was maintained at a temperature less than about -20°C , so that the pressure never exceeded approximately 300 psig. There was found to be considerably less water in the sample if it were drawn from the vapor over the liquid rather than allowing the liquid to be drawn off and to evaporate into the cell. It appeared as though most of the water in the CO_2 container was trapped in the liquid CO_2 ; and if the liquid was drawn off the H_2O was allowed to enter the absorption cell. Of course, as the vapor was drawn off the pressure, and thus the temperature, in the dewar dropped considerably and the filling rate decreased. When the pressure dropped below about 50 psig, the valves were closed and the dewar was allowed to warm up. Approximately two days were required to fill the large absorption cell to 2.5 atmospheres from two dewars of CO_2 .

The cell could have been filled in much less time by drawing off the liquid CO_2 rather than the vapor, but the slower technique was chosen because of the drier samples which could be obtained. The amount of H_2O in the N_2 was much less than that in the CO_2 ; therefore, it was not necessary to take extreme care to reduce the amount of H_2O put into the sample with the N_2 which was drawn off as a liquid into a heat exchanger where it evaporated and entered the cell.

Before introducing a sample into the cell, the cell was pumped for several hours in order to get rid of most of the H_2O which was adsorbed on the walls. This can be very important, depending on the previous sample in the cell. Even after using this much care, the absorption by traces of H_2O in the sample was bothersome at some wavenumbers.

$\text{CO}_2 + \text{N}_2$ mixtures were mixed in the cell; CO_2 was added to the desired pressure, N_2 was then added and the mixing was accomplished by fans inside the absorption cells. No attempt was made to change the relative abundances of the different isotopes of C or O in the samples studied.

Sample pressures less than approximately 0.06 atm were measured with a U-tube oil manometer, and a U-tube H_2 manometer was used for pressures in the range $0.06 < P < 2$ atm. All pressures > 2 atm were measured with a bourdon-type gauge.

Since CO_2 varies significantly from a perfect gas at some of the higher pressures used in this investigation, it was necessary to account for the Van der Waals' forces giving rise to the deviation from the perfect gas. In calculating the absorber thickness u , the following equation was used:

$$u = W p L \frac{273}{296} \quad (\text{atm cm}) \text{ STP.} \quad (1)$$

- L is the path length in centimeters,
- p is the partial pressure of CO_2 in atmospheres,
- $273/296$ accounts for the difference in density between room temperature and standard temperature,
- W is a correction term which accounts for the Van der Waals' forces and is given adequately for the pressures used in this investigation by

$$W = 1.00 + 0.0047 p. \quad (2)$$

All measurements were made with the samples at room temperature, approximately 296°K.

In order to relate the pressure of a sample to the half-width of the absorption lines, it is necessary to account for the different broadening abilities of CO₂ and N₂ when dealing with mixtures of these two gases. Burch, Gryvnak and Williams⁶ have used an equivalent pressure P_e given by

$$P_e = 1.3 p + (P - p) \quad (3)$$

where P is the total pressure. It is noted that the equivalent pressure approaches the total pressure for a very dilute mixture of CO₂ in N₂.

Since the simple classical theory predicts that the half-width of a line is proportional to the density of molecules, equation (3) should probably be modified to account for Van der Waals' forces which cause the density to increase faster than the pressure. The adjustment can be made by modifying (3) in the following way:

$$P_e = 1.3 W p + (P - p) \quad (3')$$

The factor 1.3 is not valid in all portions of the spectrum, but it has been found to represent satisfactorily the different broadening abilities throughout most portions of the band not too distant from the absorption lines producing the absorption. Winters, Silverman and Benedict⁷ have found, for example, that the self-broadening factor should be considerably greater than 1.3 when dealing with regions such as that on the high wavenumber side of the 2350 cm⁻¹ CO₂ band, where all of the absorption is due to the wings of very strong lines whose centers are several cm⁻¹ away. This results from the fact that the shape of a self-broadened line several cm⁻¹ from the line center is quite different from that of a nitrogen-broadened line. Although (3') does not give an accurate representation of an equivalent pressure which is proportional to the line half-width for all regions in the spectrum, it is adequate in most cases and serves to simplify the analysis of the data. As more accurate methods of accounting for the self-broadening are found a better value for equivalent pressure for the samples in any spectral region can be obtained.

2.2 RECORDING AND CALIBRATION OF SPECTRA

The spectrometer is "home made" and was contained in a tank which could be evacuated to essentially eliminate absorption due to atmospheric gases outside of the absorption cell. The spectrometer tank, as well as another tank containing the radiation source and chopper, were connected to the absorption cell by means of flexible bellows so that all of the optical path external to the absorption cell could be evacuated.

The spectrometer is of the Ebert type with the main mirror having a 75 cm focal length. It utilized a small grating having a ruled area 64 x 64 mm with 600 lines/mm and blazed at 1.6 microns. The grating was used in the first order and a Ge filter eliminated overlapping orders of shorter wavelength. A PbS cell cooled with liquid nitrogen was used as the detector. It was not necessary to cool the detector below dry ice temperature for operation in this wavelength region, but the dewar was designed to hold the liquid nitrogen for use at longer wavelengths. Cooling by liquid nitrogen was, therefore, more convenient and was used since the signal-to-noise ratio was approximately the same at both temperatures. Three different resolution schedules were used in recording the spectra; the approximate spectral slitwidths for each schedule at three different wavenumbers are given in Table 2-1.

TABLE 2-1
RESOLUTION SCHEDULES

ν (cm^{-1})	Spectral Slitwidth (cm^{-1})		
	Schedule A	Schedule B	Schedule C
4600	0.30	0.39	0.54
5000	0.35	0.50	0.70
5400	0.41	0.62	0.86

Resolution Schedule A was used in several regions of the spectrum to resolve closely spaced lines; but because of the small signal available while using this schedule, considerable time was required to scan a portion of the spectrum. Therefore, Schedule B

which represents a compromise between resolution and scanning time was used for the majority of the data. Schedule C was used for some spectra of samples at several passes of the absorption cell for which the signal was low.

Background curves were obtained with the sample cell evacuated for each number of passes for which measurements were obtained. The background curve was different for different numbers of passes since the reflectivity of the mirrors in the multiple-pass optics varies with wavenumber. The appropriate background curve which represents 100 percent transmittance was then fitted to each spectrum and traced on it. All the sample spectra extended beyond the region of absorption on both ends of the band so that dependable "tie-points" between a spectrum and its background could be established. The transmittance was determined from the ratio of the deflection of the sample spectrum to that of the background curve at the same wavenumber. Each spectrum was digitized, with the information put on IBM cards by the technique which is described in considerable detail in Appendix C.

Most of the points used in the wavenumber calibration were obtained from the CO₂ lines of the bands being studied since the positions of many of them have been determined with very good accuracy by Courtoy². The lines of a few other bands which have not been measured or calculated by Courtoy can be determined from the values of the molecular constants provided by him. The wavenumbers of several lines in the 002 band of CO₁₆O₁₈ which was omitted by Courtoy have been measured by Díaz⁸; his values were used for calibration in the region of this band. H₂O lines whose positions are well known were also used as standards in certain parts of the spectrum.

Lines separated by an average of about 10 cm⁻¹ were used in calibrating each spectrum; it was assumed that each spectrum was linear between the calibration points. The absorption lines used as standards are summarized in Table 5-1, and each of the lines is indicated in the spectra shown in Section 5.

SECTION 3

DISCUSSION OF ABSORPTION BANDS

3.1 IDENTIFICATION OF ABSORPTION BANDS

All of the absorption bands of the various isotopes of CO_2 which absorb appreciably in this region are listed in Table 3-1. ν_0 is the wavenumber of the center of the band for the species and the transition indicated. Portions of all but five or six of these absorption bands have been observed with good resolution in the very extensive work of Courtoy². By using samples in which the abundances of some of the less common isotopes were enhanced, Courtoy was able to observe absorption lines which would otherwise be hidden by the lines of the more common isotopes. Several of the absorption bands of these rarer isotopes do not contribute significantly to the absorption by samples having the natural abundance of each isotope.

Most of Courtoy's measurements dealt with the determination of the positions of the absorption lines with very good accuracy, but only limited information about the strengths of the lines can be obtained from his results. Kuiper⁵ has measured the absorption by large samples for about 20 absorption bands in this region and correlated some of them with bands observed in the spectrum of Venus. Kuiper also observed some absorption at about 4505 cm^{-1} which he attributed to the $002-000$ transition of $\text{C}^{13}\text{O}^{16}\text{O}^{18}$. We were not able to observe this band for the largest sample used, which was at a pressure of 2.5 atm and a path length of 933 meters. In our experiment the noise is approximately three or four percent of the signal in this portion of the spectrum when operating at the maximum path length; therefore, it is possible that the absorbance near the middle of a band in this region could be as much as 0.02 or 0.03 without being detected by us. Unfortunately, only two or three spectra of very large samples were obtained by us in this region

TABLE 3-1
CO₂ BANDS BETWEEN 4500 and 5400 cm⁻¹

ν cm ⁻¹	Transition	Species	ν cm ⁻¹	Transition	Species
4591.0 ^a	15 ¹ ₀ ^b	C ¹² ₀₂ ¹⁶	4942.484	14 ⁰ ₁ ←02 ⁰ ₀	C ¹² ₀₂ ¹⁶
4639.53 ^c	00 ⁰ ₂	C ¹² ₀ ¹⁶ ₀ ¹⁸	4953.327	14 ² ₁ ←02 ² ₀	C ¹² ₀₂ ¹⁶
4673.635	06 ¹ ₁ ←02 ² ₀	C ¹³ ₀₂ ¹⁶	4959.635	22 ⁰ ₁ ←10 ⁰ ₀	C ¹² ₀₂ ¹⁶
4685.715	06 ⁰ ₁ ←02 ⁰ ₀	C ¹³ ₀₂ ¹⁶	4965.337	13 ¹ ₁ ←01 ¹ ₀	C ¹² ₀₂ ¹⁶
4687.75 ^a	06 ⁰ ₁ ←10 ⁰ ₀	C ¹² ₀₂ ¹⁶	4976.095	22 ⁰ ₁ ←02 ⁰ ₀	C ¹³ ₀₂ ¹⁶
4692.120	04 ⁰ ₁	C ¹³ ₀ ¹⁶ ₀ ¹⁸	4977.793	12 ⁰ ₁	C ¹² ₀₂ ¹⁶
4708.477	05 ¹ ₁ ←01 ¹ ₀	C ¹³ ₀₂ ¹⁶	4991.309	20 ⁰ ₁	C ¹³ ₀₂ ¹⁶
4748.012	04 ⁰ ₁	C ¹³ ₀₂ ¹⁶	4993.520	30 ⁰ ₁ ←10 ⁰ ₀	C ¹³ ₀₂ ¹⁶
4768.49	06 ² ₁ ←02 ² ₀	C ¹² ₀₂ ¹⁶	5013.730	21 ¹ ₁ ←01 ¹ ₀	C ¹³ ₀₂ ¹⁶
4790.520	06 ⁰ ₁ ←02 ⁰ ₀	C ¹² ₀₂ ¹⁶	5028.730	22 ² ₁ ←02 ² ₀	C ¹³ ₀₂ ¹⁶
4791.208	20 ⁰ ₁	C ¹² ₀ ¹⁶ ₀ ¹⁸	5042.538	04 ⁰ ₁	C ¹² ₀ ¹⁶ ₀ ¹⁸
4807.652	05 ¹ ₁ ←01 ¹ ₀	C ¹² ₀₂ ¹⁶	5062.394	22 ⁰ ₁ ←02 ⁰ ₀	C ¹² ₀₂ ¹⁶
4814.530	12 ⁰ ₁	C ¹³ ₀ ¹⁶ ₀ ¹⁸	5099.619	20 ⁰ ₁	C ¹² ₀₂ ¹⁶
4839.704	14 ⁰ ₁ ←10 ⁰ ₀	C ¹² ₀₂ ¹⁶	5123.170	21 ¹ ₁ ←01 ¹ ₀	C ¹² ₀₂ ¹⁶
4853.578	04 ⁰ ₁	C ¹² ₀₂ ¹⁶	5139.400	22 ² ₁ ←02 ² ₀	C ¹² ₀₂ ¹⁶
4853.745	14 ⁰ ₁ ←02 ⁰ ₀	C ¹³ ₀₂ ¹⁶	5168.60	01 ¹ ₂	C ¹³ ₀₂ ¹⁶
4858.100	14 ² ₁ ←02 ² ₀	C ¹³ ₀₂ ¹⁶	5217.63 ^a	30 ⁰ ₁ ←02 ⁰ ₀	C ¹² ₀₂ ¹⁶
4871.408	13 ¹ ₁ ←01 ¹ ₀	C ¹³ ₀₂ ¹⁶	5248 ^a	02 ⁰ ₂ ←01 ¹ ₀	C ¹² ₀₂ ¹⁶
4871.860	22 ⁰ ₁ ←10 ⁰ ₀	C ¹³ ₀₂ ¹⁶	5291.12	02 ² ₂ ←01 ¹ ₀	C ¹² ₀₂ ¹⁶
4887.348	12 ⁰ ₁	C ¹³ ₀₂ ¹⁶	5315.696	01 ¹ ₂	C ¹² ₀₂ ¹⁶
4904.820	12 ⁰ ₁	C ¹² ₀ ¹⁶ ₀ ¹⁸	5349 ^a	10 ⁰ ₂ ←01 ¹ ₀	C ¹² ₀₂ ¹⁶
4924.990	20 ⁰ ₁	C ¹³ ₀ ¹⁶ ₀ ¹⁸			

^a Observed in present study only; not reported previously.

^b Absence of quantum numbers for lower state indicates they are all zero.

^c Observed by Kuiper, positions of lines measured later by Diaz. Not observed by Courtoy.

of the spectrum, so that it is not possible to be sure that there is no absorption present. But in view of the fact that our largest sample was considerably bigger than the largest used by Kuiper, and since our spectral resolution is greater than his, it seems likely that the absorption observed by him in this region is due to some gas other than CO_2 which was present in his sample. Of course, it is conceivable that the abundance of O^{18} was greater in his sample than in ours; in which case he would be able to observe the band when we would not.

Other than for this apparent disagreement near 4505 cm^{-1} , there is good qualitative agreement among the results of the present study and those of Kuiper and Courtoy.

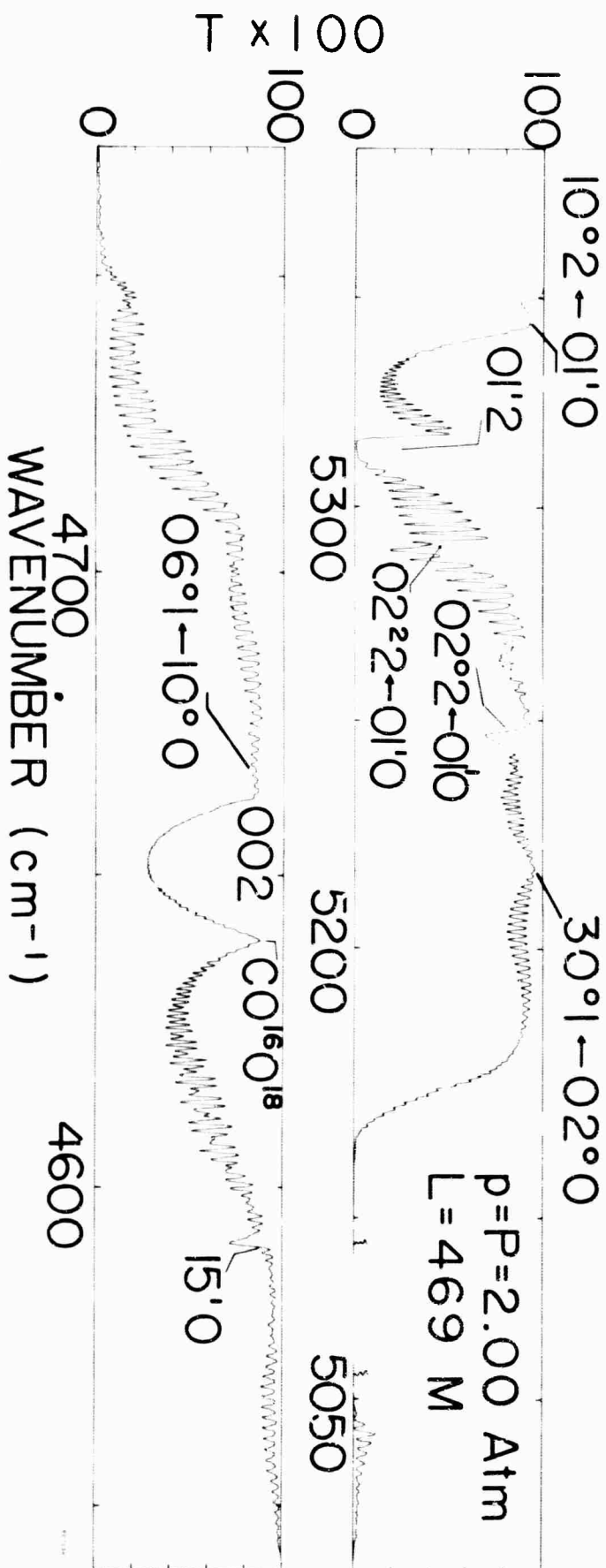
Because of the extremely large samples which were used in the present investigation, five new absorption bands which have not been reported previously were observed. These new bands are included in Table 3-1 and the identifications have been made from our spectra by Dr. William Benedict⁹. A spectrum of a very large sample which shows these new bands is given in Figure 3-1. The two panels of the figure represent the ends of the spectral region considered in this report; the sample is opaque in the regions not shown. The 002 band of $\text{CO}^{16}\text{O}^{18}$ has been observed by Kuiper⁵ and Diaz⁸, who have measured the positions of several of the lines in this band with very good accuracy. Courtoy has observed the 01^{12} and $02^{22}+01^{10}$ bands, but the other bands which are identified in the figure have not been reported previously.

3.2 BAND STRENGTHS

The strength, or intensity of an absorption band is given by $\int K(\nu) d\nu$, where $K(\nu)$ is the absorption coefficient defined by

$$T'(\nu) = \exp(-K(\nu)u).$$

$T'(\nu)$ is the true transmittance at ν which would be observed by a spectrometer with infinite resolution. The strength of a band is proportional to the transition probability and, by the use of theoretical considerations, one can calculate from it the strengths of individual lines within the band. The reverse is, of course, also true; the strength of a band, or at least the strength of one branch, can usually be calculated from the strength of one line.



RO7104 FIGURE 3-1. SPECTRUM OF A VERY LARGE SAMPLE OF CO₂ BETWEEN 4540 AND 5380 cm⁻¹

The sample is opaque in the spectral regions from approximately 4775 to 5020 cm⁻¹ and from 5050 to 5160 cm⁻¹. The broken portions of the curve represent regions where attempts have been made to correct for considerable absorption by H₂O. The 002 band of CC16018 has been observed by Kuiper and by Diaz; and Courtoy has observed the 01°2 and 02°2 ← 01°0 bands. The other bands which are identified have not been reported previously. Only a few of the lines near P20 of the 06°1 ← 10°0 band can be observed in the region indicated; the other lines of this band are overlapped by stronger lines.

The strengths of many of the bands in this spectral region could be determined from the data presented in this report. However, at the present time, the only bands whose strengths have been considered are the three major ones, $2\nu_1 + \nu_3$, $\nu_1 + 2\nu_2 + \nu_3$ and $4\nu_2 + \nu_3$ and the five new bands which have not been reported previously. The strengths of as many of the branches of these bands as it has been possible to determine are given in Table 3-2.

TABLE 3-2
BAND STRENGTHS

Transition ^a	$\nu_0 \text{ cm}^{-1}$	$\int K(\nu) d\nu \text{ (atm}^{-1} \text{ cm}^{-1} \text{ cm}^{-1}) \text{ STP}$				
		P	Q	R	Total Band ^b	Total Band Weber et al ^c
<u>Major Bands</u>						
04^0_1	4853.578	0.11	0.00	0.12	0.23	0.30
12^0_1	4977.793	0.52	0.00	0.53	1.05	1.11
20^0_1	5099.619	0.15	0.00	0.18	0.33	0.47
<u>New Bands^d</u>						
15^1_0	4591.0	0.16	0.04	0.22	0.42	
$06^0_1 \leftarrow 10^0_0$	4687.75	0.06				
$30^0_1 \leftarrow 02^2_0$	5217.63		0.00	0.27		
$02^0_2 \leftarrow 01^1_0$	5248	0.09	0.17	0.07	0.33	
$10^0_2 \leftarrow 01^1_0$	5349	0.03	0.08	0.03	0.14	

^a All bands listed in this Table pertain to the $\text{C}^{12}\text{O}_2^{16}$ molecule.

^b The uncertainty in the values for the major bands is about six or eight percent, but it is considerably higher for the new bands.

^c Values are taken from Weber, Holm and Penner³ and multiplied by 300/273 to adjust to density at STP. The authors given an uncertainty of $\pm 20\%$.

^d The strengths of the new bands were determined from our raw spectra by Benedict⁹.

The values given for the new bands should be considered tentative since they were obtained from the "raw" spectra. With some effort it is possible that more accurate values could be determined from the digitized output based on these spectra. Considerably better values could probably be obtained for the new bands from spectra with better resolution for samples containing less H₂O as an impurity.

The strengths of the major bands were determined from spectra of samples at a total pressure of approximately 10 atm. At pressures this high, the lines are wider than the spectral slit width so that the observed spectrum is not greatly different from what would be observed with infinite resolution; i.e., the observed transmittance $T(\nu)$ would be approximately the same as $T'(\nu) = \exp(-K(\nu)u)$. The band strengths were then determined from

$$\int K(\nu) d\nu = \frac{1}{u} \int -\ln T(\nu) d\nu \quad (4)$$

Strengths of other bands in this region, and possibly some revised values for those listed in Table 3-2, will appear in subsequent reports.

SECTION 4

SAMPLE PARAMETERS

The parameters for the 38 samples used in this investigation are summarized in Table 4-1. The path length L and the following pressures are expressed in both atmospheres and torr (mm Hg):

- p partial pressure of CO_2 .
- P total pressure due to CO_2 and N_2 .
- P_e the equivalent pressure which accounts for the difference between self-broadening and foreign broadening by nitrogen. It is given by
 $P_e = 1.3 W p + (P - p)$ where W is a Van der Waals correction given by equation (2).

Reproductions of the spectra of all the samples are shown in Section 5; the page numbers in which the different portions of each spectrum appear are listed in the last six columns of Table 4-1.

The resolution schedule corresponds to the spectral slitwidth of the spectrometer used to record the spectrum. Each schedule is related to slitwidth according to Table 2-1. Spectra of a few samples were obtained with two different resolution schedules; in these cases the same letter was used to identify the sample, but a star was used to denote the different resolution.

The maximum pressure used in the long absorption cell was 2.5 atm; spectra were obtained for the maximum pressure of pure CO_2 at 4, 8, 16 and 32 passes of the cell. The other samples were chosen so as to provide groups having approximately the same value of absorber thickness, which is given by

$$u = W p L \frac{273}{296} \quad (1)$$

TABLE 4-1
SAMPLE PARAMETERS

Sam. No.	p	P	P _e	p	P	P _e
	torr	torr	torr	atm	atm	atm
a	1900	1900	2500	2.50	2.50	3.29
b	1920	1920	2520	2.53	2.53	3.32
c	1920	1920	2520	2.53	2.53	3.32
d	1900	1900	2500	2.50	2.50	3.29
e	1522	1522	1995	2.00	2.00	2.62
f	760	760	992	1.000	1.000	1.31
g	760	760	992	1.000	1.000	1.31
g*	760	760	992	1.000	1.000	1.31
h	88	88	114	0.116	0.116	0.150
i	175	175	228	0.230	0.230	0.300
i*	175	175	228	0.230	0.230	0.300
j	175	707	759	0.230	0.930	0.999
k	29.4	29.4	38.2	0.0387	0.0387	0.0503
l	58.5	58.5	76.1	0.0770	0.0770	0.1001
l*	58.5	58.5	76.1	0.0770	0.0770	0.1001
m	58.5	212	230	0.0770	0.279	0.303
n	58.5	743	761	0.0770	0.978	1.001
o	8.95	8.95	11.8	0.0118	0.0118	0.0155
p	17.8	17.8	23.1	0.0234	0.0234	0.0304
q	17.8	72	77.3	0.0234	0.0948	0.102
r	17.8	223	228	0.0234	0.293	0.300
r	507	507	661	0.667	0.667	0.870
t	5.8	5.8	7.54	0.00763	0.00763	0.00992
u	5.8	21	22.7	0.00763	0.0276	0.0299
v	165	165	214	0.217	0.217	0.282
w	165	710	760	0.217	0.934	1.000
x	165	2280	2330	0.217	3.00	3.07
y	49.8	49.8	64.7	0.0655	0.0655	0.0851
z	49.8	213	228	0.0655	0.280	0.300
aa	49.8	776	791	0.0655	1.02	1.04
ab	49.8	2280	2295	0.0655	3.00	3.02
ac	49.8	7600	7615	0.0655	10.00	10.02
ad	49.8	213	228	0.0655	0.280	0.300
ae	49.8	2280	2295	0.0655	30.0	3.02
af	25	7600	7610	0.0329	10.00	10.01
ag	100	7600	7630	0.132	10.00	10.04
ah	3.13	25	25.9	0.00412	0.0329	0.0341
ai	3.13	225	226	0.00412	0.296	0.297
aj	3.13	760	761	0.00412	1.000	1.001
ak	3.13	2280	2280	0.00412	3.00	3.00
al	3.13	7600	7600	0.00412	10.00	10.00

TABLE 4-1 (cont.)

Sam. No.	L Path	u atm cm STP	Res. Sch.	Pages on which spectra appear					
				5375-5185	5185-5030	5030-4905	4905-4795	4795-4665	4665-4535
				cm ⁻¹	cm ⁻¹	cm ⁻¹	cm ⁻¹	cm ⁻¹	cm ⁻¹
a	933	217,000	B	5-2	5-4			5-16	5-20
b	469	111,000	B	5-2	5-4			5-16	5-20
c	237	55,700	B	5-2	5-4	5-8		5-16	5-20
d	121	28,100	B	5-3	5-5	5-9		5-17	5-21
e	469	87,100	B	5-3	5-5			5-17	5-21
f	933	86,200	B	5-3	5-5	5-9		5-17	5-21
g	237	21,900	B	5-2	5-4	5-8	5-12	5-16	5-20
g*	237	21,900	A	5-3				5-17	5-21
h	933	9,940	C	5-3	5-5	5-9	5-13	5-17	5-21
i	469	9,950	B	5-2	5-4	5-8	5-12	5-16	5-20
i*	469	9,950	A					5-18	
j	469	9,950	B	5-2	5-4	5-8	5-12	5-16	5-20
k	933	3,320	C	5-2	5-4	5-8	5-12	5-16	
l	469	3,320	B	5-2	5-4	5-8	5-12	5-16	5-20
l*	469	3,320	A		5-6	5-10	5-14	5-18	
m	469	3,320	B	5-3	5-5	5-9	5-13	5-17	5-21
n	469	3,320	B	5-3	5-5	5-9	5-13	5-17	5-21
o	933	1,010	C		5-4	5-8	5-12	5-16	
p	469	1,010	B		5-7	5-11	5-15	5-19	
q	469	1,010	B		5-7	5-11	5-15	5-19	
r	469	1,010	B		5-7	5-11	5-15	5-19	
s	16.5	1,010	B		5-6	5-10	5-14	5-18	
t	469	329	B		5-6	5-10	5-14	5-18	
u	469	329	B		5-4	5-8	5-12	5-16	
v	16.5	330	B		5-5	5-9	5-13	5-17	
w	16.5	330	B		5-5	5-9	5-13	5-17	
x	16.5	330	B		5-4	5-8	5-12	5-16	
y	16.5	99.6	B		5-7	5-11	5-15	5-19	
	16.5	99.6	B		5-6	5-10	5-14	5-18	
aa	16.5	99.6	B		5-4	5-8	5-12	5-16	
ab	16.5	99.6	B		5-6	5-10	5-14	5-18	
ac	16.5	99.6	B		5-5	5-9	5-13		
ad	4.16	25	B		5-6	5-10	5-14		
ae	4.16	25	B		5-5	5-9	5-13		
af	8.26	25	B		5-5	5-9	5-13		
ag	4.16	50.3	B		5-5	5-9	5-13		
ah	16.5	6.26	B			5-11			
ai	16.5	6.26	B			5-9			
aj	16.5	6.26	B			5-9			
ak	16.5	6.26	B			5-9			
al	16.5	6.26	B			5-8			

SECTION 5

TRANSMISSION SPECTRA

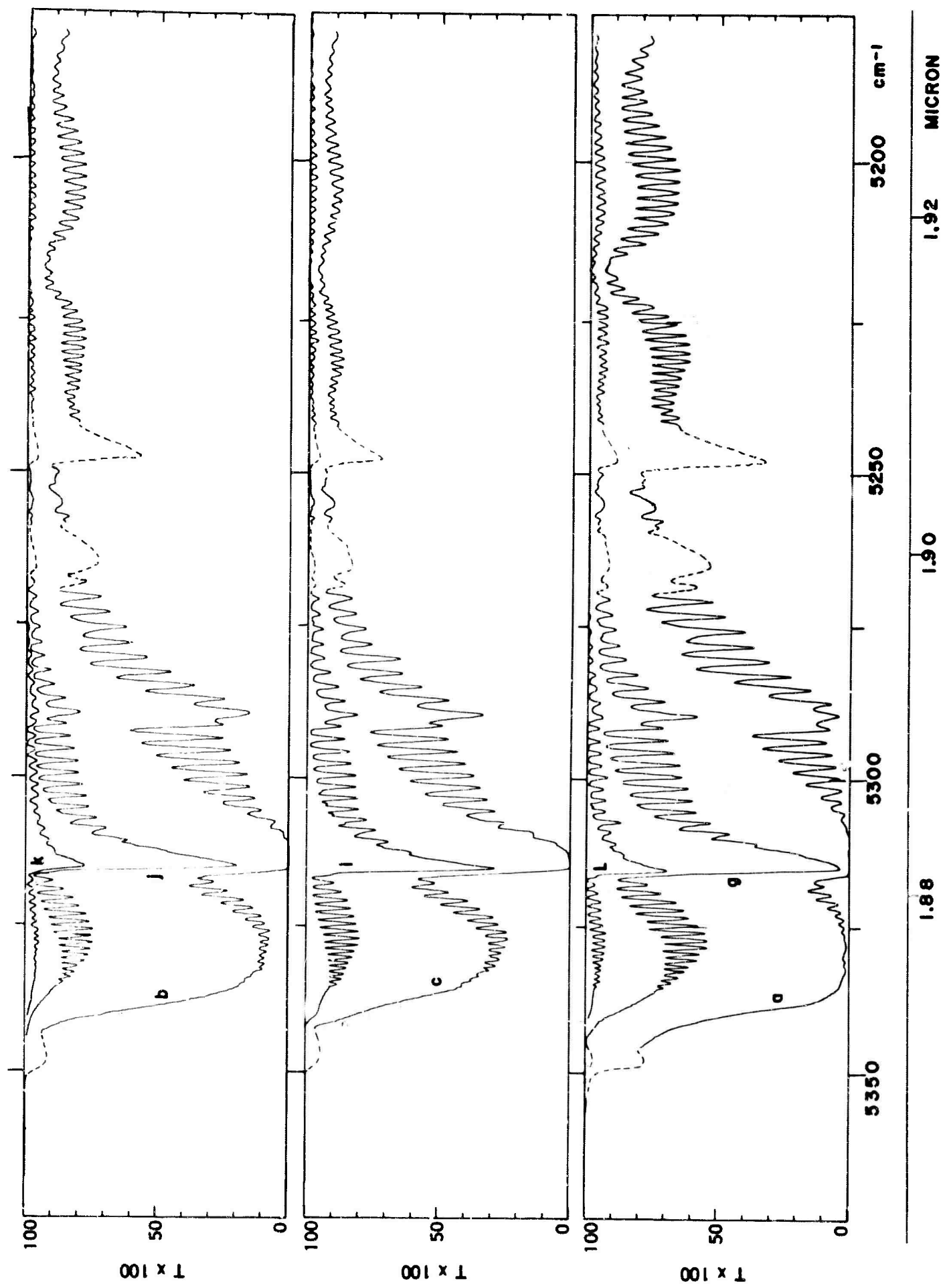
Reproductions of the spectra of 38 samples are shown in this section. The spectra were traced from the originals by use of the equipment described in Appendix C. The spectra are so long that they were divided into the following six regions:

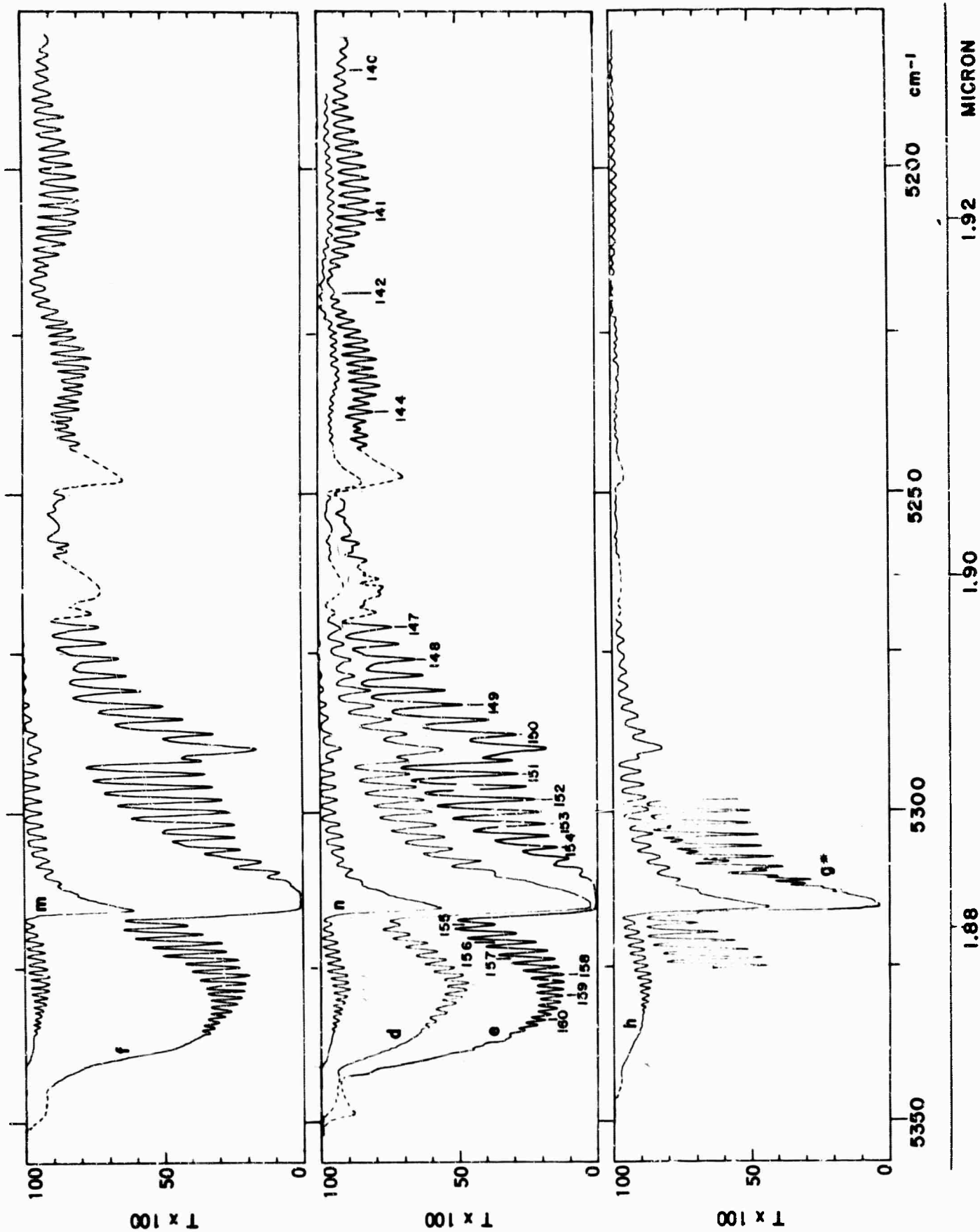
<u>Spectral Region</u>	<u>Pages</u>
5375 - 5185 cm^{-1}	5-2, 5-3
5185 - 5030	5-4 to 5-7
5030 - 4905	5-8 to 5-11
4905 - 4795	5-12 to 5-15
4795 - 4665	5-16 to 5-19
4665 - 4535	5-20, 5-21

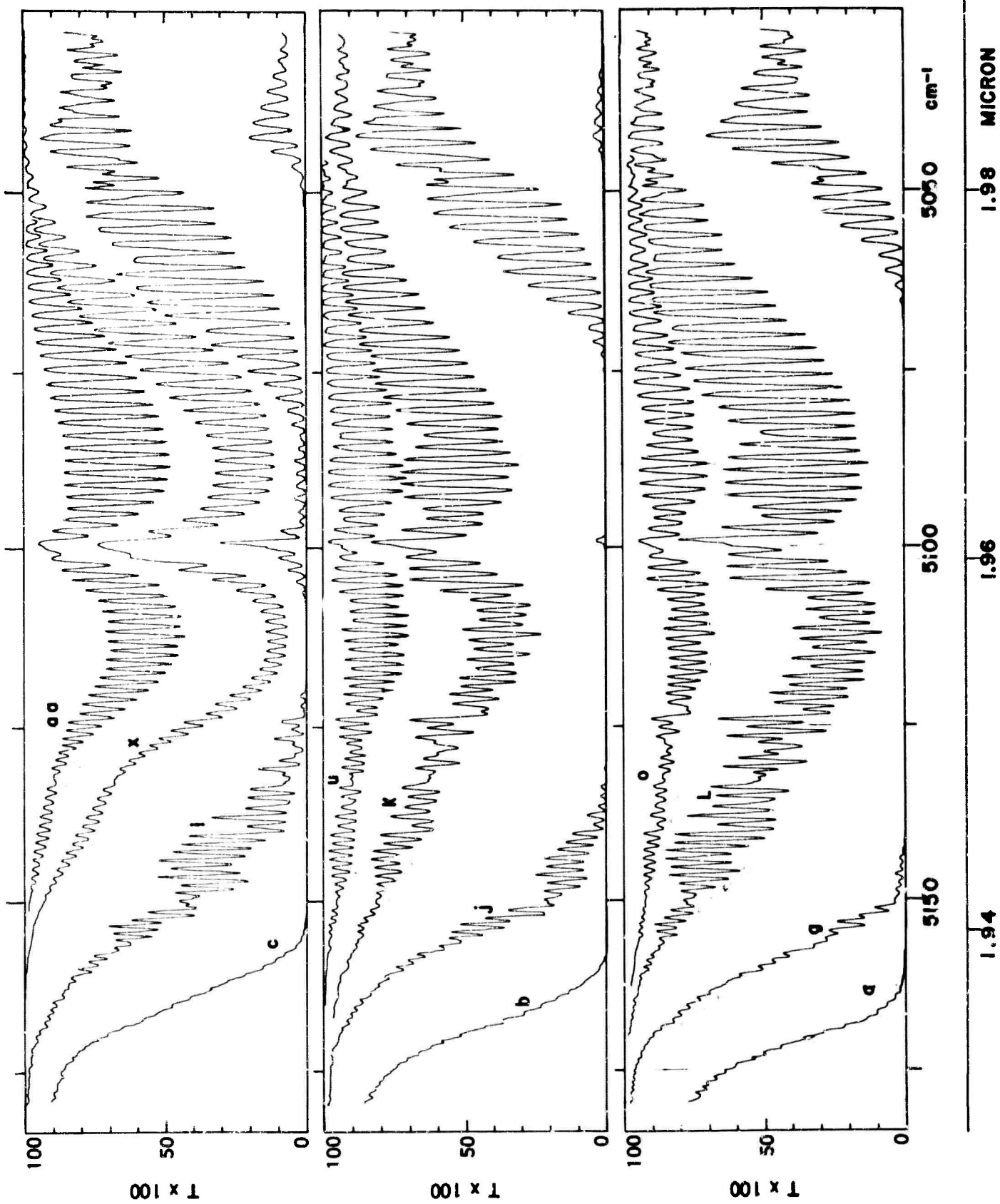
The wavenumber scale is accurate to approximately $\pm 1 \text{ cm}^{-1}$; more accurate values of line positions can be obtained from Courtoy or by interpolating between the designated calibration points. In some cases the original spectra of samples represented in the same panel in a figure were obtained at quite different times when the wavenumber calibration of the spectrometer was different, and the retraced spectra do not always "nest" properly. There is a slight periodic error in the screw which rotates the grating; therefore, the calibration shift was not linear in wavenumber. Calibration points were taken sufficiently close together that it can be assumed the spectra are linear between them; but the periodic error in the screw would in some cases cause significant error in calibration if one were to assume a spectrum was linear over an interval as wide as 20 or 30 cm^{-1} . The positions of known H_2O and CO_2 lines used for calibration are indicated by numbers in the figures, and the corresponding wavenumbers are listed in Table 5-1.

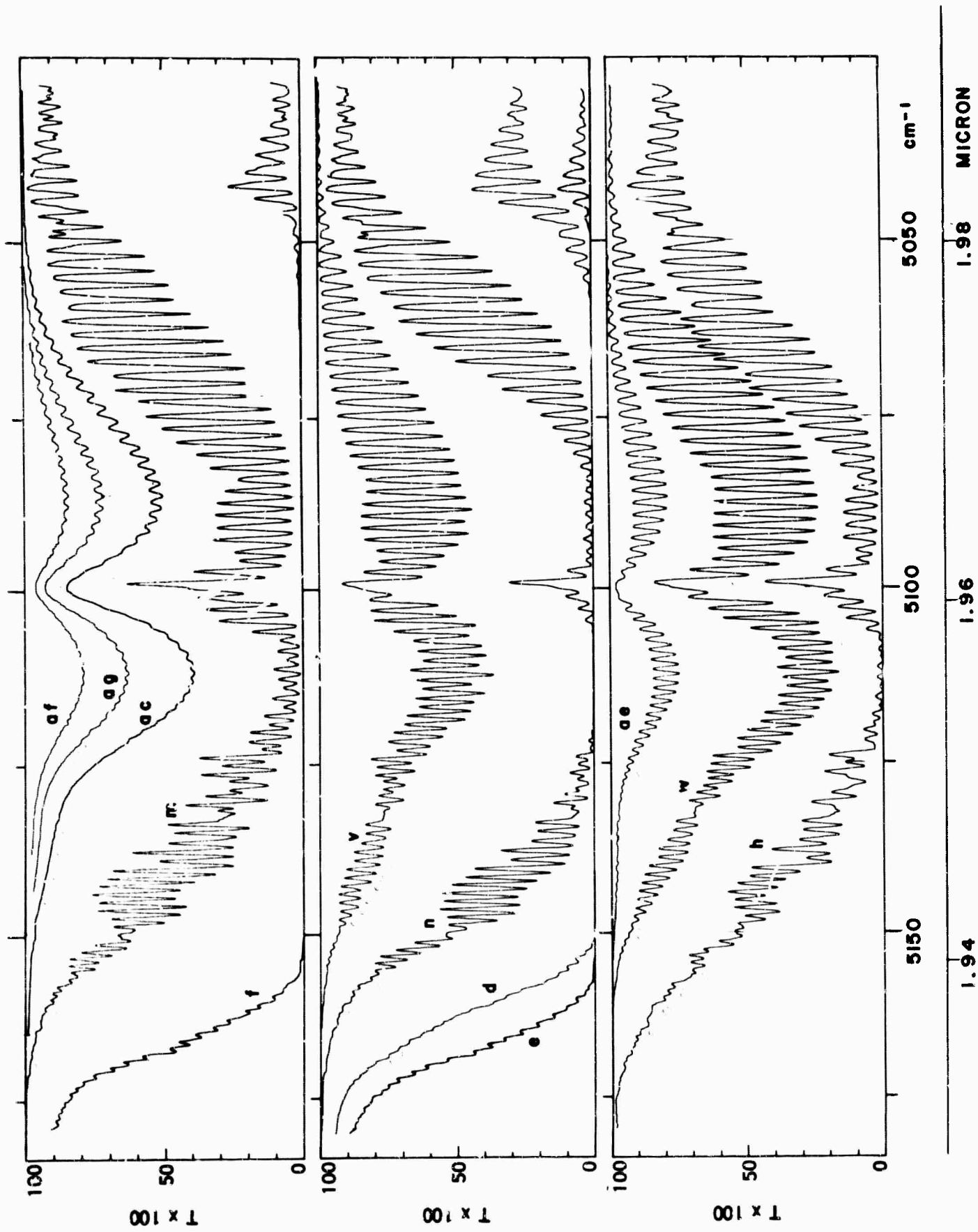
Each sample is designated by the same letter as that given in Table 4-1. The spectra designated by starred letters correspond to the same sample as the unstarred letter, but the spectrum was obtained with higher resolution in order to resolve more lines and to determine more accurate positions of the line centers. No calculations of transmittance or integrated absorptance were made from the spectra marked with stars.

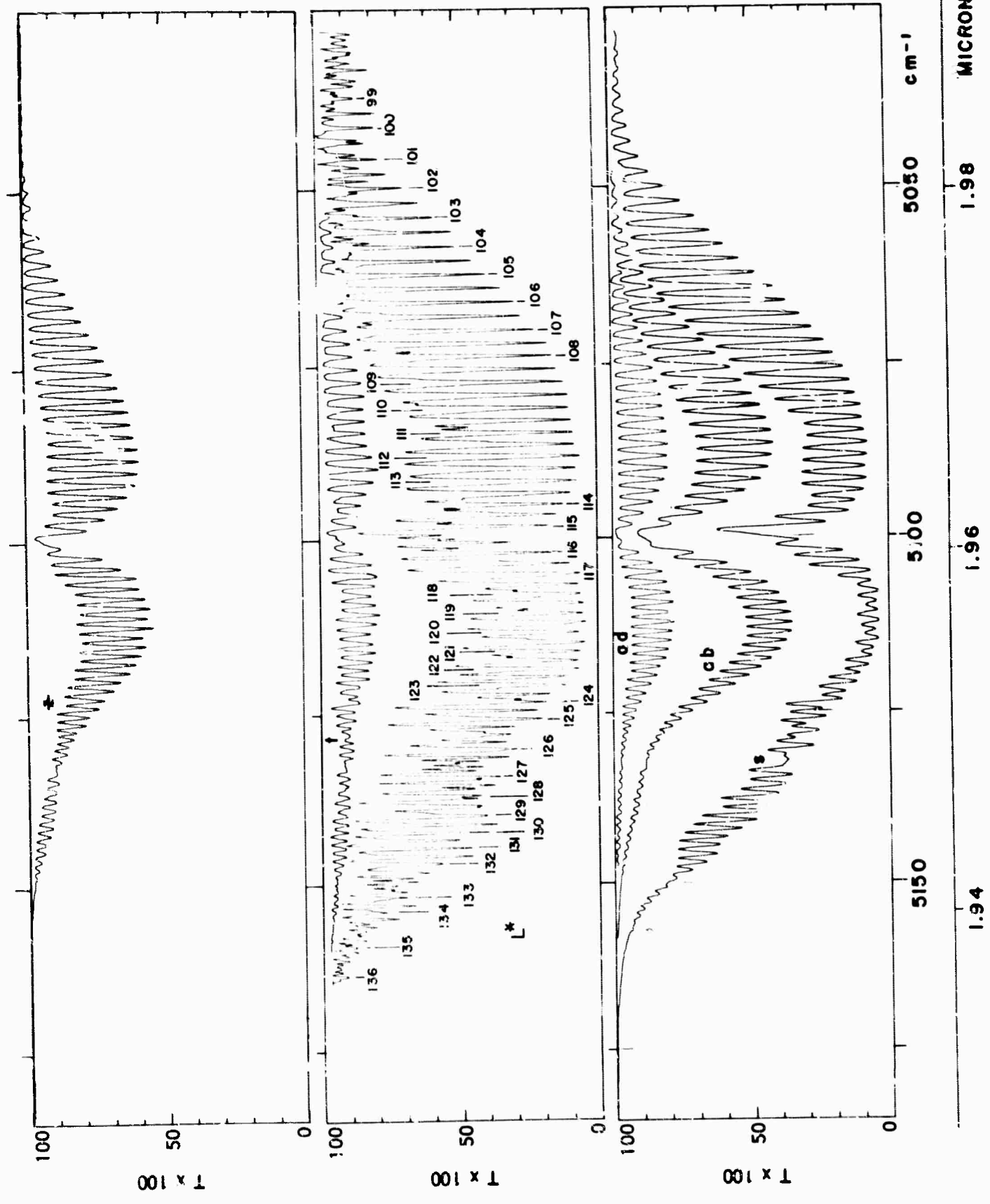
Regions where the CO_2 absorption was difficult to observe because of overlapping absorption by H_2O are indicated by dotted lines. Portions of some spectra were omitted where the absorption was small, particularly when spectra of larger samples having greater absorption in the same region are included.

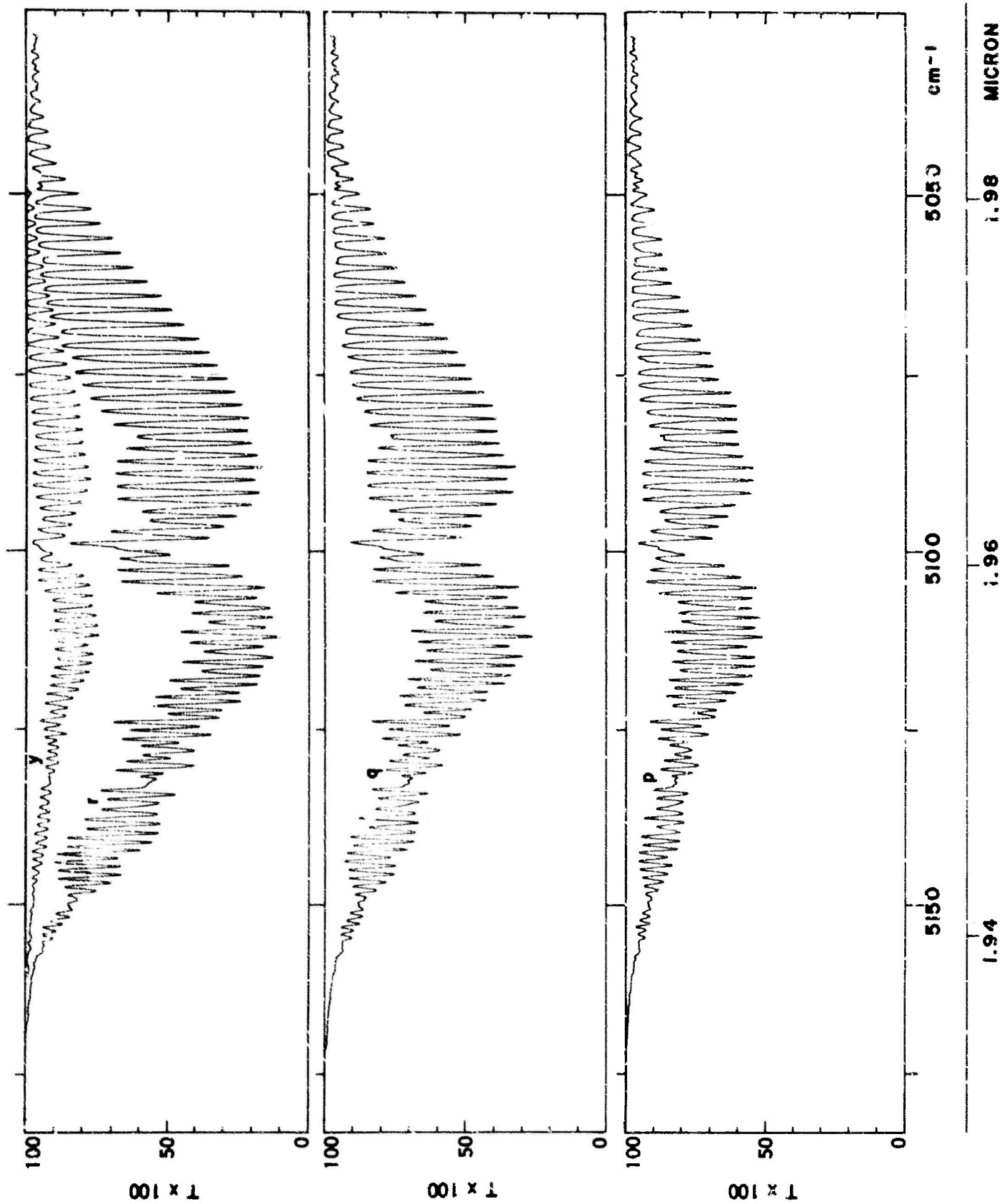


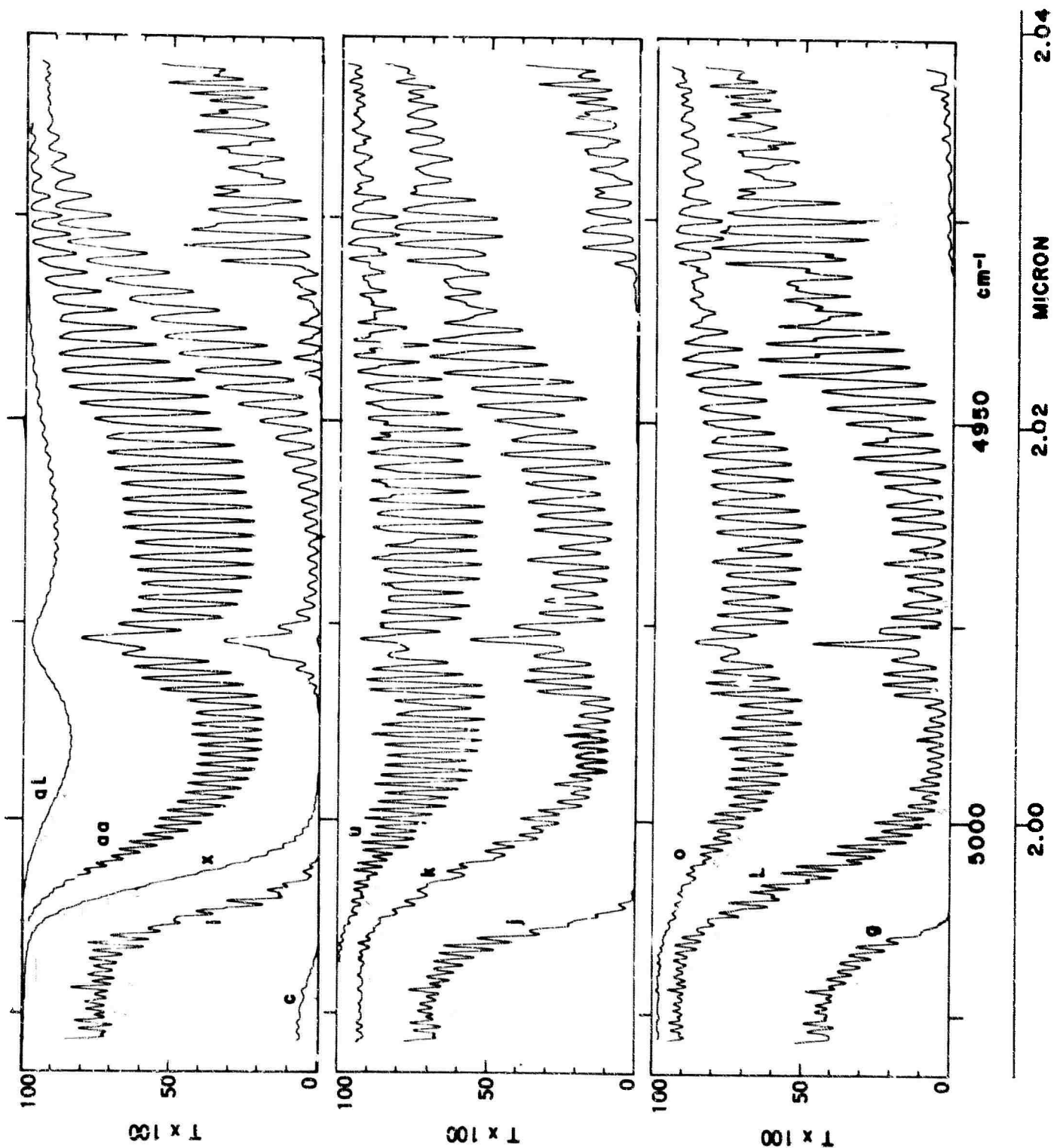


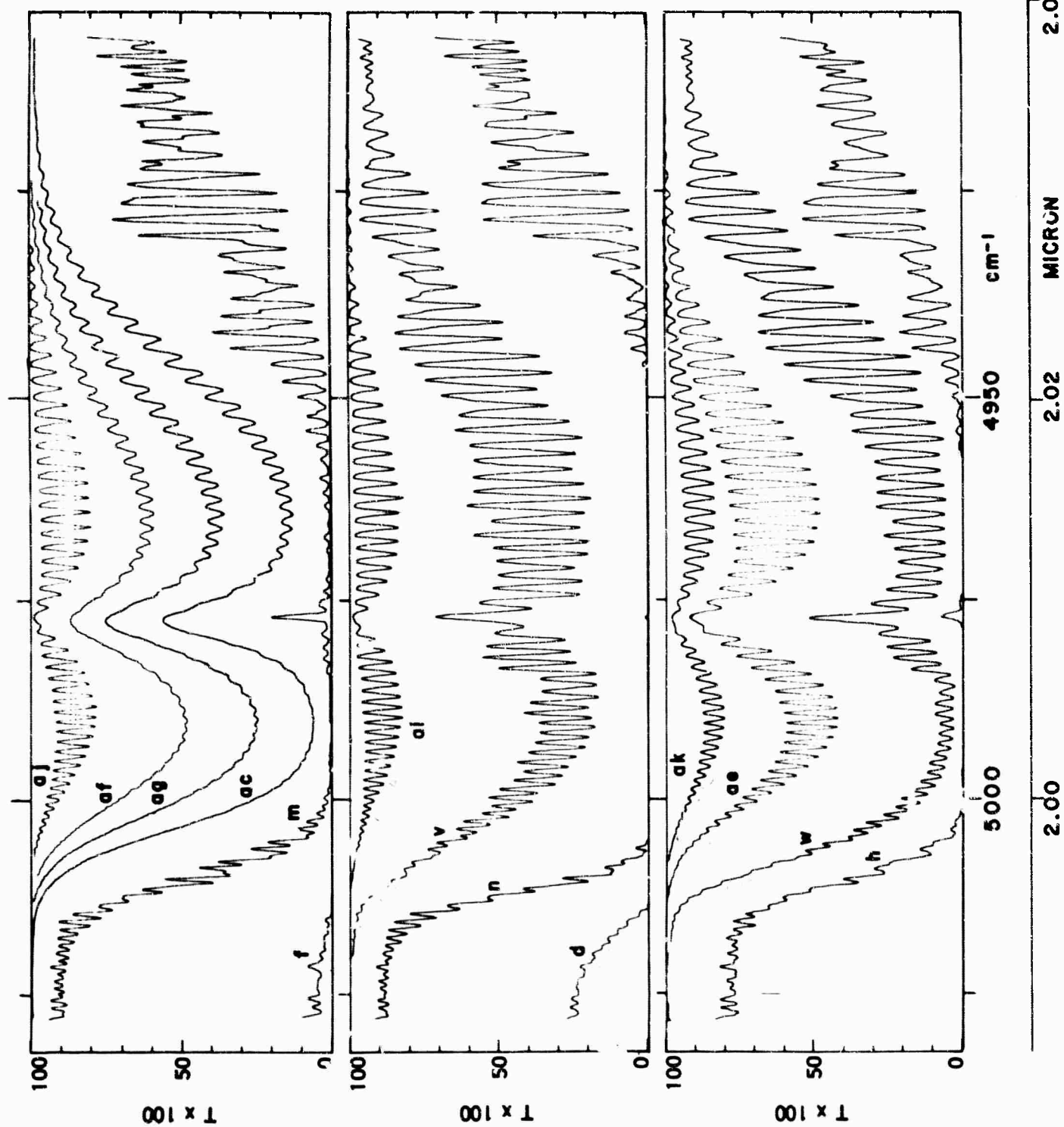


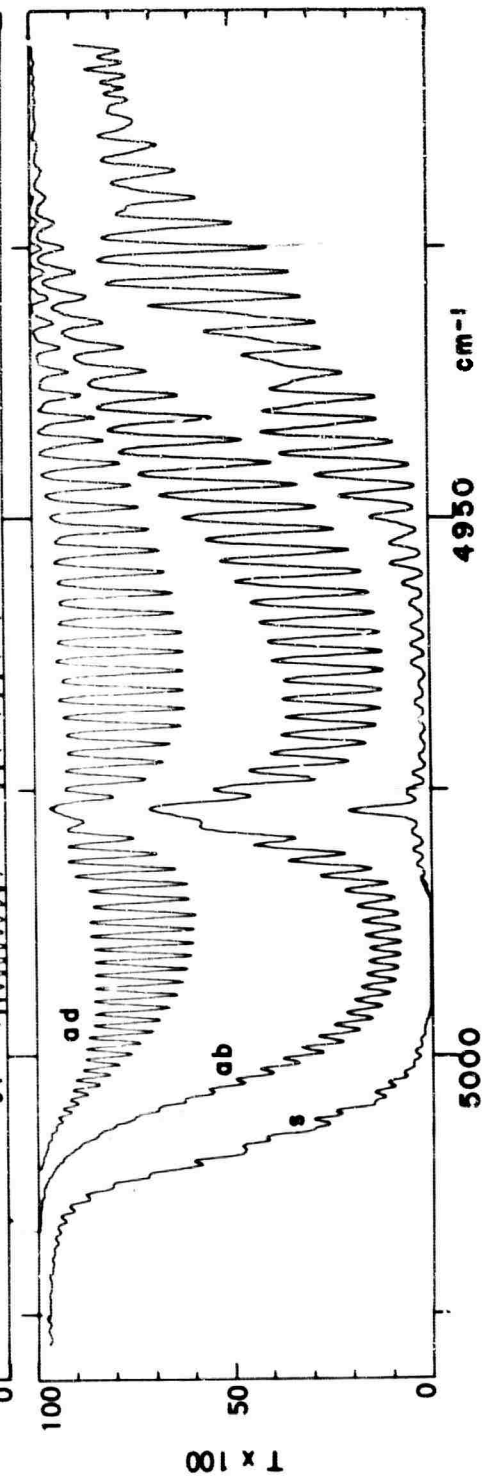
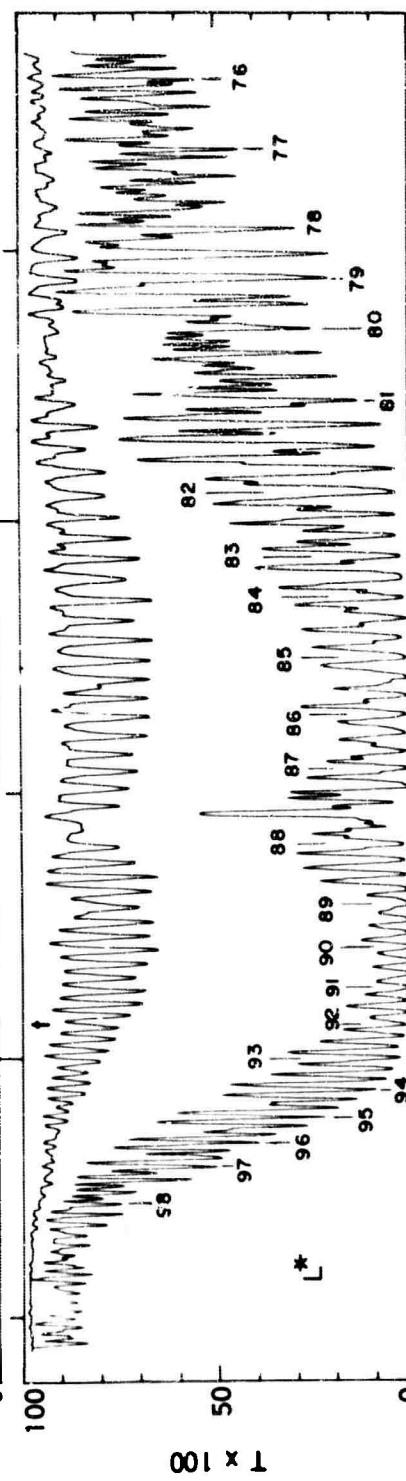
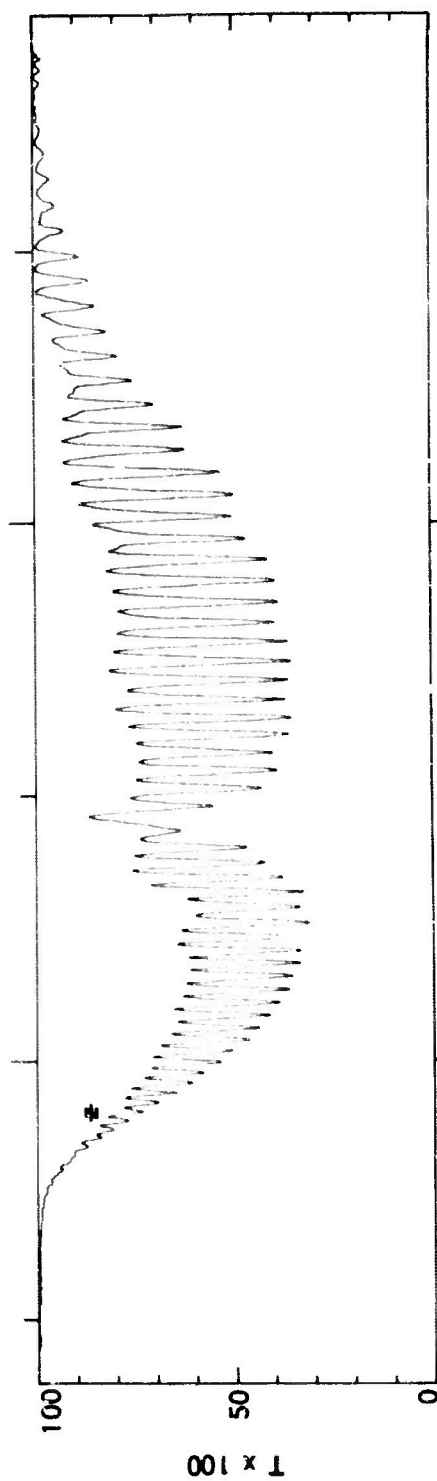


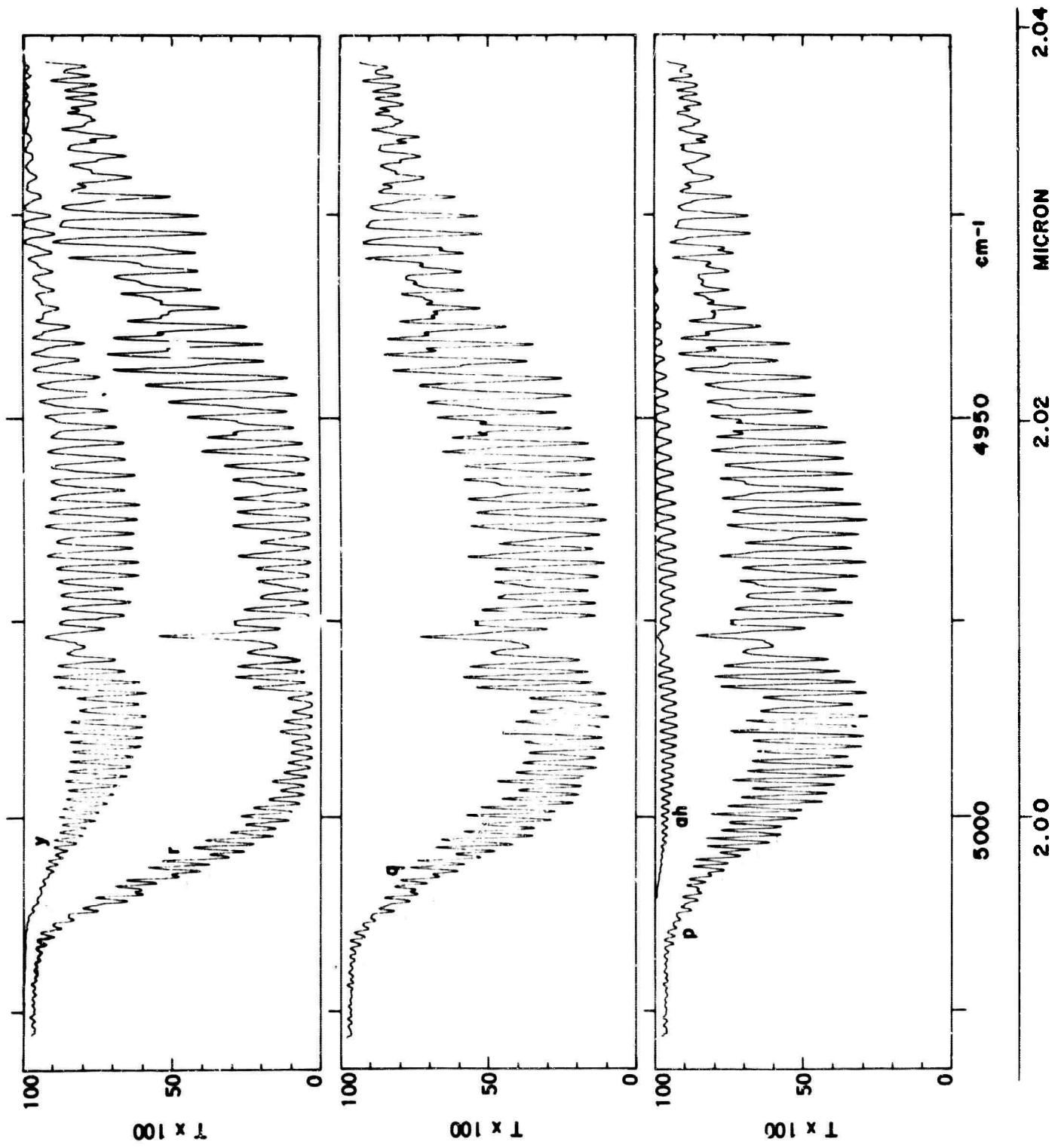


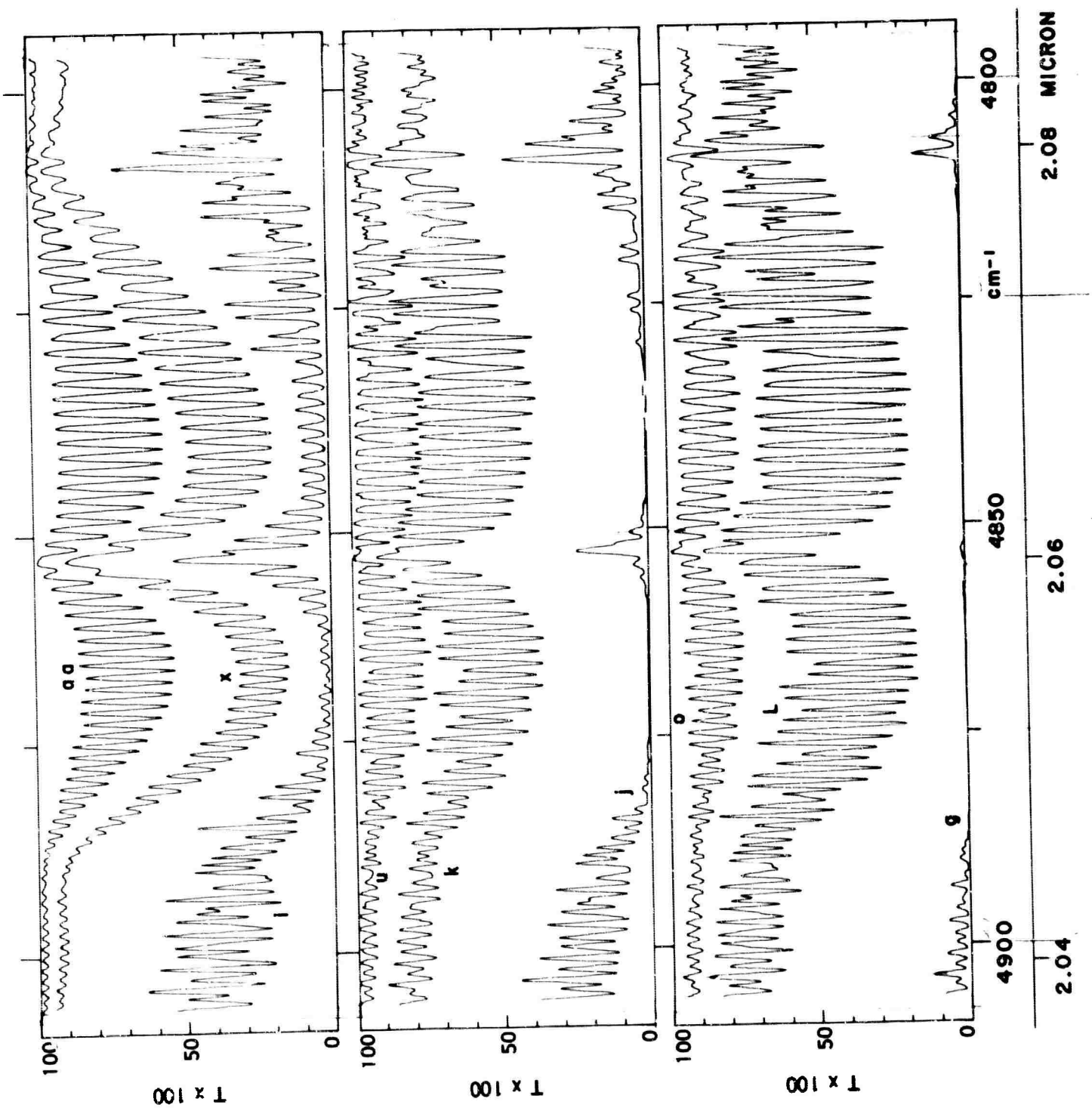


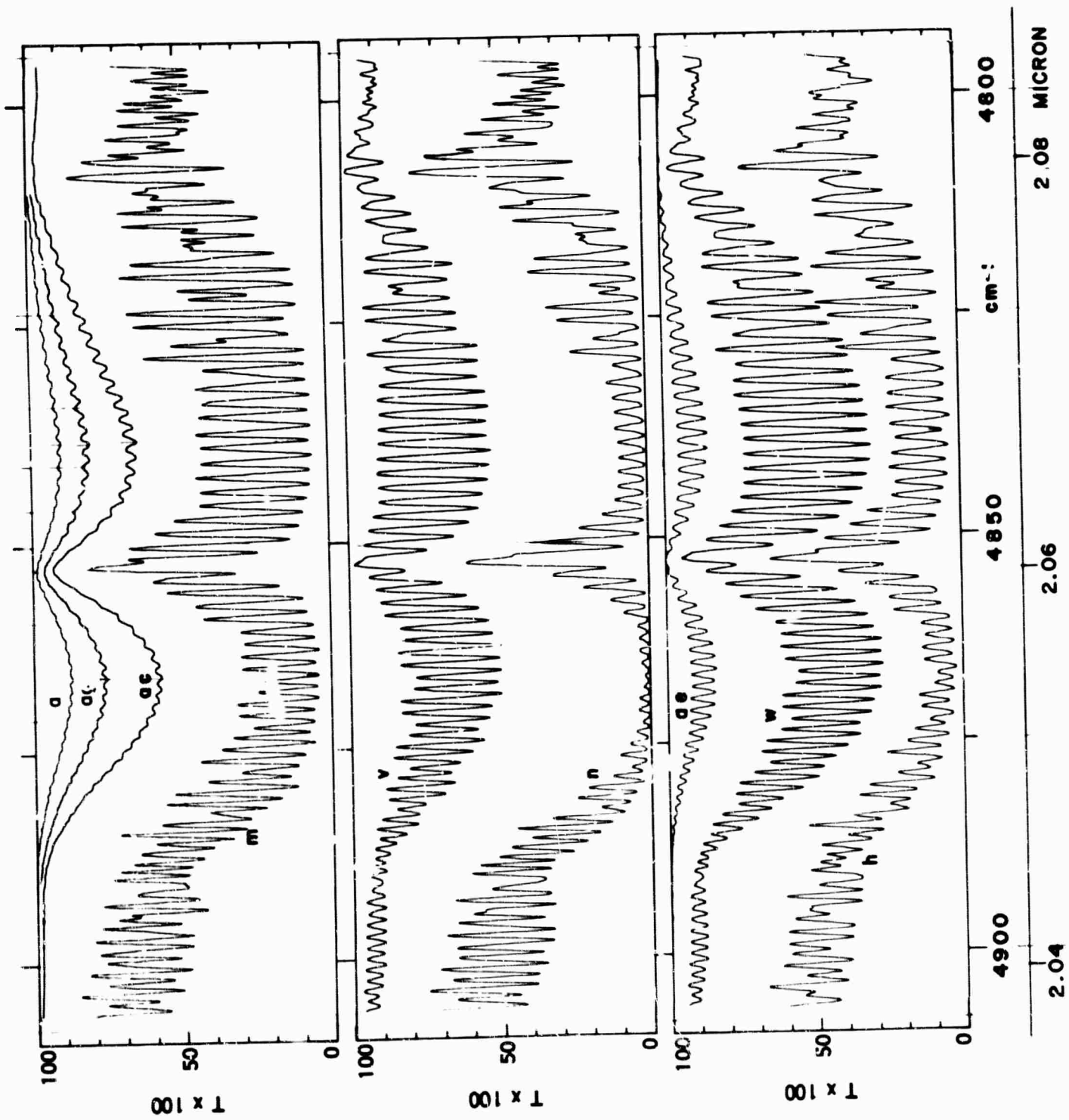


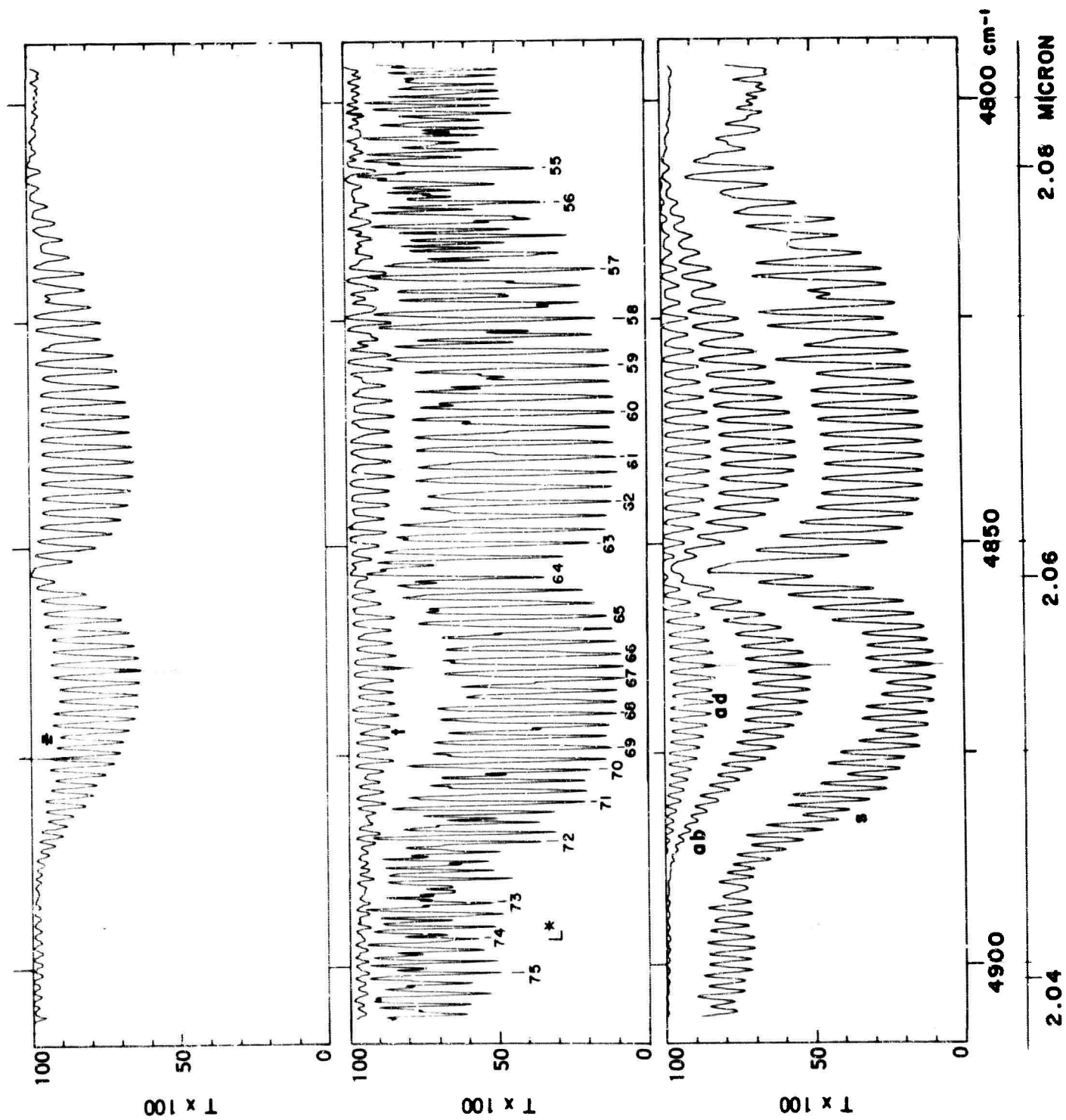


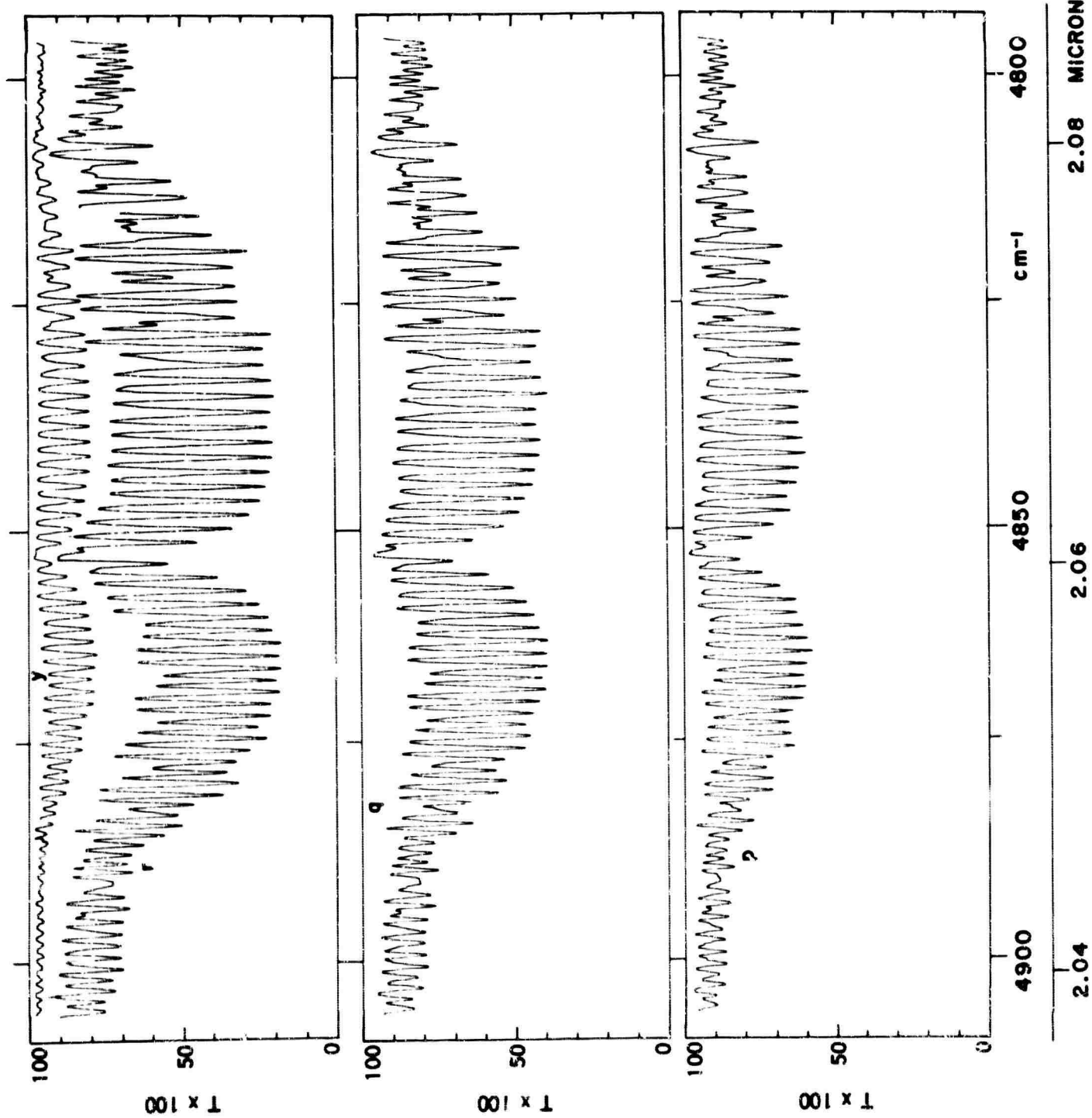


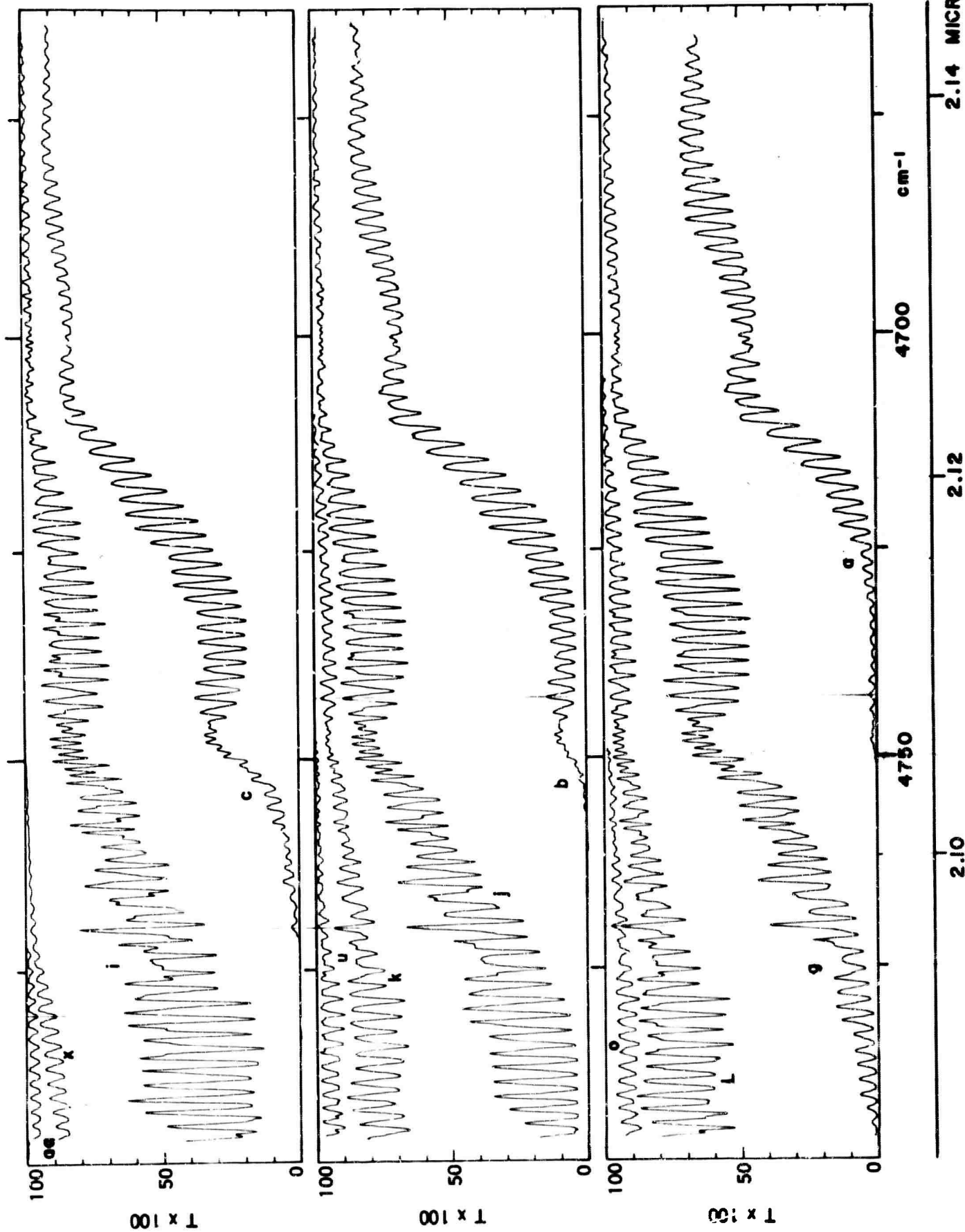


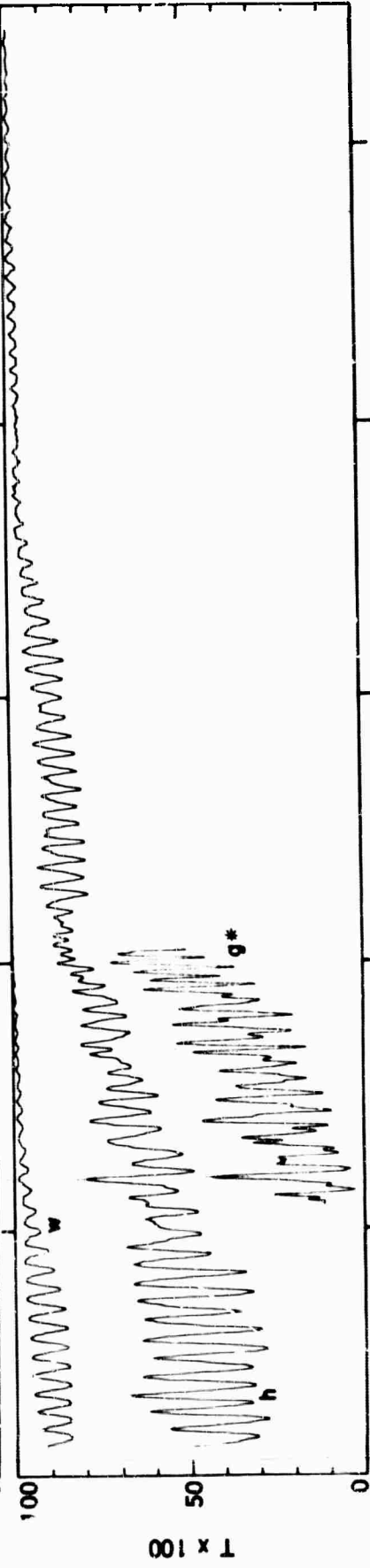
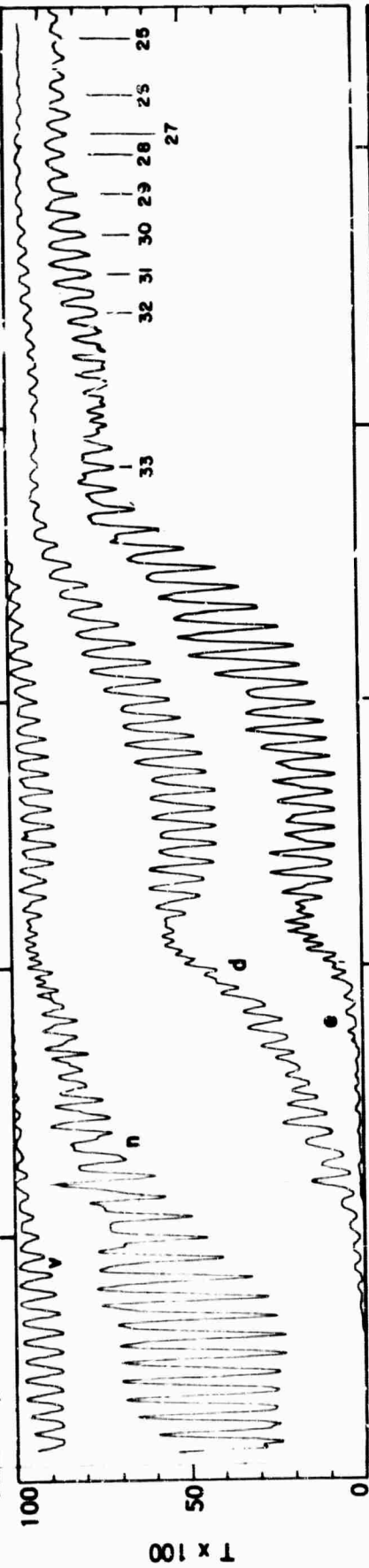
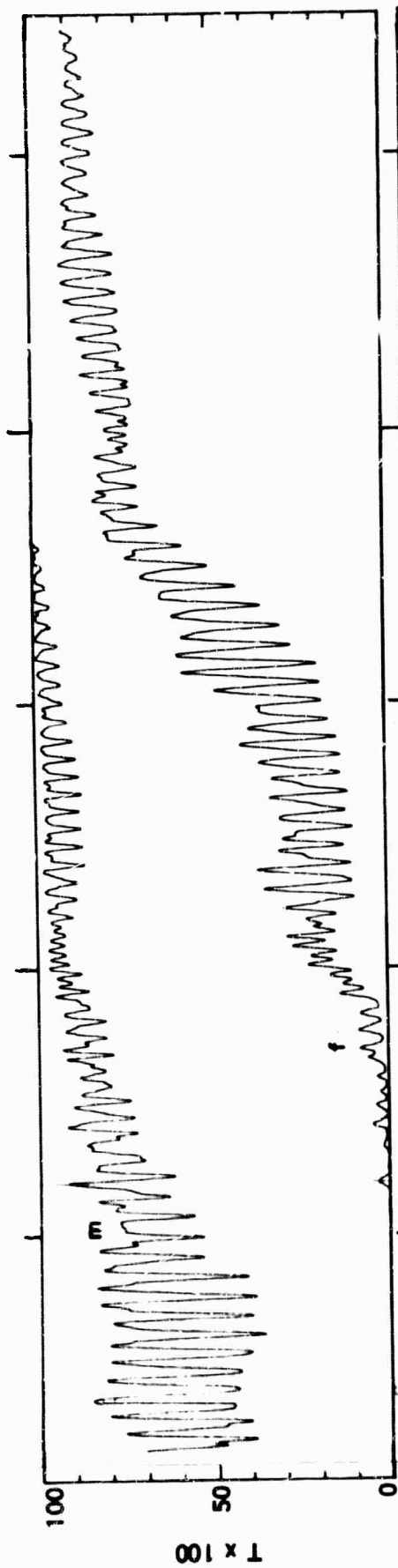








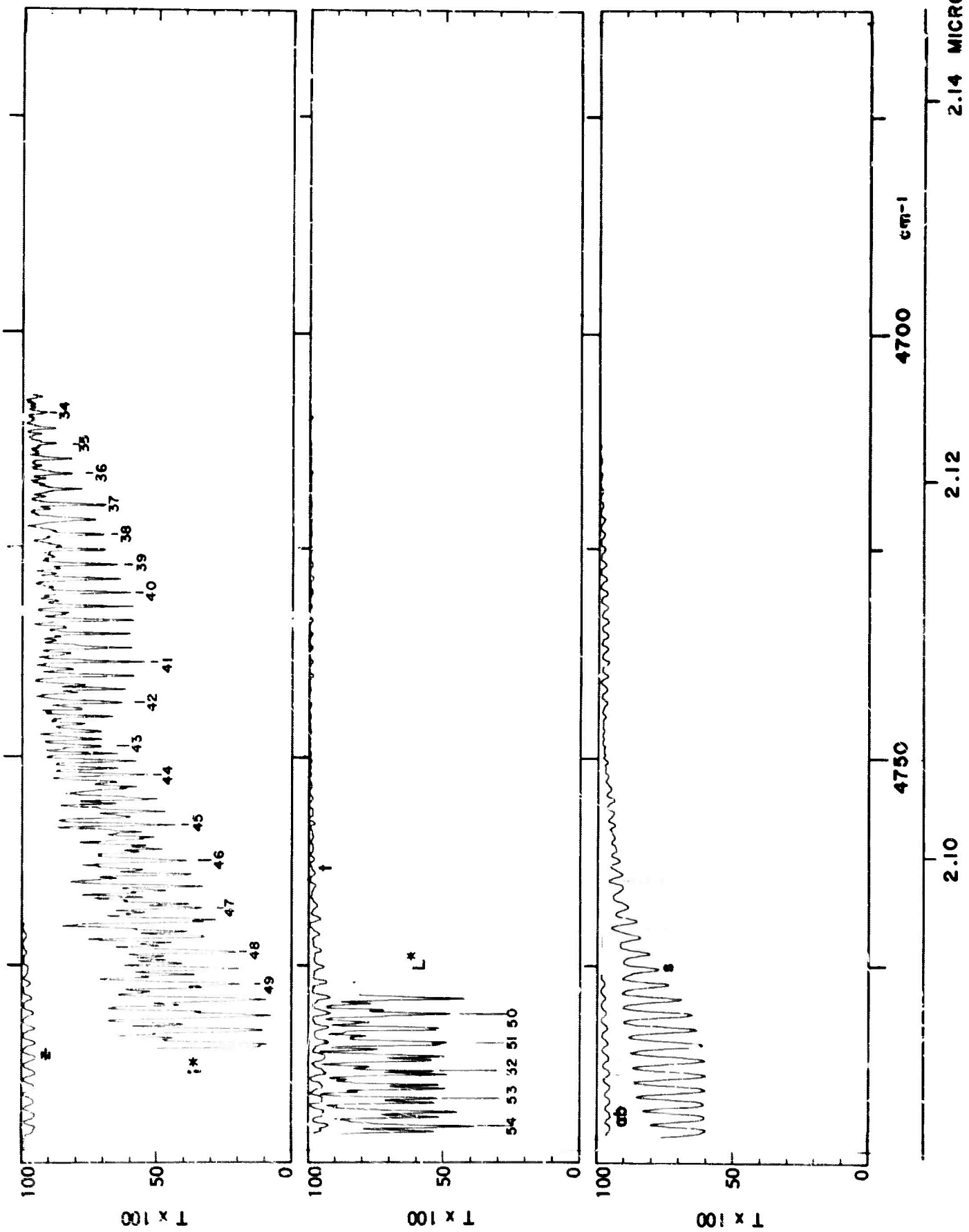


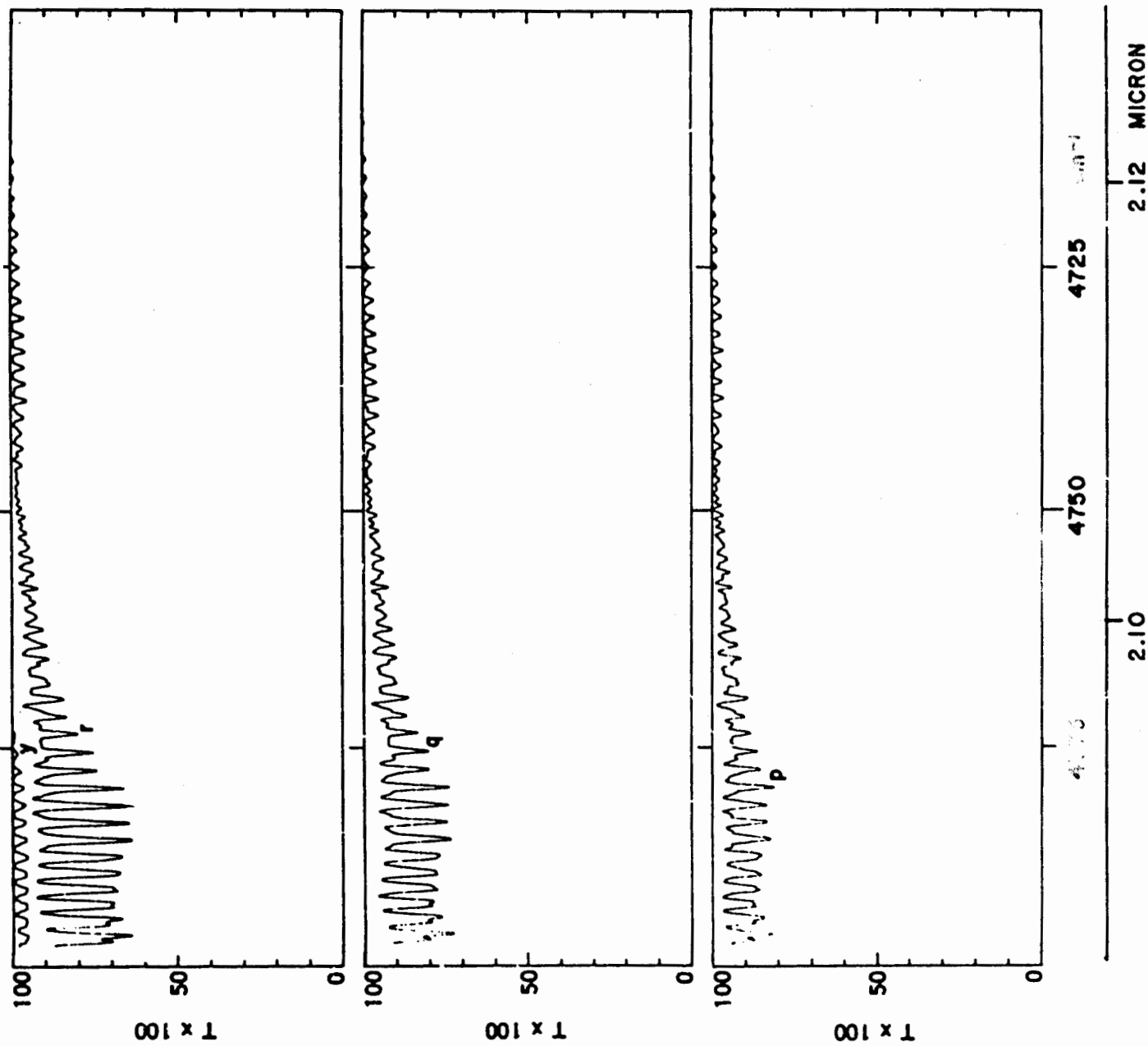


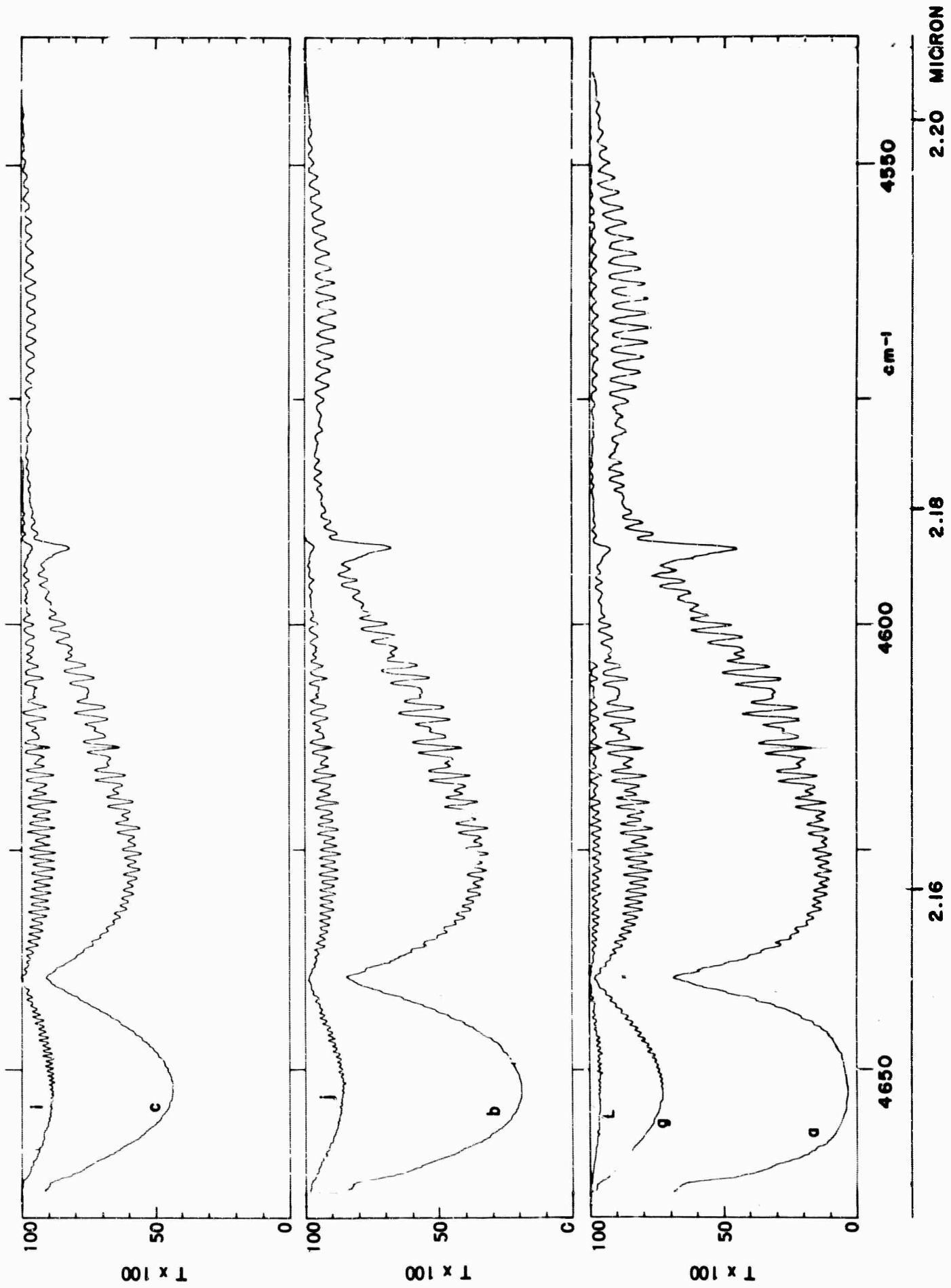
2.14 MICRON

2.12

2.10







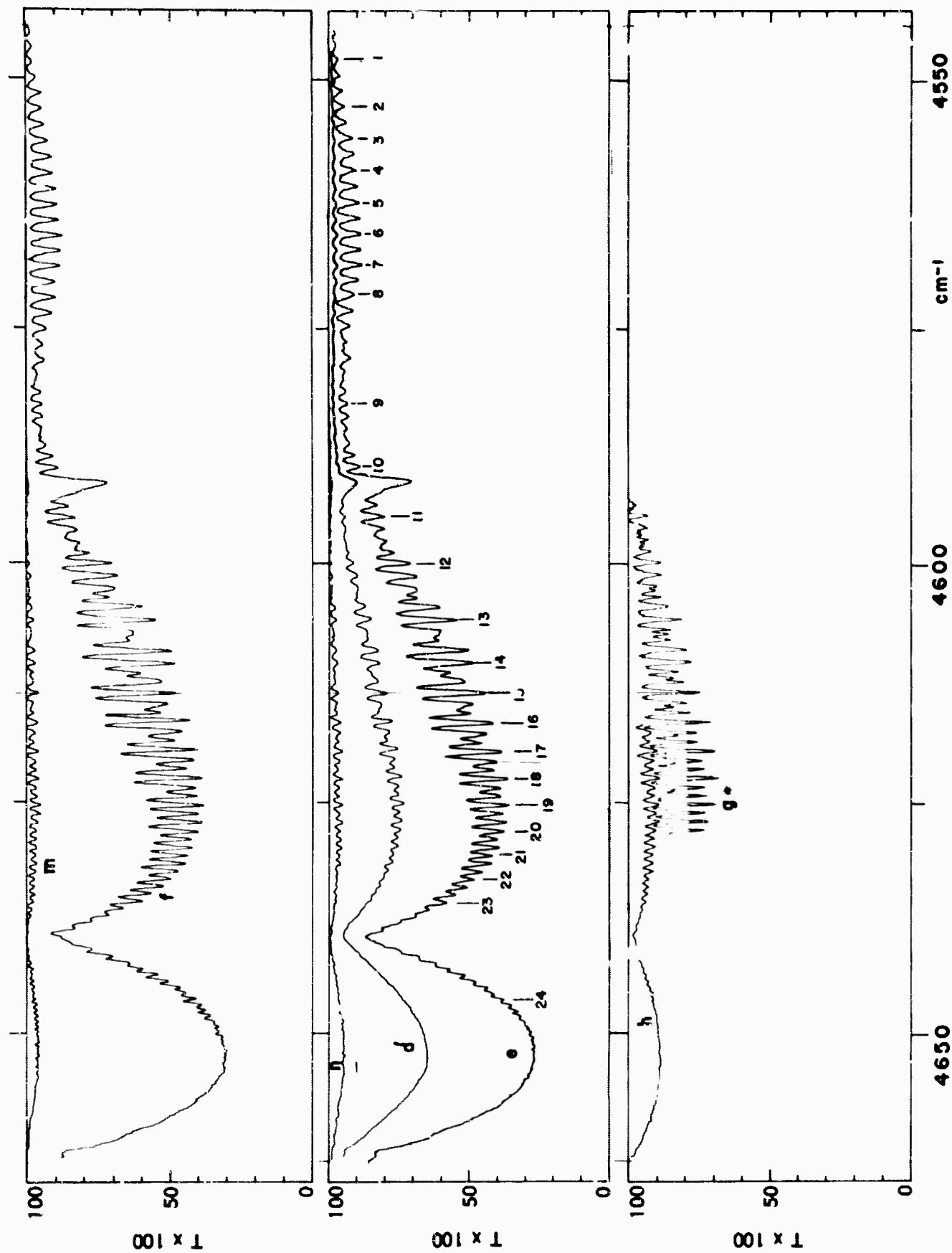


TABLE 5-1
CALIBRATION TABLE

Line No.	ν cm^{-1}	Line No.	ν cm^{-1}	Line No.	ν cm^{-1}	Line No.	ν cm^{-1}	Line No.	ν cm^{-1}
1	4548.1	36	4716.3	71	4880.7	106	5066.2	141	5206.6H
2	4553.0	37	4719.9	72	4885.4	107	5070.1	142	5218.4H
3	4556.2	38	4723.4	73	4892.8	108	5073.9	143	5225.3H
4	4559.4	39	4726.8	74	4896.9	109	5077.6	144	5236.9
5	4562.6	40	4730.2	75	4900.9	110	5081.2	145	5254.8H
6	4565.8	41	4738.5	76	4909.1	111	5084.8	146	5264.1H
7	4569.0	42	4743.3	77	4915.1	112	5088.2	147	5270.7
8	4572.1	43	4748.8	78	4922.8	113	5091.6	148	5275.7
9	4583.2	44	4751.8	79	4927.5	114	5094.9	149	5282.8
10	4589.6	45	4757.9	80	4932.1	115	5098.1	150	5287.4
11	4594.7	46	4762.3	81	4938.7	116	5101.9	151	5293.8
12	4599.7	47	4767.8	82	4947.1	117	5104.9	152	5297.9
13	4605.6	48	4773.3	83	4953.2	118	5107.8	153	5301.9
14	4610.1	49	4777.2	84	4957.0	119	5110.7	154	5305.6
15	4613.4	50	4780.9	85	4962.6	120	5113.4	155	5318.0
16	4616.6	51	4784.4	86	4967.9	121	5116.1	156	5320.8
17	4619.6	52	4787.9	87	4973.0	122	5118.6	157	5323.5
18	4622.6	53	4791.4	88	4980.1	123	5121.1	158	5326.0
19	4625.4	54	4794.7	89	4985.9	124	5123.5	159	5329.5
20	4628.1	55	4807.4	90	4989.9	125	5125.9	160	5333.4
21	4630.8	56	4811.4	91	4993.7	126	5130.3	161	5344.9H
22	4633.3	57	4819.1	92	4997.2	127	5134.5	162	5348.7H
23	4635.7	58	4824.7	93	5000.4	128	5137.2	163	5364.2H
24	4646.3	59	4830.2	94	5003.4	129	5139.8	164	5377.0H
25	4666.6	60	4835.5	95	5006.1	130	5142.4		
26	4671.6	61	4840.6	96	5008.5	131	5144.8		
27	4675.1	62	4845.6	97	5010.7	132	5147.1		
28	4676.9	63	4850.4	98	5014.2	133	5151.9		
29	4690.6	64	4854.4	99	5037.3	134	5154.1		
30	4684.1	65	4858.9	100	5041.6	135	5159.4		
31	4687.6	66	4862.3	101	5045.9	136	5163.7H		
32	4691.0	67	4866.2	102	5050.1	137	5170.5H		
33	4705.3	68	4870.4	103	5054.2	138	5176.2H		
34	4709.1	69	4874.4	104	5058.3	139	5180.6H		
35	4712.7	70	4877.0	105	5062.3	140	5184.9H		

^HThe calibration lines designated with an H are taken from H₂O;
while the remainder are from CO₂.

SECTION 6

TABLE OF TRANSMITTANCES

Table 6-1 consists of values of transmittance, in percent, recorded at intervals of 0.1 cm^{-1} in the region from 4536 to 5362 cm^{-1} . The interval is sufficiently small that the original spectrum could be approximated very closely by plotting the tabulated values and joining the points with straight lines. The first column gives the wavenumber in cm^{-1} , and the second column gives the corresponding wavelength in microns. The tables were made by photographing a portion of the computer output which was obtained from the spectra by the technique described in Appendix C.

The CO_2 partial pressure p , the equivalent pressure P_e and the absorber thickness u for each sample are shown at the top of the column corresponding to that sample. These sample parameters are repeated on alternate pages for convenience. Each sample is designated by the same letter as in Table 4-1 and in the spectra in Section 5.

Values of transmittance are not tabulated over the entire spectral region for several of the smaller samples since the absorbance in some portions is very small. All tabulated values for a given wavenumber appear on a single page. Several of the larger samples are opaque over rather wide spectral intervals; in such cases a note stating that $T = 0$ is included rather than repeating all the hundreds of zero's which would otherwise appear.

The wavenumbers of most of the line centers are accurate to $\pm 0.1 \text{ cm}^{-1}$ in the tables. Only a few are in error by more than 0.2 cm^{-1} .

THIS PAGE LEFT BLANK INTENTIONALLY

THIS PAGE LEFT BLANK INTENTIONALLY

Table 6-1

[illegible]

000000	000001	000002	000003	000004	000005	000006	000007	000008	000009	000010	000011	000012	000013	000014	000015	000016	000017	000018	000019	000020	000021	000022	000023	000024	000025	000026	000027	000028	000029
000030	000031	000032	000033	000034	000035	000036	000037	000038	000039	000040	000041	000042	000043	000044	000045	000046	000047	000048	000049	000050	000051	000052	000053	000054	000055	000056	000057	000058	000059
000060	000061	000062	000063	000064	000065	000066	000067	000068	000069	000070	000071	000072	000073	000074	000075	000076	000077	000078	000079	000080	000081	000082	000083	000084	000085	000086	000087	000088	000089
000090	000091	000092	000093	000094	000095	000096	000097	000098	000099	000100	000101	000102	000103	000104	000105	000106	000107	000108	000109	000110	000111	000112	000113	000114	000115	000116	000117	000118	000119
000120	000121	000122	000123	000124	000125	000126	000127	000128	000129	000130	000131	000132	000133	000134	000135	000136	000137	000138	000139	000140	000141	000142	000143	000144	000145	000146	000147	000148	000149
000150	000151	000152	000153	000154	000155	000156	000157	000158	000159	000160	000161	000162	000163	000164	000165	000166	000167	000168	000169	000170	000171	000172	000173	000174	000175	000176	000177	000178	000179
000180	000181	000182	000183	000184	000185	000186	000187	000188	000189	000190	000191	000192	000193	000194	000195	000196	000197	000198	000199	000200	000201	000202	000203	000204	000205	000206	000207	000208	000209
000210	000211	000212	000213	000214	000215	000216	000217	000218	000219	000220	000221	000222	000223	000224	000225	000226	000227	000228	000229	000230	000231	000232	000233	000234	000235	000236	000237	000238	000239
000240	000241	000242	000243	000244	000245	000246	000247	000248	000249	000250	000251	000252	000253	000254	000255	000256	000257	000258	000259	000260	000261	000262	000263	000264	000265	000266	000267	000268	000269
000270	000271	000272	000273	000274	000275	000276	000277	000278	000279	000280	000281	000282	000283	000284	000285	000286	000287	000288	000289	000290	000291	000292	000293	000294	000295	000296	000297	000298	000299
000300	000301	000302	000303	000304	000305	000306	000307	000308	000309	000310	000311	000312	000313	000314	000315	000316	000317	000318	000319	000320	000321	000322	000323	000324	000325	000326	000327	000328	000329
000330	000331	000332	000333	000334	000335	000336	000337	000338	000339	000340	000341	000342	000343	000344	000345	000346	000347	000348	000349	000350	000351	000352	000353	000354	000355	000356	000357	000358	000359
000360	000361	000362	000363	000364	000365	000366	00036																						

Table 6-1 (continued)

[illegible]

BLANK PAGE

4584.0	4584.1	4584.2	4584.3	4584.4	4584.5	4584.6	4584.7	4584.8	4584.9	4585.0	4585.1	4585.2	4585.3	4585.4	4585.5	4585.6	4585.7	4585.8	4585.9	4586.0	4586.1	4586.2	4586.3	4586.4	4586.5	4586.6	4586.7	4586.8	4586.9	4587.0	4587.1	4587.2	4587.3	4587.4	4587.5	4587.6	4587.7	4587.8	4587.9	4588.0	4588.1	4588.2	4588.3	4588.4	4588.5	4588.6	4588.7	4588.8	4588.9	4589.0	4589.1	4589.2	4589.3	4589.4	4589.5	4589.6	4589.7	4589.8	4589.9	4590.0	4590.1	4590.2	4590.3	4590.4	4590.5	4590.6	4590.7	4590.8	4590.9	4591.0	4591.1	4591.2	4591.3	4591.4	4591.5	4591.6	4591.7	4591.8	4591.9	4592.0	4592.1	4592.2	4592.3	4592.4	4592.5	4592.6	4592.7	4592.8	4592.9	4593.0	4593.1	4593.2	4593.3	4593.4	4593.5	4593.6	4593.7	4593.8	4593.9	4594.0	4594.1	4594.2	4594.3	4594.4	4594.5	4594.6	4594.7	4594.8	4594.9	4595.0	4595.1	4595.2	4595.3	4595.4	4595.5	4595.6	4595.7	4595.8	4595.9	4596.0	4596.1	4596.2	4596.3	4596.4	4596.5	4596.6	4596.7	4596.8	4596.9	4597.0	4597.1	4597.2	4597.3	4597.4	4597.5	4597.6	4597.7	4597.8	4597.9	4598.0	4598.1	4598.2	4598.3	4598.4	4598.5	4598.6	4598.7	4598.8	4598.9	4599.0	4599.1	4599.2	4599.3	4599.4	4599.5	4599.6	4599.7	4599.8	4599.9	4600.0	4600.1	4600.2	4600.3	4600.4	4600.5	4600.6	4600.7	4600.8	4600.9	4601.0	4601.1	4601.2	4601.3	4601.4	4601.5	4601.6	4601.7	4601.8	4601.9	4602.0	4602.1	4602.2	4602.3	4602.4	4602.5	4602.6	4602.7	4602.8	4602.9	4603.0	4603.1	4603.2	4603.3	4603.4	4603.5	4603.6	4603.7	4603.8	4603.9	4604.0	4604.1	4604.2	4604.3	4604.4	4604.5	4604.6	4604.7	4604.8	4604.9	4605.0	4605.1	4605.2	4605.3	4605.4	4605.5	4605.6	4605.7	4605.8	4605.9	4606.0	4606.1	4606.2	4606.3	4606.4	4606.5	4606.6	4606.7	4606.8	4606.9	4607.0	4607.1	4607.2	4607.3	4607.4	4607.5	4607.6	4607.7	4607.8	4607.9	4608.0	4608.1	4608.2	4608.3	4608.4	4608.5	4608.6	4608.7	4608.8	4608.9	4609.0	4609.1	4609.2	4609.3	4609.4	4609.5	4609.6	4609.7	4609.8	4609.9	4610.0	4610.1	4610.2	4610.3	4610.4	4610.5	4610.6	4610.7	4610.8	4610.9	4611.0	4611.1	4611.2	4611.3	4611.4	4611.5	4611.6	4611.7	4611.8	4611.9	4612.0	4612.1	4612.2	4612.3	4612.4	4612.5	4612.6	4612.7	4612.8	4612.9	4613.0	4613.1	4613.2	4613.3	4613.4	4613.5	4613.6	4613.7	4613.8	4613.9	4614.0	4614.1	4614.2	4614.3	4614.4	4614.5	4614.6	4614.7	4614.8	4614.9	4615.0	4615.1	4615.2	4615.3	4615.4	4615.5	4615.6	4615.7	4615.8	4615.9	4616.0	4616.1	4616.2	4616.3	4616.4	4616.5	4616.6	4616.7	4616.8	4616.9	4617.0	4617.1	4617.2	4617.3	4617.4	4617.5	4617.6	4617.7	4617.8	4617.9	4618.0	4618.1	4618.2	4618.3	4618.4	4618.5	4618.6	4618.7	4618.8	4618.9	4619.0	4619.1	4619.2	4619.3	4619.4	4619.5	4619.6	4619.7	4619.8	4619.9	4620.0	4620.1	4620.2	4620.3	4620.4	4620.5	4620.6	4620.7	4620.8	4620.9	4621.0	4621.1
--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------	--------

Table 6-1 (continued)

Star Name	RA	Dec	Distance (pc)	Parallax (mas)	Proper Motion (mas/yr)	Radial Velocity (km/s)	Mass (M _☉)	Age (Myr)	Distance (kpc)	Galactic Coordinates (l, b)	Cluster Membership
1. HD 101061	10 10 10.1	+10 10 10.1	101.061	9.900	0.000	0.000	1.000	10.1	1.01061	(10.1061, 10.1010)	Member
2. HD 101062	10 10 10.2	+10 10 10.2	101.062	9.900	0.000	0.000	1.000	10.2	1.01062	(10.1062, 10.1010)	Member
3. HD 101063	10 10 10.3	+10 10 10.3	101.063	9.900	0.000	0.000	1.000	10.3	1.01063	(10.1063, 10.1010)	Member
4. HD 101064	10 10 10.4	+10 10 10.4	101.064	9.900	0.000	0.000	1.000	10.4	1.01064	(10.1064, 10.1010)	Member
5. HD 101065	10 10 10.5	+10 10 10.5	101.065	9.900	0.000	0.000	1.000	10.5	1.01065	(10.1065, 10.1010)	Member
6. HD 101066	10 10 10.6	+10 10 10.6	101.066	9.900	0.000	0.000	1.000	10.6	1.01066	(10.1066, 10.1010)	Member
7. HD 101067	10 10 10.7	+10 10 10.7	101.067	9.900	0.000	0.000	1.000	10.7	1.01067	(10.1067, 10.1010)	Member
8. HD 101068	10 10 10.8	+10 10 10.8	101.068	9.900	0.000	0.000	1.000	10.8	1.01068	(10.1068, 10.1010)	Member
9. HD 101069	10 10 10.9	+10 10 10.9	101.069	9.900	0.000	0.000	1.000	10.9	1.01069	(10.1069, 10.1010)	Member
10. HD 101070	10 10 11.0	+10 10 11.0	101.070	9.900	0.000	0.000	1.000	11.0	1.01070	(10.1070, 10.1010)	Member
11. HD 101071	10 10 11.1	+10 10 11.1	101.071	9.900	0.000	0.000	1.000	11.1	1.01071	(10.1071, 10.1010)	Member
12. HD 101072	10 10 11.2	+10 10 11.2	101.072	9.900	0.000	0.000	1.000	11.2	1.01072	(10.1072, 10.1010)	Member
13. HD 101073	10 10 11.3	+10 10 11.3	101.073	9.900	0.000	0.000	1.000	11.3	1.01073	(10.1073, 10.1010)	Member
14. HD 101074	10 10 11.4	+10 10 11.4	101.074	9.900	0.000	0.000	1.000	11.4	1.01074	(10.1074, 10.1010)	Member
15. HD 101075	10 10 11.5	+10 10 11.5	101.075	9.900	0.000	0.000	1.000	11.5	1.01075	(10.1075, 10.1010)	Member
16. HD 101076	10 10 11.6	+10 10 11.6	101.076	9.900	0.000	0.000	1.000	11.6	1.01076	(10.1076, 10.1010)	Member
17. HD 101077	10 10 11.7	+10 10 11.7	101.077	9.900	0.000	0.000	1.000	11.7	1.01077	(10.1077, 10.1010)	Member
18. HD 101078	10 10 11.8	+10 10 11.8	101.078	9.900	0.000	0.000	1.000	11.8	1.01078	(10.1078, 10.1010)	Member
19. HD 101079	10 10 11.9	+10 10 11.9	101.079	9.900	0.000	0.000	1.000	11.9	1.01079	(10.1079, 10.1010)	Member
20. HD 101080	10 10 12.0	+10 10 12.0	101.080	9.900	0.000	0.000	1.000	12.0	1.01080	(10.1080, 10.1010)	Member
21. HD 101081	10 10 12.1	+10 10 12.1	101.081	9.900	0.000	0.000	1.000	12.1	1.01081	(10.1081, 10.1010)	Member
22. HD 101082	10 10 12.2	+10 10 12.2	101.082	9.900	0.000	0.000	1.000	12.2	1.01082	(10.1082, 10.1010)	Member
23. HD 101083	10 10 12.3	+10 10 12.3	101.083	9.900	0.000	0.000	1.000	12.3	1.01083	(10.1083, 10.1010)	Member
24. HD 101084	10 10 12.4	+10 10 12.4	101.084	9.900	0.000	0.000	1.000	12.4	1.01084	(10.1084, 10.1010)	Member
25. HD 101085	10 10 12.5	+10 10 12.5	101.085	9.900	0.000	0.000	1.000	12.5	1.01085	(10.1085, 10.1010)	Member
26. HD 101086	10 10 12.6	+10 10 12.6	101.086	9.900	0.000	0.000	1.000	12.6	1.01086	(10.1086, 10.1010)	Member
27. HD 101087	10 10 12.7	+10 10 12.7	101.087	9.900	0.000	0.000					

44307.1	44307.2	44307.3	44307.4	44307.5	44307.6	44307.7	44307.8	44307.9	44308.0	44308.1	44308.2	44308.3	44308.4	44308.5	44308.6	44308.7	44308.8	44308.9	44309.0	44309.1	44309.2	44309.3	44309.4	44309.5	44309.6	44309.7	44309.8	44309.9	44310.0	44310.1	44310.2	44310.3	44310.4	44310.5	44310.6	44310.7	44310.8	44310.9	44311.0	44311.1	44311.2	44311.3	44311.4	44311.5	44311.6	44311.7	44311.8	44311.9	44312.0	44312.1	44312.2	44312.3	44312.4	44312.5	44312.6	44312.7	44312.8	44312.9	44313.0	44313.1	44313.2	44313.3	44313.4	44313.5	44313.6	44313.7	44313.8	44313.9	44314.0	44314.1	44314.2	44314.3	44314.4	44314.5	44314.6	44314.7	44314.8	44314.9	44315.0	44315.1	44315.2	44315.3	44315.4	44315.5	44315.6	44315.7	44315.8	44315.9	44316.0	44316.1	44316.2	44316.3	44316.4	44316.5	44316.6	44316.7	44316.8	44316.9	44317.0	44317.1	44317.2	44317.3	44317.4	44317.5	44317.6	44317.7	44317.8	44317.9	44318.0	44318.1	44318.2	44318.3	44318.4	44318.5	44318.6	44318.7	44318.8	44318.9	44319.0	44319.1	44319.2	44319.3	44319.4	44319.5	44319.6	44319.7	44319.8	44319.9	44320.0	44320.1	44320.2	44320.3	44320.4	44320.5	44320.6	44320.7	44320.8	44320.9	44321.0	44321.1	44321.2	44321.3	44321.4	44321.5	44321.6	44321.7	44321.8	44321.9	44322.0	44322.1	44322.2	44322.3	44322.4	44322.5	44322.6	44322.7	44322.8	44322.9	44323.0	44323.1	44323.2	44323.3	44323.4	44323.5	44323.6	44323.7	44323.8	44323.9	44324.0	44324.1	44324.2	44324.3	44324.4	44324.5	44324.6	44324.7	44324.8	44324.9	44325.0	44325.1	44325.2	44325.3	44325.4	44325.5	44325.6	44325.7	44325.8	44325.9	44326.0	44326.1	44326.2	44326.3	44326.4	44326.5	44326.6	44326.7	44326.8	44326.9	44327.0	44327.1	44327.2	44327.3	44327.4	44327.5	44327.6	44327.7	44327.8	44327.9	44328.0	44328.1	44328.2	44328.3	44328.4	44328.5	44328.6	44328.7	44328.8	44328.9	44329.0	44329.1	44329.2	44329.3	44329.4	44329.5	44329.6	44329.7	44329.8	44329.9	44330.0	44330.1	44330.2	44330.3	44330.4	44330.5	44330.6	44330.7	44330.8	44330.9	44331.0	44331.1	44331.2	44331.3	44331.4	44331.5	44331.6	44331.7	44331.8	44331.9	44332.0	44332.1	44332.2	44332.3	44332.4	44332.5	44332.6	44332.7	44332.8	44332.9	44333.0	44333.1	44333.2	44333.3	44333.4	44333.5	44333.6	44333.7	44333.8	44333.9	44334.0	44334.1	44334.2	44334.3	44334.4	44334.5	44334.6	44334.7	44334.8	44334.9	44335.0	44335.1	44335.2	44335.3	44335.4	44335.5	44335.6	44335.7	44335.8	44335.9	44336.0	44336.1	44336.2	44336.3	44336.4	44336.5	44336.6	44336.7	44336.8	44336.9	44337.0	44337.1	44337.2	44337.3	44337.4	44337.5	44337.6	44337.7	44337.8	44337.9	44338.0	44338.1	44338.2	44338.3	44338.4	44338.5	44338.6	44338.7	44338.8	44338.9	44339.0	44339.1	44339.2	44339.3	44339.4	44339.5	44339.6	44339.7	44339.8	44339.9	44340.0	44340.1	44340.2	44340.3	44340.4	44340.5	44340.6	44340.7	44340.8	44340.9	44341.0	44341.1
---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------

Table 6-1 (continued)

[illegible]

Table 6-1 (continued)

[illegible]

Table 6-1 (continued)

[illegible]

Table 6-1 (continued)

[illegible]

[illegible]

Table 6-1 (continued)

[illegible]

Table 6-1 (continued)

[illegible]

P-00000	P-00001	P-00002	P-00003	P-00004	P-00005	P-00006	P-00007	P-00008	P-00009	P-00010	P-00011	P-00012	P-00013	P-00014	P-00015	P-00016	P-00017	P-00018	P-00019	P-00020	P-00021	P-00022	P-00023	P-00024	P-00025	P-00026	P-00027	P-00028	P-00029	P-00030	P-00031	P-00032	P-00033	P-00034	P-00035	P-00036	P-00037	P-00038	P-00039	P-00040	P-00041	P-00042	P-00043	P-00044	P-00045	P-00046	P-00047	P-00048	P-00049	P-00050	P-00051	P-00052	P-00053	P-00054	P-00055	P-00056	P-00057	P-00058	P-00059	P-00060	P-00061	P-00062	P-00063	P-00064	P-00065	P-00066	P-00067	P-00068	P-00069	P-00070	P-00071	P-00072	P-00073	P-00074	P-00075	P-00076	P-00077	P-00078	P-00079	P-00080	P-00081	P-00082	P-00083	P-00084	P-00085	P-00086	P-00087	P-00088	P-00089	P-00090	P-00091	P-00092	P-00093	P-00094	P-00095	P-00096	P-00097	P-00098	P-00099
P-00100	P-00101	P-00102	P-00103	P-00104	P-00105	P-00106	P-00107	P-00108	P-00109	P-00110	P-00111	P-00112	P-00113	P-00114	P-00115	P-00116	P-00117	P-00118	P-00119	P-00120	P-00121	P-00122	P-00123	P-00124	P-00125	P-00126	P-00127	P-00128	P-00129	P-00130	P-00131	P-00132	P-00133	P-00134	P-00135	P-00136	P-00137	P-00138	P-00139	P-00140	P-00141	P-00142	P-00143	P-00144	P-00145	P-00146	P-00147	P-00148	P-00149	P-00150	P-00151	P-00152	P-00153	P-00154	P-00155	P-00156	P-00157	P-00158	P-00159	P-00160	P-00161	P-00162	P-00163	P-00164	P-00165	P-00166	P-00167	P-00168	P-00169	P-00170	P-00171	P-00172	P-00173	P-00174	P-00175	P-00176	P-00177	P-00178	P-00179	P-00180	P-00181	P-00182	P-00183	P-00184	P-00185	P-00186	P-00187	P-00188	P-00189	P-00190	P-00191	P-00192	P-00193	P-00194	P-00195	P-00196	P-00197	P-00198	P-00199
P-00200	P-00201	P-00202	P-00203	P-00204	P-00205	P-00206	P-00207	P-00208	P-00209	P-00210	P-00211	P-00212	P-00213	P-00214	P-00215	P-00216	P-00217	P-00218	P-00219	P-00220	P-00221	P-00222	P-00223	P-00224	P-00225	P-00226	P-00227	P-00228	P-00229	P-00230	P-00231	P-00232	P-00233	P-00234	P-00235	P-00236	P-00237	P-00238	P-00239	P-00240	P-00241	P-00242	P-00243	P-00244	P-00245	P-00246	P-00247	P-00248	P-00249	P-00250	P-00251	P-00252	P-00253	P-00254	P-00255	P-00256	P-00257	P-00258	P-00259	P-00260	P-00261	P-00262	P-00263	P-00264	P-00265	P-00266	P-00267	P-00268	P-00269	P-00270	P-00271	P-00272	P-00273	P-00274	P-00275	P-00276	P-00277	P-00278	P-00279	P-00280	P-00281	P-00282	P-00283	P-00284	P-00285	P-00286	P-00287	P-00288	P-00289	P-00290	P-00291	P-00292	P-00293	P-00294	P-00295	P-00296	P-00297	P-00298	P-00299
P-00300	P-00301	P-00302	P-00303	P-00304	P-00305	P-00306	P-00307	P-00308	P-00309	P-00310	P-00311	P-00312	P-00313	P-00314	P-00315	P-00316	P-00317	P-00318	P-00319	P-00320	P-00321	P-00322	P-00323	P-00324	P-00325	P-00326	P-00327	P-00328	P-00329	P-00330	P-00331	P-00332	P-00333	P-00334	P-00335	P-00336	P-00337	P-00338	P-00339	P-00340	P-00341	P-00342	P-00343	P-00344	P-00345	P-00346	P-00347	P-00348	P-00349	P-00350	P-00351	P-00352	P-00353	P-00354	P-00355	P-00356	P-00357	P-00358	P-00359	P-00360	P-00361	P-00362	P-00363	P-00364	P-00365	P-00366	P-00367	P-00368	P-00369	P-00370	P-00371	P-00372	P-00373	P-00374	P-00375	P-00376	P-00377	P-00378	P-00379	P-00380	P-00381	P-00382	P-00383	P-00384	P-00385	P-00386	P-00387	P-00388	P-00389	P-00390	P-00391	P-00392	P-00393	P-00394	P-00395	P-00396	P-00397	P-00398	P-00399
P-00400	P-00401	P-00402	P-00403	P-00404	P-00405	P-00406	P-00407	P-00408	P-00409	P-00410	P-00411	P-00412	P-00413	P-00414	P-00415	P-00416	P-00417	P-00418	P-00419	P-00420	P-00421	P-00422	P-00423	P-00424	P-00425	P-00426	P-00427	P-00428	P-00429	P-00430	P-00431	P-00432	P-00433	P-00434	P-00435	P-00436	P-00437	P-00438	P-00439	P-00440	P-00441	P-00442	P-00443	P-00444	P-00445	P-00446	P-00447	P-00448	P-00449	P-00450	P-00451	P-00452	P-00453	P-00454	P-00455	P-00456	P-00457	P-00458	P-00459	P-00460	P-00461	P-00462	P-00463	P-00464	P-00465	P-00466	P-00467	P-00468	P-00469	P-00470	P-00471	P-00472	P-00473	P-00474	P-00475	P-00476	P-00477	P-00478	P-00479	P-00480	P-00481	P-00482	P-00483	P-00484	P-00485	P-00486	P-00487	P-00488	P-00489	P-00490	P-00491	P-00492	P-00493	P-00494	P-00495	P-00496	P-00497	P-00498	P-00499
P-00500	P-00501	P-00502	P-00503	P-00504	P-00505	P-00506	P-00507	P-00508	P-00509	P-00510	P-00511	P-00512	P-00513	P-00514	P-00515	P-00516	P-00517	P-00518	P-00519	P-00520	P-00521	P-00522	P-00523	P-00524	P-00525	P-00526	P-00527	P-00528	P-00529	P-00530	P-00531	P-00532	P-00533	P-00534	P-00535	P-00536	P-00537	P-00538	P-00539	P-00540	P-00541	P-00542	P-00543	P-00544	P-00545	P-00546	P-00547	P-00548	P-00549	P-00550	P-00551	P-00552	P-00553	P-00554	P-00555	P-00556	P-00557	P-00558	P-00559	P-00560	P-00561	P-00562	P-00563	P-00564	P-00565	P-00566	P-00567	P-00568	P-00569	P-00570	P-00571	P-00572	P-00573	P-00574	P-00575	P-00576	P-00577	P-00578	P-00579	P-00580	P-00581	P-00582	P-00583	P-00584	P-00585	P-00586	P-00587	P-00588	P-00589	P-00590	P-00591	P-00592	P-00593	P-00594	P-00595	P-00596	P-00597	P-00598	P-00599
P-00600	P-00601	P-00602	P-00603	P-00604	P-00605	P-00606	P-00607	P-00608	P-00609	P-00610	P-00611	P-00612	P-00613	P-00614	P-00615	P-00616	P-00617	P-00618	P-00619	P-00620	P-00621	P-00622	P-00623	P-00624	P-00625	P-00626	P-00627	P-00628	P-00629	P-00630	P-00631	P-00632	P-00633	P-00634	P-00635	P-00636	P-00637	P-00638	P-00639	P-00640	P-00641	P-00642	P-00643	P-00644	P-00645	P-00646	P-00647	P-00648	P-00649	P-00650	P-00651	P-00652	P-00653	P-00654	P-00655	P-00656	P-00657	P-00658	P-00659	P-00660	P-00661	P-00662	P-00663	P-00664	P-00665	P-00666	P-00667	P-00668	P-00669	P-00670	P-00671	P-00672	P-00673	P-00674	P-00675	P-00676	P-00677	P-00678	P-00679	P-00680	P-00681	P-00682	P-00683	P-00684	P-00685	P-00686	P-00687	P-00688	P-00689	P-00690	P-00691	P-00692	P-00693	P-00694	P-00695	P-00696	P-00697	P-00698	P-00699
P-00700	P-00701	P-00702	P-00703	P-00704	P-00705	P-00706	P-00707	P-00708	P-00709	P-00710	P-00711	P-00712	P-00713	P-00714	P-00715	P-00716	P-00717	P-00718	P-00719	P-00720	P-00721	P-00722	P-00723	P-00724	P-00725	P-00726	P-00727	P-00728	P-00729	P-00730	P-00731	P-00732	P-00733	P-00734	P-00735	P-00736	P-00737	P-00738	P-00739	P-00740	P-00741	P-00742	P-00743	P-00744	P-00745	P-00746	P-00747	P-00748	P-00749	P-00750	P-00751	P-00752	P-00753	P-00754	P-00755	P-00756	P-00757	P-00758	P-00759	P-00760	P-00761	P-00762	P-00763	P-00764	P-00765	P-00766	P-00767	P-00768	P-00769	P-00770	P-00771	P-00772	P-00773	P-00774	P-00775	P-00776	P-00777	P-00778	P-00779	P-00780	P-00781	P-00782	P-00783	P-00784	P-00785	P-00786	P-00787	P-00788	P-00789	P-00790	P-00791	P-00792	P-00793	P-00794	P-00795	P-00796	P-00797	P-00798	P-00799
P-00800	P-00801	P-00802	P-00803	P-00804	P-00805	P-00806	P-00807	P-00808	P-00809	P-00810	P-00811	P-00812	P-00813	P-00814	P-00815	P-00816	P-00817	P-00818	P-00819	P-00820	P-00821	P-00822	P-00823	P-00824	P-00825	P-00826	P-00827	P-00828	P-00829	P-00830	P-00831	P-00832	P-00833	P-00834	P-00835	P-00836	P-00837	P-00838	P-00839	P-00840	P-00841	P-00842	P-00843	P-00844	P-00845	P-00846	P-00847	P-00848	P-00849	P-00850	P-00851	P-00852	P-00853	P-00854	P-00855	P-00856	P-00857	P-00858	P-00859	P-00860	P-00861	P-00862	P-00863	P-00864	P-00865	P-00866	P-00867	P-00868	P-00869	P-00870	P-00871	P-00872	P-00873	P-00874	P-00875	P-00876	P-00877	P-00878	P-00879	P-00880	P-00881	P-00882	P-00883	P-00884	P-00885	P-00886	P-00887	P-00888	P-00889	P-00890	P-00891	P-00892	P-00893	P-00894	P-00895	P-00896	P-00897	P-00898	P-00899
P-00900	P-00901	P-00902	P-00903	P-00904	P-00905	P-00906	P-00907	P-00908	P-00909	P-00910	P-00911	P-00912	P-00913	P-00914	P-00915	P-00916	P-00917	P-00918	P-00919	P-00920	P-00921	P-00922	P-00923	P-00924	P-00925	P-00926	P-00927	P-00928	P-00929	P-00930	P-00931	P-00932	P-00933	P-00934	P-00935	P-00936	P-00937	P-00938	P-00939	P-00940	P-00941	P-00942	P-00943	P-00944	P-00945	P-00946	P-00947	P-00948	P-00949	P-00950	P-00951	P-00952	P-00953	P-00954	P-00955	P-00956	P-00957	P-00958	P-00959	P-00960	P-00961	P-00962	P-00963	P-00964	P-00965	P-00966	P-00967	P-00968	P-00969	P-00970	P-00971	P-00972	P-00973	P-00974	P-00975	P-00976	P-00977	P-00978	P-00979	P-00980	P-00981	P-00982	P-00983	P-00984	P-00985	P-00986	P-00987	P-00988	P-00989	P-00990	P-00991	P-00992	P-00993	P-00994	P-00995	P-00996	P-00997	P-00998	P-00999
P-01000	P-01001	P-01002	P-01003	P-01004	P-01005	P-01006	P-01007	P-01008	P-01009	P-01010	P-01011	P-01012	P-01013	P-01014	P-01015	P-01016	P-01017	P-01018	P-01019	P-01020	P-01021	P-01022	P-01023	P-01024	P-01025	P-01026	P-01027	P-01028	P-01029	P-01030	P-01031	P-01032	P-01033	P-01034	P-01035	P-01036	P-01037	P-01038	P-01039	P-01040	P-01041	P-01042	P-01043	P-01044	P-01045	P-01046	P-01047	P-01048	P-01049	P-01050	P-01051	P-01052	P-01053	P-01054	P-01055	P-01056	P-01057	P-01058	P-01059	P-01060	P-01061	P-01062	P-01063	P-01064	P-01065	P-01066	P-01067	P-01068	P-01069	P-01070	P-01071	P-01072	P-01073	P-01074	P-01075	P-01076	P-01077	P-01078	P-01079	P-01080	P-01081	P-01082	P-01083	P-01084	P-01085	P-01086	P-01087	P-01088	P-01089	P-01090	P-01091	P-01092	P-01093	P-01094	P-01095	P-01096	P-01097	P-01098	P-01099
P-01100	P-01101	P-01102	P-01103	P-01104	P-01105	P-01106	P-01107	P-01108	P-01109	P-01110	P-01111	P-01112	P-01113	P-01114	P-01115	P-01116	P-01117	P-01118	P-01119	P-01120	P-01																																																																														

Table 6-1 (continued)

[illegible]

Table 6-1 (continued)

[illegible]

Table 6-1 (continued)

[illegible]

● 4 ● 2016 年第 12 期

Table 6-1 (continued)

[illegible]

44881.0	44881.1	44881.2	44881.3	44881.4	44881.5	44881.6	44881.7	44881.8	44881.9	44882.0	44882.1	44882.2	44882.3	44882.4	44882.5	44882.6	44882.7	44882.8	44882.9	44883.0	44883.1	44883.2	44883.3	44883.4	44883.5	44883.6	44883.7	44883.8	44883.9	44884.0	44884.1	44884.2	44884.3	44884.4	44884.5	44884.6	44884.7	44884.8	44884.9	44885.0	44885.1	44885.2	44885.3	44885.4	44885.5	44885.6	44885.7	44885.8	44885.9	44886.0	44886.1	44886.2	44886.3	44886.4	44886.5	44886.6	44886.7	44886.8	44886.9	44887.0	44887.1	44887.2	44887.3	44887.4	44887.5	44887.6	44887.7	44887.8	44887.9	44888.0	44888.1	44888.2	44888.3	44888.4	44888.5	44888.6	44888.7	44888.8	44888.9	44889.0	44889.1	44889.2	44889.3	44889.4	44889.5	44889.6	44889.7	44889.8	44889.9	44890.0	44890.1	44890.2	44890.3	44890.4	44890.5	44890.6	44890.7	44890.8	44890.9	44891.0	44891.1	44891.2	44891.3	44891.4	44891.5	44891.6	44891.7	44891.8	44891.9	44892.0	44892.1	44892.2	44892.3	44892.4	44892.5	44892.6	44892.7	44892.8	44892.9	44893.0	44893.1	44893.2	44893.3	44893.4	44893.5	44893.6	44893.7	44893.8	44893.9	44894.0	44894.1	44894.2	44894.3	44894.4	44894.5	44894.6	44894.7	44894.8	44894.9	44895.0	44895.1	44895.2	44895.3	44895.4	44895.5	44895.6	44895.7	44895.8	44895.9	44896.0	44896.1	44896.2	44896.3	44896.4	44896.5	44896.6	44896.7	44896.8	44896.9	44897.0	44897.1	44897.2	44897.3	44897.4	44897.5	44897.6	44897.7	44897.8	44897.9	44898.0	44898.1	44898.2	44898.3	44898.4	44898.5	44898.6	44898.7	44898.8	44898.9	44899.0	44899.1	44899.2	44899.3	44899.4	44899.5	44899.6	44899.7	44899.8	44899.9	44900.0	44900.1	44900.2	44900.3	44900.4	44900.5	44900.6	44900.7	44900.8	44900.9	44901.0	44901.1	44901.2	44901.3	44901.4	44901.5	44901.6	44901.7	44901.8	44901.9	44902.0	44902.1	44902.2	44902.3	44902.4	44902.5	44902.6	44902.7	44902.8	44902.9	44903.0	44903.1	44903.2	44903.3	44903.4	44903.5	44903.6	44903.7	44903.8	44903.9	44904.0	44904.1	44904.2	44904.3	44904.4	44904.5	44904.6	44904.7	44904.8	44904.9	44905.0	44905.1	44905.2	44905.3	44905.4	44905.5	44905.6	44905.7	44905.8	44905.9	44906.0	44906.1	44906.2	44906.3	44906.4	44906.5	44906.6	44906.7	44906.8	44906.9	44907.0	44907.1	44907.2	44907.3	44907.4	44907.5	44907.6	44907.7	44907.8	44907.9	44908.0	44908.1	44908.2	44908.3	44908.4	44908.5	44908.6	44908.7	44908.8	44908.9	44909.0	44909.1	44909.2	44909.3	44909.4	44909.5	44909.6	44909.7	44909.8	44909.9	44910.0	44910.1	44910.2	44910.3	44910.4	44910.5	44910.6	44910.7	44910.8	44910.9	44911.0	44911.1	44911.2	44911.3	44911.4	44911.5	44911.6	44911.7	44911.8	44911.9	44912.0	44912.1	44912.2	44912.3	44912.4	44912.5	44912.6	44912.7	44912.8	44912.9	44913.0	44913.1	44913.2	44913.3	44913.4	44913.5	44913.6	44913.7	44913.8	44913.9	44914.0	44914.1	44914.2	44914.3	44914.4	44914.5	44914.6	44914.7	44914.8	44914.9	44915.0
---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------	---------

Table 6-1 (continued)

[illegible]

Table 6-1 (continued)

[illegible]

Table 6-1 (continued)

[illegible]

Table 6-1 (continued)

[illegible]

Table 6-1 (continued)

[illegible]

[illegible]

Table 6-1 (continued)

	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54	55	56	57	58	59	60	61	62	63	64	65	66	67	68	69	70	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86	87	88	89	90	91	92	93	94	95	96	97	98	99	100
0	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54	55	56	57	58	59	60	61	62	63	64	65	66	67	68	69	70	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86	87	88	89	90	91	92	93	94	95	96	97	98	99	100
1	100	101	102	103	104	105	106	107	108	109	110	111	112	113	114	115	116	117	118	119	120	121	122	123	124	125	126	127	128	129	130	131	132	133	134	135	136	137	138	139	140	141	142	143	144	145	146	147	148	149	150	151	152	153	154	155	156	157	158	159	160	161	162	163	164	165	166	167	168	169	170	171	172	173	174	175	176	177	178	179	180	181	182	183	184	185	186	187	188	189	190	191	192	193	194	195	196	197	198	199	200
2	200	201	202	203	204	205	206	207	208	209	210	211	212	213	214	215	216	217	218	219	220	221	222	223	224	225	226	227	228	229	230	231	232	233	234	235	236	237	238	239	240	241	242	243	244	245	246	247	248	249	250	251	252	253	254	255	256	257	258	259	260	261	262	263	264	265	266	267	268	269	270	271	272	273	274	275	276	277	278	279	280	281	282	283	284	285	286	287	288	289	290	291	292	293	294	295	296	297	298	299	300
3	300	301	302	303	304	305	306	307	308	309	310	311	312	313	314	315	316	317	318	319	320	321	322	323	324	325	326	327	328	329	330	331	332	333	334	335	336	337	338	339	340	341	342	343	344	345	346	347	348	349	350	351	352	353	354	355	356	357	358	359	360	361	362	363	364	365	366	367	368	369	370	371	372	373	374	375	376	377	378	379	380	381	382	383	384	385	386	387	388	389	390	391	392	393	394	395	396	397	398	399	400
4	400	401	402	403	404	405	406	407	408	409	410	411	412	413	414	415	416	417	418	419	420	421	422	423	424	425	426	427	428	429	430	431	432	433	434	435	436	437	438	439	440	441	442	443	444	445	446	447	448	449	450	451	452	453	454	455	456	457	458	459	460	461	462	463	464	465	466	467	468	469	470	471	472	473	474	475	476	477	478	479	480	481	482	483	484	485	486	487	488	489	490	491	492	493	494	495	496	497	498	499	500
5	500	501	502	503	504	505	506	507	508	509	510	511	512	513	514	515	516	517	518	519	520	521	522	523	524	525	526	527	528	529	530	531	532	533	534	535	536	537	538	539	540	541	542	543	544	545	546	547	548	549	550	551	552	553	554	555	556	557	558	559	560	561	562	563	564	565	566	567	568	569	570	571	572	573	574	575	576	577	578	579	580	581	582	583	584	585	586	587	588	589	590	591	592	593	594	595	596	597	598	599	600
6	600	601	602	603	604	605	606	607	608	609	610	611	612	613	614	615	616	617	618	619	620	621	622	623	624	625	626	627	628	629	630	631	632	633	634	635	636	637	638	639	640	641	642	643	644	645	646	647	648	649	650	651	652	653	654	655	656	657	658	659	660	661	662	663	664	665	666	667	668	669	670	671	672	673	674	675	676	677	678	679	680	681	682	683	684	685	686	687	688	689	690	691	692	693	694	695	696	697	698	699	700
7	700	701	702	703	704	705	706	707	708	709	710	711	712	713	714	715	716	717	718	719	720	721	722	723	724	725	726	727	728	729	730	731	732	733	734	735	736	737	738	739	740	741	742	743	744	745	746	747	748	749	750	751	752	753	754	755	756	757	758	759	760	761	762	763	764	765	766	767	768	769	770	771	772	773	774	775	776	777	778	779	780	781	782	783	784	785	786	787	788	789	790	791	792	793	794	795	796	797	798	799	800
8	800	801	802	803	804	805	806	807	808	809	810	811	812	813	814	815	816	817	818	819	820	821	822	823	824	825	826	827	828	829	830	831	832	833	834	835	836	837	838	839	840	841	842	843	844	845	846	847	848	849	850	851	852	853	854	855	856	857	858	859	860	861	862	863	864	865	866	867	868	869	870	871	872	873	874	875	876	877	878	879	880	881	882	883	884	885	886	887	888	889	890	891	892	893	894	895	896	897	898	899	900
9	900	901	902	903	904	905	906	907	908	909	910	911	912	913	914	915	916	917	918	919	920	921	922	923	924	925	926	927	928	929	930	931	932	933	934	935	936	937	938	939	940	941	942	943	944	945	946	947	948	949	950	951	952	953	954	955	956	957	958	959	960	961	962	963	964	965	966	967	968	969	970	971	972	973	974	975	976	977	978	979	980	981	982	983	984	985	986	987	988	989	990	991	992	993	994	995	996	997	998	999	1000

[illegible]

Table 6-1 (continued)

[illegible]

[illegible]

Table 6-1 (continued)

[illegible]

0	1	2	3	4	5	6	7	8	9	A	B	C	D	E	F	G	H	I	J	K	L	M	N	O	P	Q	R	S	T	U	V	W	X	Y	Z
00000	00001	00002	00003	00004	00005	00006	00007	00008	00009	00010	00011	00012	00013	00014	00015	00016	00017	00018	00019	00020	00021	00022	00023	00024	00025	00026	00027	00028	00029	00030	00031	00032	00033	00034	
00035	00036	00037	00038	00039	00040	00041	00042	00043	00044	00045	00046	00047	00048	00049	00050	00051	00052	00053	00054	00055	00056	00057	00058	00059	00060	00061	00062	00063	00064	00065	00066	00067	00068	00069	
00070	00071	00072	00073	00074	00075	00076	00077	00078	00079	00080	00081	00082	00083	00084	00085	00086	00087	00088	00089	00090	00091	00092	00093	00094	00095	00096	00097	00098	00099	00100	00101	00102	00103	00104	
00105	00106	00107	00108	00109	00110	00111	00112	00113	00114	00115	00116	00117	00118	00119	00120	00121	00122	00123	00124	00125	00126	00127	00128	00129	00130	00131	00132	00133	00134	00135	00136	00137	00138	00139	
00140	00141	00142	00143	00144	00145	00146	00147	00148	00149	00150	00151	00152	00153	00154	00155	00156	00157	00158	00159	00160	00161	00162	00163	00164	00165	00166	00167	00168	00169	00170	00171	00172	00173	00174	
00175	00176	00177	00178	00179	00180	00181	00182	00183	00184	00185	00186	00187	00188	00189	00190	00191	00192	00193	00194	00195	00196	00197	00198	00199	00200	00201	00202	00203	00204	00205	00206	00207	00208	00209	
00210	00211	00212	00213	00214	00215	00216	00217	00218	00219	00220	00221	00222	00223	00224	00225	00226	00227	00228	00229	00230	00231	00232	00233	00234	00235	00236	00237	00238	00239	00240	00241	00242	00243	00244	
00245	00246	00247	00248	00249	00250	00251	00252	00253	00254	00255	00256	00257	00258	00259	00260	00261	00262	00263	00264	00265	00266	00267	00268	00269	00270	00271	00272	00273	00274	00275	00276	00277	00278	00279	
00280	00281	00282	00283	00284	00285	00286	00287	00288	00289	00290	00291	00292	00293	00294	00295	00296	00297	00298	00299	00300	00301	00302	00303	00304	00305	00306	00307	00308	00309	00310	00311	00312	00313	00314	
00315	00316	00317	00318	00319	00320	00321	00322	00323	00324	00325	00326	00327	00328	00329	00330	00331	00332	00333	00334	00335	00336	00337	00338	00339	00340	00341	00342	00343	00344	00345	00346	00347	00348	00349	
00350	00351	00352	00353	00354	00355	00356	00357	00358	00359	00360	00361	00362	00363	00364	00365	00366	00367	00368	00369	00370	00371	00372	00373	00374	00375	00376	00377	00378	00379	00380	00381	00382	00383	00384	

Table 6-1 (continued)

[illegible]

Table 6-1 (continued)

[illegible]

Table 6-1 (continued)

[illegible]

1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54	55	56	57	58	59	60	61	62	63	64	65	66	67	68	69	70	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86	87	88	89	90	91	92	93	94	95	96	97	98	99	100
101	102	103	104	105	106	107	108	109	110	111	112	113	114	115	116	117	118	119	120	121	122	123	124	125	126	127	128	129	130	131	132	133	134	135	136	137	138	139	140	141	142	143	144	145	146	147	148	149	150	151	152	153	154	155	156	157	158	159	160	161	162	163	164	165	166	167	168	169	170	171	172	173	174	175	176	177	178	179	180	181	182	183	184	185	186	187	188	189	190	191	192	193	194	195	196	197	198	199	200
201	202	203	204	205	206	207	208	209	210	211	212	213	214	215	216	217	218	219	220	221	222	223	224	225	226	227	228	229	230	231	232	233	234	235	236	237	238	239	240	241	242	243	244	245	246	247	248	249	250	251	252	253	254	255	256	257	258	259	260	261	262	263	264	265	266	267	268	269	270	271	272	273	274	275	276	277	278	279	280	281	282	283	284	285	286	287	288	289	290	291	292	293	294	295	296	297	298	299	300
301	302	303	304	305	306	307	308	309	310	311	312	313	314	315	316	317	318	319	320	321	322	323	324	325	326	327	328	329	330	331	332	333	334	335	336	337	338	339	340	341	342	343	344	345	346	347	348	349	350	351	352	353	354	355	356	357	358	359	360	361	362	363	364	365	366	367	368	369	370	371	372	373	374	375	376	377	378	379	380	381	382	383	384	385	386	387	388	389	390	391	392	393	394	395	396	397	398	399	400
401	402	403	404	405	406	407	408	409	410	411	412	413	414	415	416	417	418	419	420	421	422	423	424	425	426	427	428	429	430	431	432	433	434	435	436	437	438	439	440	441	442	443	444	445	446	447	448	449	450	451	452	453	454	455	456	457	458	459	460	461	462	463	464	465	466	467	468	469	470	471	472	473	474	475	476	477	478	479	480	481	482	483	484	485	486	487	488	489	490	491	492	493	494	495	496	497	498	499	500
501	502	503	504	505	506	507	508	509	510	511	512	513	514	515	516	517	518	519	520	521	522	523	524	525	526	527	528	529	530	531	532	533	534	535	536	537	538	539	540	541	542	543	544	545	546	547	548	549	550	551	552	553	554	555	556	557	558	559	560	561	562	563	564	565	566	567	568	569	570	571	572	573	574	575	576	577	578	579	580	581	582	583	584	585	586	587	588	589	590	591	592	593	594	595	596	597	598	599	600
601	602	603	604	605	606	607	608	609	610	611	612	613	614	615	616	617	618	619	620	621	622	623	624	625	626	627	628	629	630	631	632	633	634	635	636	637	638	639	640	641	642	643	644	645	646	647	648	649	650	651	652	653	654	655	656	657	658	659	660	661	662	663	664	665	666	667	668	669	670	671	672	673	674	675	676	677	678	679	680	681	682	683	684	685	686	687	688	689	690	691	692	693	694	695	696	697	698	699	700
701	702	703	704	705	706	707	708	709	710	711	712	713	714	715	716	717	718	719	720	721	722	723	724	725	726	727	728	729	730	731	732	733	734	735	736	737	738	739	740	741	742	743	744	745	746	747	748	749	750	751	752	753	754	755	756	757	758	759	760	761	762	763	764	765	766	767	768	769	770	771	772	773	774	775	776	777	778	779	780	781	782	783	784	785	786	787	788	789	790	791	792	793	794	795	796	797	798	799	800
801	802	803	804	805	806	807	808	809	810	811	812	813	814	815	816	817	818	819	820	821	822	823	824	825	826	827	828	829	830	831	832	833	834	835	836	837	838	839	840	841	842	843	844	845	846	847	848	849	850	851	852	853	854	855	856	857	858	859	860	861	862	863	864	865	866	867	868	869	870	871	872	873	874	875	876	877	878	879	880	881	882	883	884	885	886	887	888	889	890	891	892	893	894	895	896	897	898	899	900
901	902	903	904	905	906	907	908	909	910	911	912	913	914	915	916	917	918	919	920	921	922	923	924	925	926	927	928	929	930	931	932	933	934	935	936	937	938	939	940	941	942	943	944	945	946	947	948	949	950	951	952	953	954	955	956	957	958	959	960	961	962	963	964	965	966	967	968	969	970	971	972	973	974	975	976	977	978	979	980	981	982	983	984	985	986	987	988	989	990	991	992	993	994	995	996	997	998	999	1000

Table 6-1 (continued)

Table 6-1 (continued)

Wavelength, μ	1.0	1.5	2.0	3.0	4.0	5.0	6.0	7.0	8.0	9.0	10.0	12.0	15.0	20.0	30.0	40.0	50.0	60.0	70.0	80.0	90.0	100.0	120.0	150.0	200.0	300.0	400.0	500.0	600.0	700.0	800.0	900.0	1000.0
Refractive index, n	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000
Extinction coefficient, k	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000
Attenuation coefficient, α	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000
Phase shift, δ	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000
Reflection coefficient, R	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000	0.000000
Transmission coefficient, T	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000	1.000000

T = 0 FOR WAVELENGTHS λ > 100 μ AND λ IN THIS FREQUENCY REGION

Table 6-1 (continued)

[illegible]

Table 6-1 (continued)

[illegible]

[illegible]

Table 6-1 (continued)

	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54	55	56	57	58	59	60	61	62	63	64	65	66	67	68	69	70	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86	87	88	89	90	91	92	93	94	95	96	97	98	99	100
1	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54	55	56	57	58	59	60	61	62	63	64	65	66	67	68	69	70	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86	87	88	89	90	91	92	93	94	95	96	97	98	99	100

Table 6-1 (continued)

[illegible]

Table 6-1 (continued)

[illegible]

1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54	55	56	57	58	59	60	61	62	63	64	65	66	67	68	69	70	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86	87	88	89	90	91	92	93	94	95	96	97	98	99	100
101	102	103	104	105	106	107	108	109	110	111	112	113	114	115	116	117	118	119	120	121	122	123	124	125	126	127	128	129	130	131	132	133	134	135	136	137	138	139	140	141	142	143	144	145	146	147	148	149	150	151	152	153	154	155	156	157	158	159	160	161	162	163	164	165	166	167	168	169	170	171	172	173	174	175	176	177	178	179	180	181	182	183	184	185	186	187	188	189	190	191	192	193	194	195	196	197	198	199	200
201	202	203	204	205	206	207	208	209	210	211	212	213	214	215	216	217	218	219	220	221	222	223	224	225	226	227	228	229	230	231	232	233	234	235	236	237	238	239	240	241	242	243	244	245	246	247	248	249	250	251	252	253	254	255	256	257	258	259	260	261	262	263	264	265	266	267	268	269	270	271	272	273	274	275	276	277	278	279	280	281	282	283	284	285	286	287	288	289	290	291	292	293	294	295	296	297	298	299	300
301	302	303	304	305	306	307	308	309	310	311	312	313	314	315	316	317	318	319	320	321	322	323	324	325	326	327	328	329	330	331	332	333	334	335	336	337	338	339	340	341	342	343	344	345	346	347	348	349	350	351	352	353	354	355	356	357	358	359	360	361	362	363	364	365	366	367	368	369	370	371	372	373	374	375	376	377	378	379	380	381	382	383	384	385	386	387	388	389	390	391	392	393	394	395	396	397	398	399	400
401	402	403	404	405	406	407	408	409	410	411	412	413	414	415	416	417	418	419	420	421	422	423	424	425	426	427	428	429	430	431	432	433	434	435	436	437	438	439	440	441	442	443	444	445	446	447	448	449	450	451	452	453	454	455	456	457	458	459	460	461	462	463	464	465	466	467	468	469	470	471	472	473	474	475	476	477	478	479	480	481	482	483	484	485	486	487	488	489	490	491	492	493	494	495	496	497	498	499	500
501	502	503	504	505	506	507	508	509	510	511	512	513	514	515	516	517	518	519	520	521	522	52																																																																													

Table 6-1 (continued)

[illegible]

[illegible]

Table 6-1 (continued)

Base No.	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54	55	56	57	58	59	60	61	62	63	64	65	66	67	68	69	70	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86	87	88	89	90	91	92	93	94	95	96	97	98	99	100
Base No.	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54	55	56	57	58	59	60	61	62	63	64	65	66	67	68	69	70	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86	87	88	89	90	91	92	93	94	95	96	97	98	99	100
Base No.	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54	55	56	57	58	59	60	61	62	63	64	65	66	67	68	69	70	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86	87	88	89	90	91	92	93	94	95	96	97	98	99	100
Base No.	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54	55	56	57	58	59	60	61	62	63	64	65	66	67	68	69	70	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86	87	88	89	90	91	92	93	94	95	96	97	98	99	100
Base No.	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	24	25	26	27	28	29	30	31	32	33	34	35	36	37	38	39	40	41	42	43	44	45	46	47	48	49	50	51	52	53	54	55	56	57	58	59	60	61	62	63	64	65	66	67	68	69	70	71	72	73	74	75	76	77	78	79	80	81	82	83	84	85	86	87	88	89	90	91	92	93	94	95	96	97	98	99	100

Table 6-1 (continued)

[illegible]

SECTION 7

TABLE OF INTEGRATED ABSORPTANCE

Values of the integrated absorptance are presented in Table 7-1 for the samples which are listed in Table 4-1 and whose spectra and tables of transmittance are included in Sections 5 and 6, respectively. Except for wavenumbers near the ends of the spectral region of interest, all the tabulations for a given spectral interval are contained on two successive pages. A single page was sufficient to contain all the tabulations for the samples considered near the ends of the spectral region covered by the table.

The lower limit of integration ν^1 , which is shown at the top of each column, was chosen at a point where there was essentially no absorption. The integrated absorptance between any two wavenumbers listed can be found by subtracting the values tabulated at the two points. The accuracy of a given value of the integrated absorptance does not justify all of the significant figures which are carried in the table. But all these figures must be retained in order to avoid losing information about the integrated absorptance over narrow intervals which is obtained by subtracting the tabulated values. The table was formed by photographing part of the computer output with the proper headings pieced in above each section.

The integrated absorptance at a given ν was calculated by the computer from the values of transmittance given in Table 6-1 by the following expression.

$$\int_{\nu^1}^{\nu} [1 - T(\nu)] d\nu = \left[\sum_{\nu=\nu^1}^{\nu} [1 - T(\nu)] \Delta\nu \right] - \frac{[1 - T(\nu^1)]}{2} + \frac{[1 - T(\nu)]}{2} \Delta\nu,$$

where $\Delta\nu = 0.1 \text{ cm}^{-1}$, the wavenumber interval between successive values of $T(\nu)$ in Table 6-1.

The important sample parameters are listed at the heading of each page.

In regions where the structure in the spectra is regular, the integrated absorptance was calculated at wavenumbers midway between the line centers. Therefore, in the case of no overlapping, the difference between two successive tabulated values is the equivalent width of the absorption line in the interval. In some other regions where the structure was not particularly regular, the integrated absorptance was calculated at wavenumbers corresponding to absorptance minima. In still other regions where there was but little structure or where the absorptance minima would shift as the pressure or absorber thickness was changed, the values were calculated at integral wavenumbers or possibly at integral half-wavenumbers, depending on the amount of information one might expect to obtain from closer readings.

Table 7-1
 $\int (1-T) dz$ [illegible]

Table 7-1 $\left[\int_0^1 (1-T) dz \right]$ (continued)

Table 7-1 $\int_0^{\infty} (1-T) d\nu$ (continued)									
ν (cm ⁻¹)	ν (cm ⁻¹)	ν (cm ⁻¹)	ν (cm ⁻¹)	ν (cm ⁻¹)	ν (cm ⁻¹)	ν (cm ⁻¹)	ν (cm ⁻¹)	ν (cm ⁻¹)	ν (cm ⁻¹)
4648.4 4648.5 4648.6 4648.7 4648.8 4648.9 4649.0 4649.1 4649.2 4649.3 4649.4 4649.5 4649.6 4649.7 4649.8 4649.9 4650.0 4650.1 4650.2 4650.3 4650.4 4650.5 4650.6 4650.7 4650.8 4650.9 4651.0 4651.1 4651.2 4651.3 4651.4 4651.5 4651.6 4651.7 4651.8 4651.9 4652.0 4652.1 4652.2 4652.3 4652.4 4652.5 4652.6 4652.7 4652.8 4652.9 4653.0 4653.1 4653.2 4653.3 4653.4 4653.5 4653.6 4653.7 4653.8 4653.9 4654.0 4654.1 4654.2 4654.3 4654.4 4654.5 4654.6 4654.7 4654.8 4654.9 4655.0 4655.1 4655.2 4655.3 4655.4 4655.5 4655.6 4655.7 4655.8 4655.9 4656.0 4656.1 4656.2 4656.3 4656.4 4656.5 4656.6 4656.7 4656.8 4656.9 4657.0 4657.1 4657.2 4657.3 4657.4 4657.5 4657.6 4657.7 4657.8 4657.9 4658.0 4658.1 4658.2 4658.3 4658.4 4658.5 4658.6 4658.7 4658.8 4658.9 4659.0 4659.1 4659.2 4659.3 4659.4 4659.5 4659.6 4659.7 4659.8 4659.9 4660.0 4660.1 4660.2 4660.3 4660.4 4660.5 4660.6 4660.7 4660.8 4660.9 4661.0 4661.1 4661.2 4661.3 4661.4 4661.5 4661.6 4661.7 4661.8 4661.9 4662.0 4662.1 4662.2 4662.3 4662.4 4662.5 4662.6 4662.7 4662.8 4662.9 4663.0 4663.1 4663.2 4663.3 4663.4 4663.5 4663.6 4663.7 4663.8 4663.9 4664.0 4664.1 4664.2 4664.3 4664.4 4664.5 4664.6 4664.7 4664.8 4664.9 4665.0 4665.1 4665.2 4665.3 4665.4 4665.5 4665.6 4665.7 4665.8 4665.9 4666.0 4666.1 4666.2 4666.3 4666.4 4666.5 4666.6 4666.7 4666.8 4666.9 4667.0 4667.1 4667.2 4667.3 4667.4 4667.5 4667.6 4667.7 4667.8 4667.9 4668.0 4668.1 4668.2 4668.3 4668.4 4668.5 4668.6 4668.7 4668.8 4668.9 4669.0 4669.1 4669.2 4669.3 4669.4 4669.5 4669.6 4669.7 4669.8 4669.9 4670.0 4670.1 4670.2 4670.3 4670.4 4670.5 4670.6 4670.7 4670.8 4670.9 4671.0 4671.1 4671.2 4671.3 4671.4 4671.5 4671.6 4671.7 4671.8 4671.9 4672.0 4672.1 4672.2 4672.3 4672.4 4672.5 4672.6 4672.7 4672.8 4672.9 4673.0 4673.1 4673.2 4673.3 4673.4 4673.5 4673.6 4673.7 4673.8 4673.9 4674.0 4674.1 4674.2 4674.3 4674.4 4674.5 4674.6 4674.7 4674.8 4674.9 4675.0 4675.1 4675.2 4675.3 4675.4 4675.5 4675.6 4675.7 4675.8 4675.9 4676.0 4676.1 4676.2 4676.3 4676.4 4676.5 4676.6 4676.7 4676.8 4676.9 4677.0 4677.1 4677.2 4677.3 4677.4 4677.5 4677.6 4677.7 4677.8 4677.9 4678.0 4678.1 4678.2 4678.3 4678.4 4678.5 4678.6 4678.7 4678.8 4678.9 4679.0 4679.1 4679.2 4679.3 4679.4 4679.5 4679.6 4679.7 4679.8 4679.9 4680.0 4680.1 4680.2 4680.3 4680.4 4680.5 4680.6 4680.7 4680.8 4680.9 4681.0 4681.1 4681.2 4681.3 4681.4 4681.5 4681.6 4681.7 4681.8 4681.9 4682.0 4682.1 4682.2 4682.3 4682.4 4682.5 4682.6 4682.7 4682.8 4682.9 4683.0 4683.1 4683.2 4683.3 4683.4 4683.5 4683.6 4683.7 4683.8 4683.9 4684.0 4684.1 4684.2 4684.3 4684.4 4684.5 4684.6 4684.7 4684.8 4684.9 4685.0 4685.1 4685.2 4685.3 4685.4 4685.5 4685.6 4685.7 4685.8 4685.9 4686.0 4686.1 4686.2 4686.3 4686.4 4686.5 4686.6 4686.7 4686.8 4686.9 4687.0 4687.1 4687.2 4687.3 4687.4 4687.5 4687.6 4687.7 4687.8 4687.9 4688.0 4688.1 4688.2 4688.3 4688.4 4688.5 4688.6 4688.7 4688.8 4688.9 4689.0 4689.1 4689.2 4689.3 4689.4 4689.5 4689.6 4689.7 4689.8 4689.9 4690.0 4690.1 4690.2 4690.3 4690.4 4690.5 4690.6 4690.7 4690.8 4690.9 4691.0 4691.1 4691.2 4691.3 4691.4 4691.5 4691.6 4691.7 4691.8 4691.9 4692.0 4692.1 4692.2 4692.3 4692.4 4692.5 4692.6 4692.7 4692.8 4692.9 4693.0 4693.1 4693.2 4693.3 4693.4 4693.5 4693.6 4693.7 4693.8 4693.9 4694.0 4694.1 4694.2 4694.3 4694.4 4694.5 4694.6 4694.7 4									

Table 7-1 $\left[\int_0^1 (1-T) dz \right]$ (continued)

[illegible]

Table 7-1 $\left[\int_0^1 (1-T) dz \right]$ (continued)[illegible]

Table 7-1 $\left[\int_0^{\infty} (1-T) dz \right]$ (continued)

[illegible]

Table 7-1 $\left[\int_0^1 (1-T) d\nu \right]$ (continued)

[illegible]

Table 7-1
 $\int_0^1 (1-T) dz$ (continued)

[illegible]

Table 7-1 $[\int_0^1 (1-T) dz]$ (continued)

[illegible]

Table 7-1 $\int (1-T) dz$ (continued)[illegible]

Table 7-1 $\left[\int_0^1 (1-T) dz \right]$ (continued)

[illegible]

Table 7-1 $\left[\int_{\mu}^{\infty} (I-T) dz \right]$ (continued)

[illegible]

(continued)

[illegible]

SECTION 8

REFERENCES

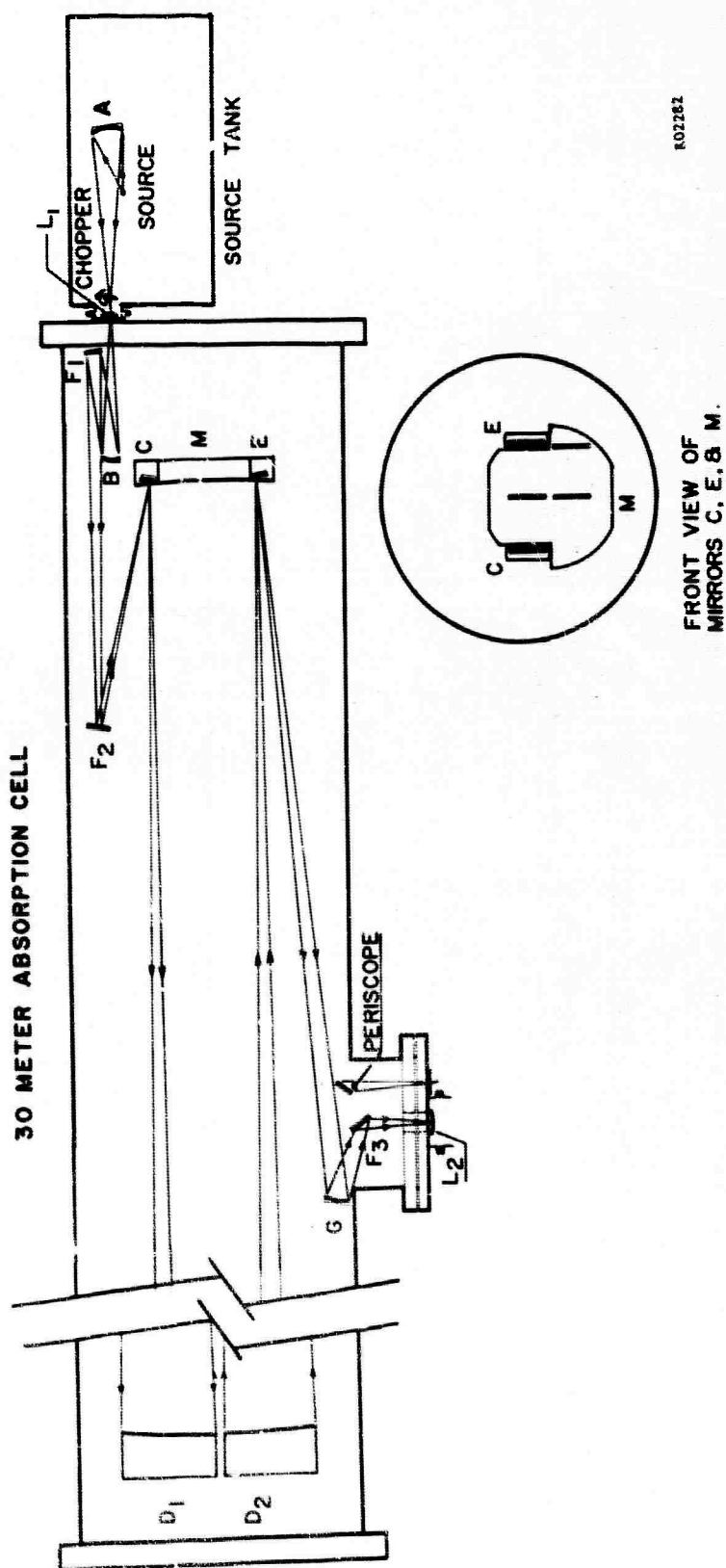
1. J. N. Howard, D. E. Burch and Dudley Williams, J. Opt. Soc. Am. 46, 237 (1956).
2. C. P. Courtoy, Annales de la Societe Scientifique de Bruxelles, Serie 1, pp 5-230 (27 March 1959). Also, C. P. Courtoy, Canad. J. Phys. 35, 608 (1957).
3. D. Weber, R. J. Holm and S. S. Penner, J. Chem. Phys. 20, 1820 (1952).
4. J. C. Breeze and C. C. Ferriso, J. Chem. Phys. 40, No. 5, 1276 (1964).
5. G. P. Kuiper, Communications of Lunar and Planetary Laboratory, Vol. 1, No. 14-16, Univ. of Arizona Press (1962).
6. D. E. Burch, D. A. Gryvnak and Dudley Williams, Appl. Opt. 1, 759 (1962).
7. B. H. Winters, S. Silverman and W. S. Benedict, J. Quan. Spec. 4, 527 (1964).
8. F. A. Diaz, J. Opt. Soc. Am 53, 203 (1963).
9. Dr. William S. Benedict, Johns Hopkins University (Private Communication).

APPENDIX A

30-METER MULTIPLE-PASS CELL AND SOURCE TANK

Figure A-1 shows an optical diagram of a multiple-pass absorption cell with a base length of approximately 30 meters and an adjoining vacuum tank which contains a radiation source, chopper and mirrors. Light from a source is focused by mirror A on field lens L_1 which serves as a window. From there the light passes on to mirror B which forms an enlarged image of the source on mirror C. The light then passes through the multiple-pass optical system which is of the type described by White (J. Opt. Soc. Am. 32, 285 (1942)). After traversing the multiple-pass system a certain number of times, which depends on the settings of mirrors D_1 and D_2 , an image of the source is formed on mirror E. The light then is directed to mirror G which re-images the source on field lens L_2 , which also serves as the exit window. From here the radiation is focused on the entrance slit of one of two spectrometers which are available in the laboratory. F_1 , F_2 and F_3 represent flat mirrors, and all mirrors are front surface aluminized.

The principle of the White multiple-pass system can be explained with the aid of the lower portion of Figure 1, which shows a front view of mirrors C, E and M, with the positions of the images shown on M when the system is adjusted for 8 passes. Light from the image on C travels to D_1 whose radius of curvature is equal to the distance between M and D_1 ; from there the light is reflected to form the lower right-hand image on M. M is a spherical mirror whose radius of curvature is, like D_1 and D_2 , equal to the distance (29.0 meters) between M and D_1 or D_2 . The center of curvature of M is mid-way between D_1 and D_2 ; therefore, all the light from the lower right-hand image falls on D_2 , which re-focuses the beam to the image in the upper center portion of M. The light then goes to D_1 and to the lower left image on M, on to D_2 and back to E, from where it is directed to G and out of the cell. The optical path from L_1 to C and from E to L_2 is 5 meters; therefore the total path inside the cell is $5 + 29 N$ meters, where N is the number of passes.



RO2282

RO2282 FIGURE A-1. OPTICAL DIAGRAM OF THE 30-METER MULTIPLE-PASS ABSORPTION CELL AND SOURCE TANK

The lower portion shows a front view of mirrors C, E and M with the image pattern corresponding to 8 passes; the images are approximately 15 by 1.2 cm. The distance from M to D₁ and D₂ is 29 meters.

The number of passes of the beam before it strikes E is changed by rotating D₂ about a vertical axis. In many cells of this type, the number of passes is varied by moving both D₁ and D₂; however, it is only necessary to move one. D₁ and D₂ are D-shaped, both having been cut from a single mirror of 81 cm diameter, rather than from two circular mirrors, in order to obtain the maximum possible light gathering power with the 91 cm diameter tube in which the optics are installed. M is cut from a 46 cm diameter mirror.

Since the f-ratio of the multiple-pass system is large (approximately 75), it is necessary to use large images on C, E and M and to reduce them to the height of the entrance slit of the spectrometer in order to fill it properly. The images on C, E and M are approximately 15 cm high x 1.2 cm wide and are reduced to about 3.0 cm high at lens L₂. The image is further reduced by optics outside the cell to match the height of the entrance slit of the spectrometer being used. The entrance optics work in just the opposite manner; an enlarged image of the source is formed at L₁, and this image is further enlarged and formed on C. Mirrors B and G are similar, as are C and E.

Vignetting is avoided by using lenses L₁ and L₂ instead of flat windows and by using spherical mirrors C and E, rather than flats as is frequently done. Lens L₁ images mirror A on B; and C images B on D₁. Similarly, E images D₂ on G, which is imaged on another mirror that is outside the cell, but not shown in Figure A-1. The need for curved mirrors C and E, rather than flats, and for lenses L₁ and L₂ is not apparent from Figure 1, which shows only the rays in a horizontal plane. However, the need is clear when one considers the height of the images on C, E and M. For example, if C were replaced by a flat mirror, the necessary height of the beam at B in order for light to go from every point on the 15 cm high image to every usable point on D₁ is approximately 20 cm. The maximum height of the portion of D₁ which is used is about 60 cm, and the optical distance from B to C is approximately 210 cm, thus the height of the beam on B is only about

$$60 \times \frac{210}{2900} \approx 4.4 \text{ cm}$$

when C is used with B and D₁ at conjugate points. Similarly, if L₁ were a plane window, the necessary height of the beam at A would be much greater. A similar argument applies for E and lens L₂.

The loss that would arise from a slight error in the focal lengths of L_1 and L_2 is small; therefore, loss due to chromatic aberrations in the two lenses is negligible. The lenses, which are made from KBr, were ground to have the proper focal length for an intermediate wavelength of 3μ , so that the system can be used satisfactorily from the visible to 25μ . Since the f-ratio of the beam inside the cell is large and the reflections at B, C, E and G are nearly normal, these mirrors can be spherical without introducing a significant amount of aberration.

The front of M can be viewed through a small window and a flat mirror which is labelled as a periscope in Figure A-1. In this manner, it is possible to see the images on M and to determine the number of passes to which the system has been adjusted. Vertical and azimuthal adjustments can be made on both D_1 and D_2 from outside the cell by means of rotary seals. The adjustments on D_1 are used to position the first image on M; and the vertical adjustment on D_2 is used to place the centers of curvature of D_1 and D_2 at the same height. The azimuthal adjustment on D_2 is much finer and is used to change the number of passes.

The long cell is provided with a two-stage mechanical pump which is capable of attaining a pressure of a few tenths of a micron of Hg. The maximum pressure which can be used safely is about 2.5 atmospheres. Pressures are measured by a McLeod gauge, an oil manometer or an Hg manometer, depending on the pressure range. Electrical heating wire and insulation have been installed so that it is possible to heat the cell to approximately 70°C . Two fans inside the cell provide mixing when samples consisting of two or more gases are used.

It should be noted that the absorption cell was designed around the long pipe which was installed in the basement of the building as it was being erected before the authors were employed by Aeronutronic. For this reason, some of the things about the cell are not as they would be if the tube had not been previously installed. The cell could be improved somewhat if it were made of glass-lined steel or of stainless steel which could be honed to provide a very smooth inner surface. With either of these surfaces the adsorption of gases, particularly H_2O , on the walls would be less troublesome than with the rough steel walls which were used. It would also be more desirable if the cell were somewhat smaller and designed to hold higher pressures. If it were smaller, it would also be more practical to cool it to temperatures corresponding to the upper atmosphere.

The source tank is constructed of steel pipe having a 45 cm diameter and 6 mm wall thickness. A steel plate which has been welded in position approximately 8 cm below the horizontal diameter provides

a surface for mounting a chopper, a radiation source and mirrors. The tank can be connected to the long cell by the use of flexible bellows and can be evacuated to avoid errors due to absorption by atmospheric gases. The interior of the tank can be viewed through a window in the end plate farthest from the absorption cell. Small adjustments on mirror A, or on any other mirror which might be used, can be made from the exterior by the use of a rotating vacuum seal.

APPENDIX B

ONE-METER MULTIPLE-PASS ABSORPTION CELL

A shorter multiple-pass absorption cell has been constructed to supplement the 30-meter one in our laboratory. The shorter one, like the long one, has multiple-pass optics using the principle described by White (J. Opt. Soc. Am. 32, 285 (1942)), with two mirrors at one end and a single mirror at the other end where the radiation enters and exits. The distance between the mirrors is 102.7 cm; and with an additional 4.8 cm at the side of the cell, the total path for N passes is

$$L(\text{cm}) = 4.8 + N (102.7).$$

Each of the two mirrors at the end opposite the windows has a useful area approximately 6 cm x 6 cm. Plano-convex lenses serve as windows to reduce vignetting. The height of the image at the single mirror is limited by the windows to approximately 4.5 cm. By means of a relatively simple set of mirrors, the image size can be reduced and the divergence of the beam increased in order to properly fill either of our monochromators, a Perkin-Elmer Model 99 or a "home-made" one utilizing a grating. The mirrors in the cell are coated with gold, rather than aluminum, because of the higher reflectivity of gold between 0.7 and 2 microns. In order to change the number of passes, the two mirrors at one end are rotated about vertical axes. This is accomplished by the use of a screw having left-hand threads on one end and right-hand threads on the other; the nuts are spring loaded against arms which extend approximately 70 cm in front of each mirror. The screw is coupled to an external crank through a rotating seal so that the number of passes can be changed without opening the cell or changing the sample.

A flat mirror mounted just below a small window in the top of the cell makes it possible to view the images on the single mirror near the windows. The number of passes can be determined from the image pattern.

The end of the cell containing the windows is fitted to a flexible bellows which is joined to a vacuum tank containing the remainder of the optics. Therefore, it is possible to essentially eliminate absorption by atmospheric gases outside of the cell.

The cell is constructed from a piece of stainless steel pipe 120 cm long with an I.D. of 15 cm. The inside of the pipe is honed to provide a smooth surface and reduce adsorption of gases on the wall. All of the mirror mounts are also made of stainless steel so that many gases which would corrode other metals can be studied.

The entire cell is wrapped with heating coils and insulation so that it can be heated to approximately 170°C. A separate heating circuit is supplied for each end plate and one for the main body of the cell; still another circuit supplies extra heat in the regions of the valves and thermocouple wells which would otherwise be cooler than the remainder of the cell. The resistance wire with asbestos insulation was threaded through copper tubing which was then wrapped around the cell and soft-soldered to it to provide a good thermal contact. Three thermocouples at different locations inside the cell make it possible to measure the temperature so that the power to each heater can be adjusted for a uniform temperature.

The cell can also be cooled by putting it in an insulated tub which was designed for the cell and can be filled with dry-ice, ice water or any other desired cooling material so that the cell is completely immersed in it. In most of the measurements which are anticipated, it will not be necessary to attain an exact temperature; therefore, a continuous temperature control is not essential. Dry-ice or ice water as refrigerants have the advantage that the temperature is self-controlled. Some temperatures between those of these two materials can probably be attained by using frozen mixtures of water and alcohol. At the present time, the cell has not been used at reduced temperatures.

APPENDIX C

DATA DIGITIZING SYSTEM

I INTRODUCTION

A system for analysis and presentation of spectrographic data has been developed. This system performs the following:

- (1) The output of the spectrometer consisting of pen deflection as a function of prism or grating positions can be presented directly in digital form as pairs of numbers displayed on lamp bank indicators and punched and printed on standard IBM cards.
- (2) One can take the original strip charts, make any corrections to account for the presence of any unwanted gas or impurities and fit in the appropriate background curve (the curve giving the pen deflection with the sample removed). These corrected data can be both replotted and digitized as percent deflection vs. grating or prism position. The digital values again are displayed and punched on standard IBM cards. This freedom to correct, modify and study a number of strip charts (each perhaps representative of the same phenomenon with some systematic change in some parameter of the experiment) before the material is digitized is extremely convenient and desirable.

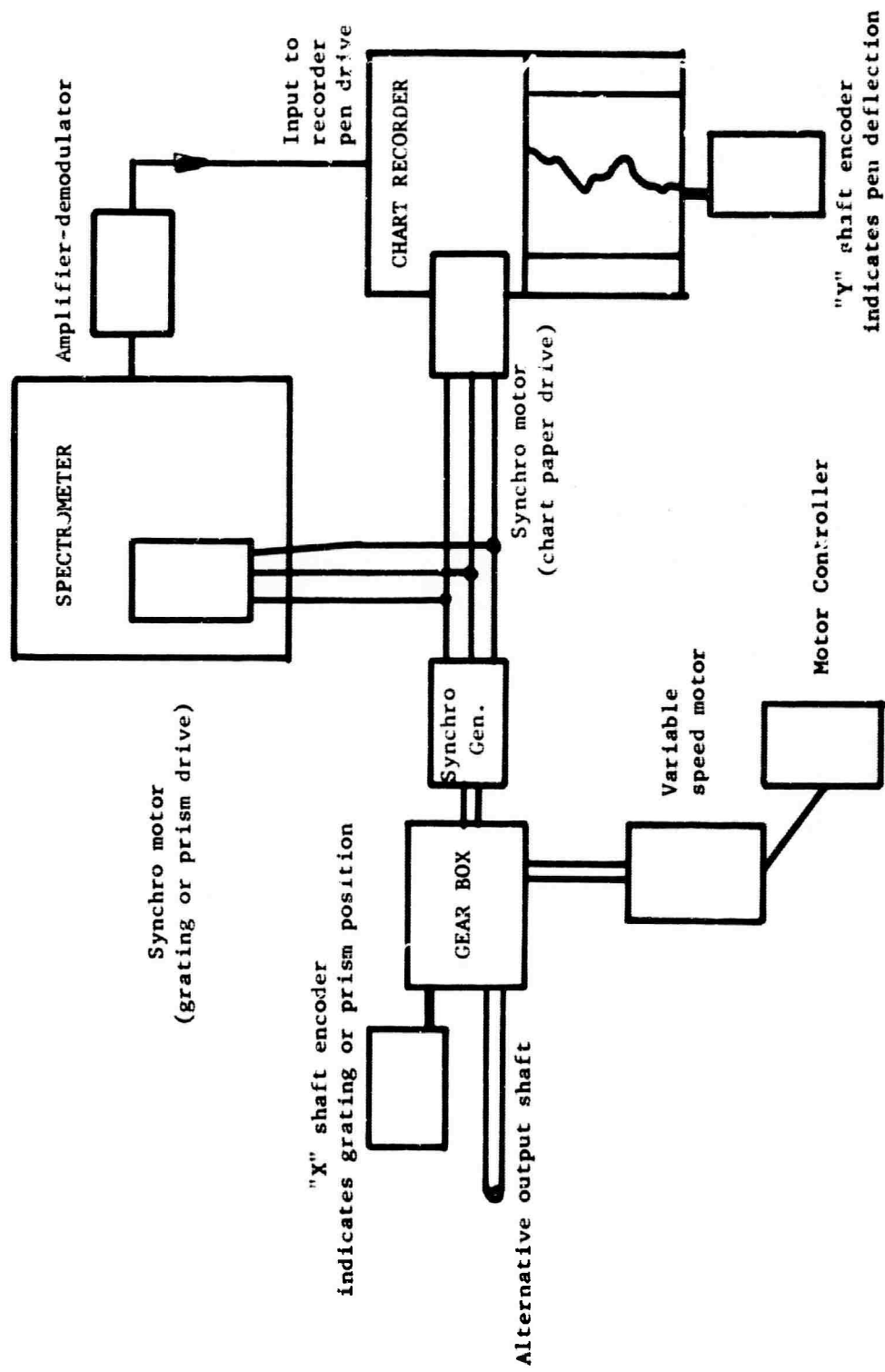
The pairs of data points can be taken at a rate slightly faster than one per second.

We will describe the system and its two basic modes of application in the following section.

II DESCRIPTION

a. The Direct Digitizing System

Figure C-1 shows the direct digitizing system. The spectrometer's prism or grating position, the chart advance of the strip chart recorder, and the X Shaft Encoder are driven directly or through "synchros" by a variable speed motor enabling one to match their speed to the nature of the spectrum. The slowest speed and the maximum density of recorded data points are selected to accurately present the spectrum in those regions where it is most complicated. By increasing the speed and/or reducing the density of recorded data points one can significantly reduce the total time required to cover the spectrum without loss of resolution. The gear ratios on the gear box are selected to establish the density of points derived from the



RO9911 FIGURE C-1. DIRECT DIGITIZING SYSTEM

X Shaft Encoder indicating the position of the grating or prism; the number of these points which are actually recorded is controlled by the Readout Control device described below. The electrical output of the spectrometer is amplified, demodulated, and applied to the chart recorder pen drive. The resulting pen deflection is measured by the Y Shaft Encoder. The outputs of the shaft encoders go to the data processor subsystem.

b. The Digitizer and Data Processor Subsystem

This subsystem was designed and manufactured by Coleman Electronic Systems, Santa Ana, California, to our specifications. It consists of the following:

(1) Shaft Encoders

Two shaft encoders are used. They indicate the angular position of their input shaft in terms of contact positions of ganged rotary switches. The one mounted on a chart recorder measures pen deflection in units 0000 to 1,000. The second encoder synchronized with the chart paper drive measures the distance along the chart paper in units 00000 to 99,999.

(2) Data Processor

On command for readout the data processor unit translates the contact positions of the shaft encoder into contact closures of several banks of relays. These contact closures are presented in two forms: as a visual, numerical display on a lamp bank; and as an output to drive an IBM summary punch.

(3) Readout Control

The density of recorded data points can be adjusted so that a readout cycle is initiated with each change of 1, 2, 5 or 10 units of the X Shaft Encoder. This enables one to increase the number of points in those regions where the spectrum is complicated.

(4) Summary Punch Control

This device samples the data stored in the data processor during the readout cycle in a sequence and form compatible with an IBM 526 Summary Punch which records the data on standard IBM cards. In addition,

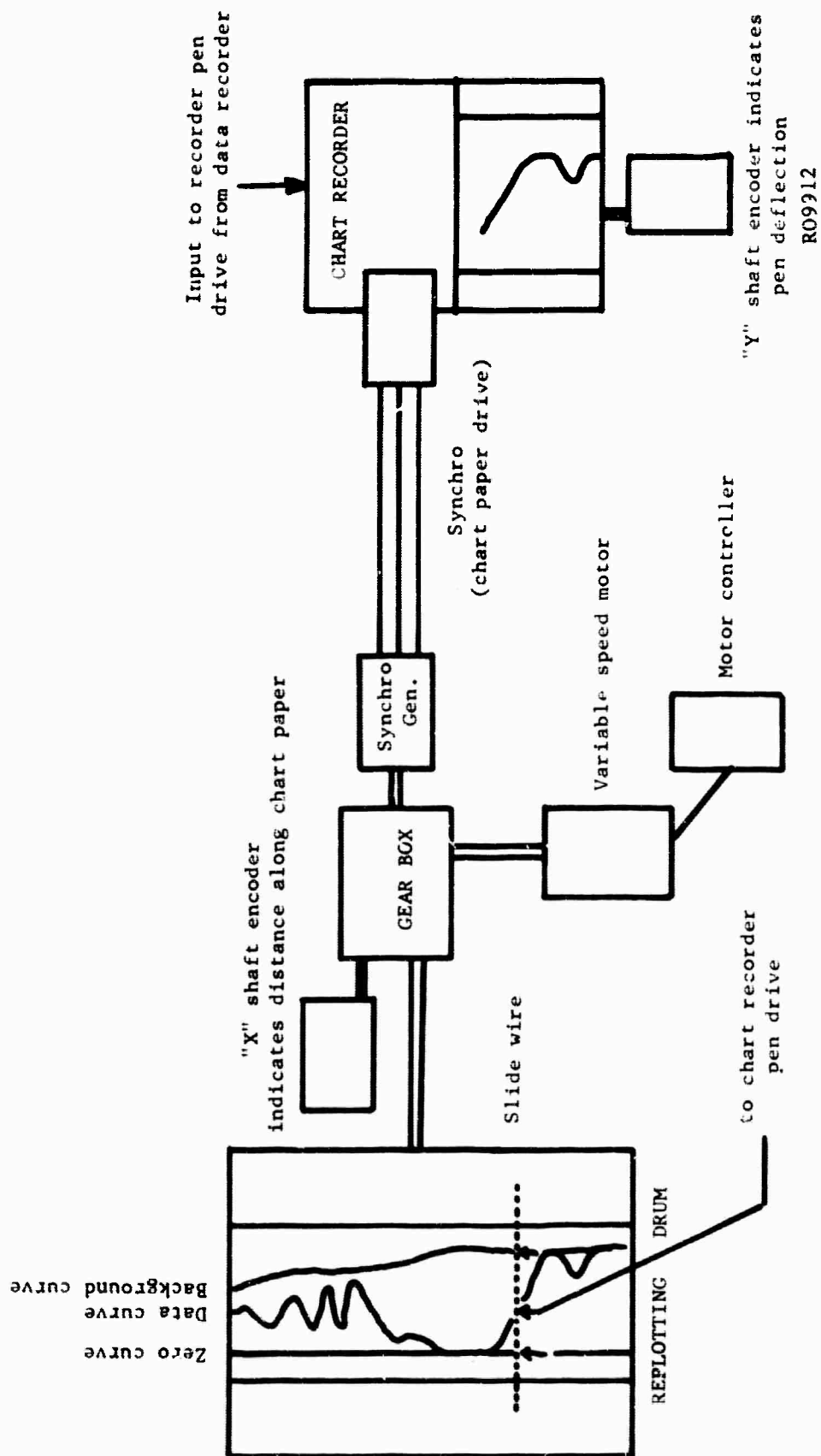
it provides for interlocking controls so that each punching sequence is completed before the following readout cycle can be initiated. The 80 columns of the IBM card can be used as convenient for recording the pairs of data points and any identifying code symbols. We have chosen to record seven pairs of data points and a five digit number to identify the spectrum on each card.

c. The Replotting and Digitizing System

Figure C-2 shows a modification of the direct digitizing system whereby the data previously recorded on strip charts can be digitized. In addition, if a zero level curve and a background curve are also recorded on the strip chart, the percent that the data curve is of the background curve can be digitized and replotted. The background curve corresponds to full scale; for example, in a simple transmission spectrum measurement of some sample it shows the spectrometer's output with the sample removed. In this application the alternative output shaft of the gear box drives a drum onto which is placed the strip chart which is to be digitized. This drum thus rotates in synchronism with the chart paper drive of the chart recorder and with the X Shaft Encoder. On a carriage mounted over the drum are three pointers which are tied to slide contacts along a slide wire--one for the background curve, one for the data curve and one for the zero level curve. A constant voltage corresponding to full scale deflection on the chart recorder is applied between the zero level slide contact and the background curve contact. The voltage from the data curve and the zero level contact thus give a measure of the percent deflection. Except to compensate for small drifts in the spectrometer characteristics, the zero point does not need to be adjusted. The data curve pointer is manually adjusted to follow the data curve; and through the use of two photodetectors in a bridge circuit and a servo system, the background pointer automatically follows the background curve.

III SUMMARY

We have described a very useful tool for the presentation and analysis of spectrographic data. The data can be directly digitized as it is being recorded or can be taken from strip chart records after corrections or modifications have been made. In addition, the percentage that a deflection shown on a strip chart is of some background curve can be recorded digitally and replotted. The output in the form of IBM punch cards enables the data to be conveniently introduced into a computer for additional analysis.



R09912 FIGURE C-2. REPLOTTING AND DIGITIZING SYSTEM

PHILCO CORPORATION

A SUBSIDIARY OF *Ford Motor Company*

AERONUTRONIC DIVISION

FORD ROAD/NEWPORT BEACH, CALIFORNIA

January 28, 1965

AD 609 921
Please find enclosed the corrected copy of page 3-5 for Report U-2955 entitled "Scientific Report - Absorption By CO₂ Between 4500 and 5400 cm⁻¹" dated 15 December 1964.

The strengths given for the bands under the sub-heading New Bands were all in error by a factor of 10⁻⁴ in the original report. Please make note in your copy or insert the attached corrected page.

Darrell E. Burch
David A. Gryvnak
Richard R. Patty

ARCHIVE COPY

Emt

TABLE 3-2
BAND STRENGTHS

Transition ^a ν_0 cm ⁻¹	$\int K(\nu) d\nu$ (atm ⁻¹ cm ⁻¹ cm ⁻¹) STP					Total Band ^b	Total Band Weber et al ^c
	P	Q	R				
<u>Major Bands</u>							
04 ⁰ ₁ 4853.578	0.11	0.00	0.12			0.23	0.30
12 ⁰ ₁ 4977.793	0.52	0.00	0.53			1.05	1.11
20 ⁰ ₁ 5099.619	0.15	0.00	0.18			0.33	0.47
<u>New Bands^d</u>							
15 ¹ ₀ 4591.0	0.16x10 ⁻⁴	0.04x10 ⁻⁴	0.22x10 ⁻⁴			0.42x10 ⁻⁴	
06 ⁰ ₁ -10 ⁰ ₀ 4587.75	0.06x10 ⁻⁴						
30 ⁰ ₁ -02 ² ₀ 5217.63		0.00x10 ⁻⁴	0.27x10 ⁻⁴				
02 ⁰ ₂ -01 ¹ ₀ 5248	0.09x10 ⁻⁴	0.17x10 ⁻⁴	0.07x10 ⁻⁴			0.33x10 ⁻⁴	
10 ⁰ ₂ -01 ¹ ₀ 5349	0.03x10 ⁻⁴	0.08x10 ⁻⁴	0.03x10 ⁻⁴			0.14x10 ⁻⁴	

^aAll bands listed in this Table pertain to the C¹²O₂¹⁶ molecule.

^bThe uncertainty in the values for the major bands is about six or eight percent, but it is considerably higher for the new bands.

^cValues are taken from Weber, Holm and Penner³ and multiplied by 300/273 to adjust to density at STP. The authors give an uncertainty of \pm 20%.

^dThe strengths of the new bands were determined from our raw spectra by Benedict⁹.